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Abstract

During the Quaternary several glaciations occurred in the mountain regions around the Mediterranean and, in recent years, new ages have better constrained their timing. However, this is not the case for the Apuan Alps, a high-rainfall mountain chain adjacent to the Mediterranean Sea. Here, in spite of the widespread evidence for glaciers, the complete lack of geochronological information hinders our understanding of glaciation history. In this paper, we utilize speleothem ages to better constrain the timing of these glacial features. We re-examine 293 uranium-thorium ages from 19 speleothems collected in five caves at different elevations. After a period of very low growth between 160 and 132 ka, the analysed speleothems grew almost continuously to ~75 ka, this period was followed by intermittent growth with lower deposition rate and presence of hiatuses until ~12.5 /12 ka. This is consistent with an ice coverage persisting over the Apuan Alps, inhibiting or interrupting the growth of speleothems via the limited availability of groundwater and the scarcity/absence of soils. This interval is much greater than the time interval that has previously been attributed to the existence of glaciers on the Apuan Alps, which has been assumed to be restricted to Marine Isotope Stage (MIS) 2. Instead, ice cover probably also appeared in the Apuan Alps during MIS 4. The phase of restarting of growth, which may implies the definitive or substantial glacier melts seem to predate the Holocene.

Keywords	Glacier; Pleniglacial; MIS2; MIS3; MIS4; Italy
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Dear Editor,

We would like to submit to *Palaeogeography, Palaeoclimatology, Palaeoecology* the research manuscript titled: **“Speleothem U/Th age constraints for the Last Glacial conditions in the Apuan Alps, northwestern Italy.”**

by Isola I., Ribolini A., Zanchetta G., Bini M., Regattieri E., Drysdale R.N., Hellstrom J.C., Bajo P., Montagna P., Edwige Pons-Branchu

The Apuan Alps, are a high-rainfall mountain chain in northwestern Italy, adjacent to the Mediterranean Sea. Until now, in spite of the widespread evidence of glaciation, the complete lack of geochronological information hinders our understanding of glaciations history and their significance in the framework of the Quaternary cold periods in the Mediterranean area.

We re-examine 293 uranium-thorium ages from 19 speleothems collected in five caves at different elevations in the Apuan Alps, and we use periods of very low or absent speleothems growth as proxy for a chronologic constraint of glaciers presence. According to this timing we can state that unfavorable conditions for speleothems deposition (low temperature) persisted on the Apuan Alps, not only during the MIS 2, but also during the MIS 4. This is consistent with evidence from the Alps and other mountain areas around the world, of a major phase of glaciation during MIS 4.

Hoping this will be of interest for the journal

Sincerely

Ilaria Isola

Comments from the editors and reviewers:

-Reviewer 1

- This is an outstanding contribution that presents important new findings on the nature of glaciations during the last two cold stages (Middle and Late Pleistocene). The paper utilises existing and new U-series data from speleothems in the Apuan Alps in Italy. The findings are very significant because the authors have identified hiatuses corresponding to periods of ice growth/ice cover in both the last two glacial cycles. I have a few minor points to make below which should help strength the paper even further.

Specific Comments

- Introduction. You should cite the recent paper by Baroni et al. 2018 which reports cosmogenic exposure ages from glacial sites very near your area.

Baroni et al. 2018. Last glacial maximum glaciers in the Northern Apennines reflect primarily the influence of southerly storm-tracks in the western Mediterranean. *Quaternary Science Reviews*, Volume 197, 1 October 2018, Pages 352-367.

Right. We included the citation

- Line 122. Please provide a brief definition of the LGM. This is defined as the interval 26.5-20/19 ka (Clark et al. 2009) or 27.5-23.3 ka (Hughes and Gibbard 2015), both definitions within MIS 2. The latter reference is in the reference list but it is not cited in the text.

Clark, P.U., Dyke, A.S., Shakun, J.D., Carlson, A.E., Clark, J., Wohlfarth, B., Mitrovica, J.X., Hostetler, S.W., McCabe, A.M., 2009. The Last Glacial Maximum. *Science* 325, 710-714.

Hughes, P.D., Gibbard, P.L. 2015. A stratigraphical basis for the Last Glacial Maximum (LGM). *Quaternary International* 383, 174-185.

We added the suggested definitions and the references

- Lines 127-129. You note “The presence of these older, cemented deposits, attributed to “Riss glaciation” (i.e. penultimate glaciation), has long been debated but no conclusive results have been forthcoming due to the absence of chronological constraints (Federici, 2005)”. In this section it might be worth citing the work elsewhere in the Apennines that has shown similar types of older cemented moraines have been dated conclusively using U-series to the Riss by Kotarba et al. (2001).

Kotarba, A., Hercman, H., Dramis, F., 2001. On the age of Campo Imperatore glaciations, Gran Sasso massif, central Italy. *Geografia Fisica e Dinamica Quaternaria* 24, 65-69.

Yes. We added this part

- Line 292. (Hughes et al., 2015). This is not cited in the reference list. Do you mean Hughes and Gibbard 2015, which is in the reference list?

Yes. We changed accordingly

- Section 5, Results. In the abstract you state that there was “very low growth between 160 and 135 ka”. But in the results it is not as clear-cut as this. I wondered why you chose the age of 135 ka in the abstract?

It is a misprint. We changed accordingly to the main text

- On this note it is interesting to compare your findings with the lake core evidence from Ioannina where Wilson et al. 2015 found that lake levels started to rise at 132-131 ka, which is c. 3 ka before the start of the interglacial recorded in pollen records. This accords with your timing of the deglaciation of the Apuan Alps indicated in the speleothem records, especially with “speleothem growth at Corchia, after a period of low values, becomes significant between ca.132 and 80 ka”.

Wilson, G.P., et al. 2015. Reconciling diverse lacustrine and terrestrial system response to penultimate deglacial warming in Southern Europe. *Geology* 43, 819-822.

Right. We added this part

- The fact that speleothems appear to start growing again after c. 13 ka is significant. What does this mean for our understanding of climate during the Younger Dryas in Mediterranean? I presume it means that your caves were not ice covered at that time. But again, I presume that this does not preclude there being small glaciers present in the highest cirques during the Younger Dryas?

Yes. It is possible that little glaciers persisted in the higher and more favorable positions, but unfortunately we aren't able to demonstrate this

- Lines 377-379. You note that your hiatuses in speleothem growth are “consistent with evidence of a phase of glaciation during MIS 4 from the Alps and other peri-Mediterranean mountain ranges”. There is now more evidence with ages of 30-40 ka from moraines in southern Greece (Pope et al. 2017) and indeed from Turkey (Sarıkaya et al. 2014) and mean moraine ages of ~50 ka from the High Atlas, Morocco (Hughes et al. 2018). This might need some minor alterations to Figure 5c. It all supports the findings that you present, so this is clearly a good thing for your paper and makes your findings very convincing that your signal of ice growth is relevant right across the Mediterranean mountains.

Hughes, P.D., Fink, D., Rodés, Á., Fenton, C.R., 2018. ^{10}Be and ^{36}Cl exposure ages and palaeoclimatic significance of glaciations in the High Atlas, Morocco. *Quaternary Science Reviews* 180, 193-213.
Pope, R.J., Hughes, P.D., Skourtsos, E., 2017. Glacial history of Mt Chelmos, Peloponnesus, Greece. In: Hughes, P.D., Woodward, J.C. (Eds.) *Quaternary glaciation in the Mediterranean Mountains*. Geological Society of London Special Publications 433, 211-236. <https://doi.org/10.1144/SP433.11>
Sarıkaya et al. 2014. An early advance of glaciers on Mount Akdäg, SW Turkey, before the global Last Glacial Maximum; insights from cosmogenic nuclides and glacier modelling. *Quaternary Science Reviews* 88, 96-109.

Right. We added all the suggested data in figure 5

-Reviewer 2

-
The manuscript by Isola et al. explores an innovative approach to establish the existence and extent of glaciers in the northern Apennines (Apuane Alps) during the Upper Pleistocene. Authors consider a large dataset of U/Th dating on speleothems and correlate the phase of interruption of speleothems growth with phases of glacier expansion or, at least, of diffuse permafrost occurrence. The same data well fit also with inferred surface temperatures. I think the paper is robust and deserves to be published in *Palaeo3*, but it requires some minor to moderate changes before it can be accepted.

General comments:

- In the introduction you discuss the importance of humidity as a trigger of speleothem growth and you cite present-day desert environments. I think that in the case of desert environments you may cite also the case of spring tufa, which occurrence significantly correlates with wet phases.

We added a sentence about tufa and associated references

- The presentation of the procedure followed to obtain MAAT is in the discussion session; I suggest to move this part in the Methods section.

We moved this part in the Method section as suggested

- The correlation between speleothems growth and occurrence of glaciers is convincing and well supported by data, but I would like to explore an alternative explanation for the lack of speleothems at least during MIS 3 and MIS4. I agree that during MIS2 glaciers were present on the Apuan Alps, as confirmed by many geomorphological observations, but their existence during the MIS3–4 periods is less supported by geomorphological data. As you propose, the absence of speleothem is related to a decrease in soil metabolism and a general lack of water. These two preconditions required by speleothems formation are limited in the case of (i) topographic surfaces covered by ice, (ii) soils freezing (permafrost), and (iii) increased environmental aridity. You did not consider the latter factor, but aridity is well attested in North Italy (namely in the Po Plain) during MIS3–4 and before MIS2. Aridity is suggested by extensive loess accumulation. I think that aridity affected also central Italy, and the Apuan Alps may have suffered a phase of cooling as well as aridification. I strongly suggest discussing this possibility in the Discussion session and possibly adding the age of loess formation (proxy for aridity) in Figure 2 or Figure 5. More in general I suggest that speleothem inhibition occurred in both glacial and periglacial conditions.

- Lines 288–298: the discussion on what happened in the Lateglacial is very interesting, as this period is poorly understood. Also in this case, I would suggest considering the possibility that environmental aridity, and not necessarily glacial expansion, tuned speleothems growth. In fact, a number of records indicate the formation of Lateglacial loess along the valleys of the Alps and the Apennines, also not far from the Apuan Alps. The age of these loess sequences seems to correlate with the phases of reduced speleothems growth.

Generally speaking we agree that an increased environmental aridity can inhibit the speleothems growth, but the Apuan Alps has a geographic position so favorable for rainfall, that only a really very strong aridity can produce the affect recorded by the described speleothems. We introduced a new part about this in the Discussion section.

Minor changes:

- line 57: 'soil development' I would say 'soil formation'

Yes. We changed accordingly

- Figure 1: can you add the main circulation systems to the inset?

We added arrows showing the main winds direction

- Figure 2: add indications of all MIS, not only MIS1/Holocene.

We added an upper bar with all the indications

- Lines 278–283: do you relate this to a proper phase of fluvial activity or to a general phase of slope instability?

As it is explained in the txt we think that this phase is related to snow/ice melting and the great availability of debris due to weathering process

- lines 412–415: can you find any correlations with the Alps?

We added information about the Alps in the Discussion section

1 **Speleothem U/Th age constraints for the Last Glacial conditions in the Apuan Alps,**
2 **northwestern Italy.**

3
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20
21 **Abstract**

22
23 During the Quaternary several glaciations occurred in the mountain regions around the
24 Mediterranean and, in recent years, new ages have better constrained their timing. However, this is
25 not the case for the Apuan Alps, a high-rainfall mountain chain adjacent to the Mediterranean Sea.
26 Here, in spite of the widespread evidence for glaciers, the complete lack of geochronological

27 information hinders our understanding of glaciation history. In this paper, we utilize speleothem
28 ages to better constrain the timing of these glacial features. We re-examine 293 uranium-thorium
29 ages from 19 speleothems collected in five caves at different elevations. After a period of very low
30 growth between 160 and 135-132 ka, the analysed speleothems grew almost continuously to ~75 ka,
31 this period was followed by intermittent growth with lower deposition rate and presence of hiatuses
32 until ~12.5 /12 ka. This is consistent with an ice coverage persisting over the Apuan Alps, inhibiting
33 or interrupting the growth of speleothems via the limited availability of groundwater and the
34 scarcity/absence of soils. This interval is much greater than the time interval that has previously
35 been attributed to the existence of glaciers on the Apuan Alps, which has been assumed to be
36 restricted to Marine Isotope Stage (MIS) 2. Instead, ice cover probably also appeared in the Apuan
37 Alps during MIS 4. The phase of restarting of growth, which may implies the definitive or
38 substantial glacier melts seem to predate the Holocene.

39 **Keywords:** Glacier; Pleniglacial; MIS2; MIS3; MIS4; Italy

41 1. Introduction

42
43 Speleothems (i.e. cave carbonate deposits) are multi-proxy paleoclimate archives, which can be
44 accurately dated back to several hundred thousand years before present thanks to U-Th isotope
45 systematic (Edwards et al. 1987; Hellstrom, 2006). Most of the paleoclimate research conducted
46 using speleothems focused on the interpretation of stable isotope and trace element proxy records
47 (e.g. Dykoski et al., 2005; Bar-Matthews et al., 1996; Wang et al., 2008; Regattieri et al., 2016).
48 However, speleothems need certain conditions to grow continuously, such as the presence of liquid
49 water, CO₂ and high Ca concentrations in drip waters (e.g. Atkinson et al., 1986; Gascoyne 1992;
50 Genty et al., 2001). This generally implies relatively wet climate conditions, surface temperature
51 above 0 °C and a well-developed soil above the cave (Gordon et al., 1989; Genty et al., 2001,2005).
52 For instance, in desert environments, phases of speleothem growth are important indicators of past

53 humid periods (Burns et al., 2001; Fleitmann et al., 2011; Vaks et al., 2006, 2007, 2010), as well as
54 tufa, which occurrence significantly correlates with wetter phases (Smith et al., 2004; Cremaschi et
55 al., 2015a). Similarly, the growth of speleothems in permafrost or glacier-dominated environments,
56 often marks interruptions of cold conditions (i.e. Berstad et al., 2002; Spötl and Mangini, 2006;
57 Pons-Branchu et al., 2010; Vaks et al., 2013). This is potentially consistent with degradation of
58 permafrost and the formation of large ice-free surfaces, favoring water infiltration and soil
59 development formation (Varks et al., 2013). Growth hiatuses within a single or multiple
60 speleothems have been interpreted as evidences of particularly dry and/or cold conditions (e.g.
61 Hodges et al., 2008; Genty et al., 2003, Baldini, 2002; Moreno et al., 2010; Mayer et al. 2012, Stoll
62 et al., 2013), even if the nature and duration of hiatuses in speleothems cannot be unambiguously
63 correlated to climatic conditions. However, stacking multiple speleothem records can give more
64 robust information about the relationship between growth cessations and climate (e.g. Stoll et al.,
65 2013).

66 During glacial periods, depressed external air temperatures at high latitudes and/or high altitudes
67 sites may cause extensive freezing at the surface above a cave, strongly limiting water infiltration
68 (Berstad et al., 2002; Genty et al., 2003,2005; Ayalon et al., 2013) and totally inhibiting soil activity
69 (McDermott, 2004). This implies a strong reduction in the CO₂ transfer to the epikarst, a
70 prerequisite for bedrock dissolution and speleothem growth. Therefore, development of glacial
71 conditions (i.e. the presence of glaciers) in the water-infiltrating area over a cave (catchment area
72 from here onwards) would be a factor for reduced or ceased speleothem growth. In such conditions,
73 the only possibility for the deposition of calcite can be related to sulphide oxidation if liquid water
74 is present (Atkinson 1983; Gascoyne and Nelson, 1983; Spötl and Mangini, 2007).

75 In the Mediterranean region there are numerous mountains where geomorphological and
76 geological evidence for the presence of glaciers has been dated (Baroni et al., 2018; Finsinger and
77 Ribolini 2001; Perez-Alberti et al., 2004; Federici et al., 2008; 2012; 2017; Kuhlemann et al., 2008;
78 Hughes et al., 2004, 2011, 2013; Giraudi et al., 2011 Giraudi and Giaccio, 2017; Ribolini al., 2011,

79 2018; Serrano et al. 2012; Hughes and Woodward, 2017 and references therein, Gromig et al.,
80 2018; Hannah et al., 2017; Akçar et al., 2017; Çiner et al., 2017; Sarikaya and Çiner, 2017).
81 However, for some regions, it is still challenging to identify the glacial cycles to which this
82 evidence belongs. This is particularly true for the Apuan Alps, a mid-latitude (44°N) mountain
83 chain in northwestern Italy where, despite the recognition of numerous glacial features, no precise
84 chronological data exist so far.

85 In this paper, we use periods of speleothem deposition as an indicator of ice-free conditions in
86 cave catchment areas, and periods of very low or absent stalagmite growth as proxy for constraining
87 the chronology of glacier presence in the Apuan Alps. To achieve this, we use previously published
88 and new U/Th ages of speleothems from five different caves located at different altitudes (Fig. 1).

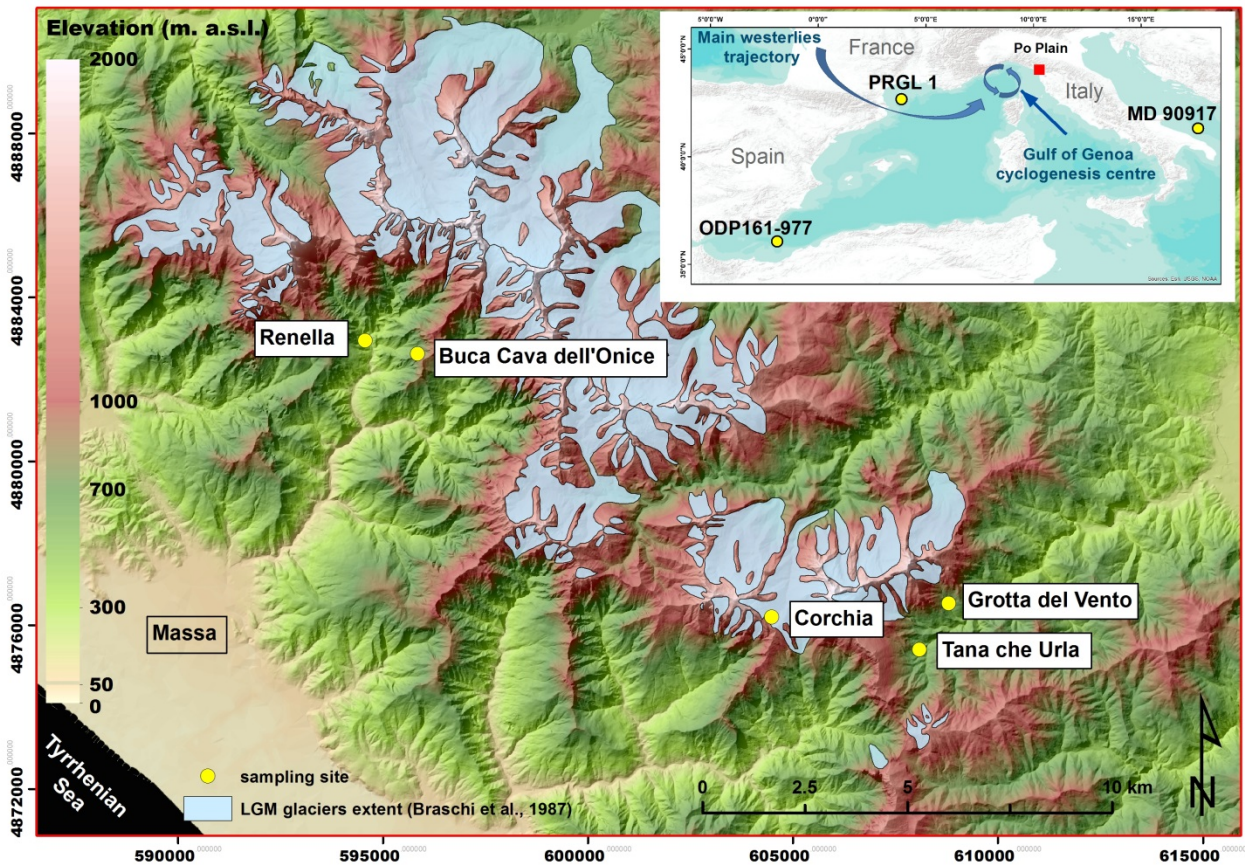
89

90 **2. Site description**

91

92 The Apuan Alps is a NW-SE-oriented mountain range rising abruptly to about 2000 m a.s.l. from
93 the narrow coastal plain bordered by the Ligurian Sea (Fig. 1). The atmospheric circulation is
94 dominated by the important cyclogenesis centre of the Gulf of Genoa (Trigo et al., 2002) and by the
95 humid westerly air masses of North Atlantic provenance (e.g. Reale and Lionello, [20132014](#); Fig.
96 [1](#)). These sources of moisture impact on the Apuan Alps mountain chain, which acts as a natural
97 barrier by forcing an adiabatic rise of air masses, resulting in high precipitation (> 2,500 mm/yr
98 Rapetti and Vittorini, 1994; Piccini et al., 2008). As with most of the Apennine chain, winter
99 rainfall amount, which is the main period of recharge of Apuan caves (Piccini et al., 2008), is
100 strongly regulated by the North Atlantic Oscillation (NAO), with a negative correlations observed
101 between NAO index and winter precipitation (López-Moreno et al., 2011).

102



103
104

Fig. 1. Glacier extent (light blue) during the LGM according to the reconstruction of Braschi et al.

105

(2001). Sampled sites are shown in yellow.

106

107 Two main tectono-metamorphic regional events have determined the structural setting of the
 108 Apuan Alps (e.g. Carmignani and Kligfield, 1990; Molli and Vaselli, 2006), where extensive areas
 109 of carbonate lithologies outcrop (Mesozoic marbles and metadolostones). The actions of glacial,
 110 fluvial and karst processes led to an Alpine-like landscape, where exokarst landforms and
 111 Quaternary cave systems are particularly developed (Piccini, 1998, 2003, 2011).

112 The Apuan Alps is one of the first places amongst the larger Apennine chain where glacial
 113 landforms have been described (Cocchi, 1872; Stopani, 1872). Following these pioneering studies,
 114 in the second half of nineteenth and throughout the twentieth century, further qualitative evidence
 115 of glacial landforms were reported (De Stefani, 1874,1875,1890; Merciai, 1912; Paci, 1935;
 116 Zaccagna, 1937; Tongiorgi and Trevisan, 1940; Beneo, 1945; Masini, 1970; Federici, 1981; 2005;
 117 Braschi al., 1987; Putzolu, 1995, Baroni et al. 2013; 2015). Most of these glacial features are now

118 under risk of destruction due to marble quarrying (Bini, 2005). Despite only traces of small glaciers
119 have been reported on the seaward side of the mountain chain (i.e. erosional landforms and a few
120 scattered glacial deposits, Federici, 1981), features associated with at least nine glaciers have been
121 identified on the northern side. These include lateral and terminal moraines, cirques and polished
122 surfaces. Different phases of expansion have been described and, in some cases, the glacier termini
123 reached exceptionally low elevations (~600 a.s.l.) considering the latitude (Braschi et al., 1987;
124 Jaurand, 1998). The best-preserved glacial features, with relatively unweathered, incoherent detrital
125 materials, are traditionally attributed to the Last Glacial Maximum (LGM) by all previous authors
126 on the basis of their fresh appearance and by analogy with the Apennines and Alpine landforms.
127 The temporal definition of LGM itself is not univocal, depending on the considered references, e.g.
128 Clark et al. (2009) defined the LGM as the duration of sea-level lowstand during the interval 26.5-
129 20/19 ka, while Hughes and Gibbard (2015) as the event between the top (end) of Greenland
130 Interstadial 3 and the base (onset) of Greenland Interstadial 2, spanning the interval 27.540-23.340
131 ka (Greenland Stadial 3). In both cases within MIS2.

132 Inner terminal moraines have been interpreted as the Late-Glacial readvance/stillstand phases
133 (Braschi et al., 1987; Federici, 2005). Cemented glacial deposits, rarely outcropping, have been
134 attributed to a pre-LGM glacial event according to their appearance, altitude and stratigraphic
135 position (Braschi et al. 1987). The presence of these older, cemented deposits, attributed to “Riss
136 glaciation” (i.e. penultimate glaciation), has long been debated but no conclusive results have been
137 forthcoming due to the absence of chronological constraints (Federici, 2005), moreover, Kotarba et
138 al. (2001) have dated in the Apennines, similar types of older cemented moraines to the Riss, using
139 U-series.

140

141 **3. Cave descriptions**

142

143 In the Apuan Alps karst massif, about 2000 cave entrances are now mapped with a composite
144 length of cave development of ~500 km (<http://www.speleotoscana.it>). About 80 caves are carved
145 below the areas potentially covered by glaciers during the Last Glacial expansion, as reconstructed
146 by Braschi et al. (1987). Despite most of the Apuan caves are mainly vertical with active drainage
147 systems, many kilometres of passages are relict and nowadays host carbonate (usually calcite)
148 concretions. Here we focus on speleothems from five caves located at different elevations (Fig. 1).

149 The largest and highest cave is Antro del Corchia (CC), a large complex cave system (~60 km
150 long ~1200 m deep) mainly developed in the Upper Triassic-Liassic marble, dolomitic marble and
151 metadolostone of the Mt Corchia syncline confined by the non-karstifiable, low-permeable rocks of
152 the Paleozoic basement (Piccini et al., 2008).

153 Tana che Urla (TCU) is a small resurgence cave with a permanent stream (~600 m of total length
154 of which almost half is submerged; +45 m of total difference in height), developed at the contact
155 between Paleozoic schist basement and Triassic metadolostone (Regattieri et al., 2012). The
156 entrance is located at ~620 m a.s.l. on the south-eastern side of Panie Massif and functions as an
157 overflow spring during intense rainfall and snowmelt.

158 Grotta Del Vento (GDV) is an almost fossil phreatic cave (~4.5 km total length and over 450 m
159 of elevation change) developed at the contact between the Paleozoic schist basement and Triassic
160 metadolostone. The entrance is located at ~630 m a.s.l. on the south-eastern side of the Panie Massif
161 (Piccini et al., 2003).

162 Buca della Renella (RL) is a small, shallow cave at the confluence of Canale Regolo and the
163 Frigido River (Drysdale et al. 2006; Zhornyak et al., 2011). The entrance is located at ~275 m a.s.l.,
164 only a few metres above the stream confluence. The cave, predominantly horizontal (~200 m
165 length), is carved in Triassic metadolostone close to the contact with the Paleozoic phyllite.

166 Buca Cava dell'Onice (BCO) is a fossil small cave (85 m long and 65 m deep), carved in Lower
167 Jurassic cherty metalimestone. The entrance is located in the northwest side of Mt. Castagnolo at
168 ~700 m a.s.l.

169

170 4. Methods

171

172 We consider 19 speleothems (ages in supplementary material, Table 1), collected from the five
173 caves described above. Nine stalagmites and two cores drilled from flowstones have been collected
174 from the “Galleria delle Stalattiti” of the Corchia Cave. This part of the cave is a near-horizontal
175 chamber carved at ~840 m a.s.l., vertically overlain by ~400 m of rock. The chamber lies hundreds
176 of metres above the present groundwater table and is characterised by a mean annual air
177 temperature of 7.8 °C (Piccini et al., 2008). Five speleothems have already been described in
178 previous papers (CC1, CC5, CC7, CC26, [CC27](#) and CC28, Drysdale et al., 2004, 2005, 2007, 2009;
179 Zanchetta et al., 2007; Bajo et al., 2017; [Isola et al., 2018](#)). We present new ages for stalagmites
180 CC4, CC7, CC22, CC27, CC53, and two flowstone cores CD4 and CD20.

181 At TCU two cores have been drilled from the same flowstone in the main gallery, about 100 m
182 from the entrance. The mean air temperature is 10.2 °C (Regattieri et al., 2014a). Speleothem
183 stratigraphy, chronology and geochemistry have been discussed by Regattieri et al. (2012, 2014a,
184 2016).

185 Two cores have been drilled from different flowstones located in the middle level of GDV
186 (Piccini et al., 2003).

187 At RL, three samples have been collected in different stratigraphic levels from a fan-like
188 flowstone deposited at the margin of an epiphreatic passage in the inner part of the cave (average air
189 temperature ca. 12 °C) and their stratigraphy and chronology have been discussed by Drysdale et al.
190 (2006) and Zhornyak et al. (2011).

191 Only one stalagmite (BCO1) has been collected in BCO, in a fossil gently dipping gallery, about
192 30 m from the entrance.

193 All ages from Corchia, Tana Che Urla and Renella have been obtained following the same
194 procedure at the University of Melbourne (Victoria, Australia). Briefly, uranium and thorium

195 isotopes were analysed using a Nu Instruments Multicollector Inductively Coupled Plasma–Mass
196 Spectrometer (MC-ICP-MS), according to the analytical method and corrections (where necessary)
197 described by Hellstrom (2003, 2006).

198 Age-depth models and speleothems growth rates were obtained from a Monte Carlo-derived
199 method (described in Drysdale et al. 2005 and Scholz et al. 2012). Dating methods for Grotta del
200 Vento are discussed by Piccini et al. (2003).

201 BCO1 $^{230}\text{Th}/^{234}\text{U}$ ages were determined at LSCE from 9 subsamples collected along the growth
202 axis of the stalagmite. Subsamples were crushed in an agate pestle and aliquots of typically ~400
203 mg were rinsed twice with MilliQ, dissolved in diluted HCl (~10%) and equilibrated with a mixed
204 ^{236}U - ^{233}U - ^{229}Th spike that was calibrated against a Harwell Uraninite solution (HU-1) assumed to
205 be at secular equilibrium. Uranium and thorium were purified and separated using Eichrom
206 UTEVA® and pre-filter resins in nitric media following a procedure modified from Pons-Branchu
207 et al. (2005) and Douville et al. (2010). The isotopes of uranium and thorium were analysed by
208 solution multi-collector ICPMS using a ThermoScientific Neptune^{Plus} hosted at the Laboratoire des
209 Sciences du Climat et de l'Environnement (LSCE, Gif-sur-Yvette, France) following the procedure
210 described by Pons-Branchu et al. (2014). Based on the measured atomic ratios, the $^{230}\text{Th}/^{234}\text{U}$ ages
211 were calculated through iterative age estimation (Ludwig and Titterton, 1994) using the ^{230}Th ,
212 ^{234}U and ^{238}U decay constants of Cheng et al. (2013) and Jaffey et al. (1971). An initial $^{230}\text{Th}/^{232}\text{Th}$
213 activity ratio of $0.9\pm 50\%$ was used for correction of the non-radiogenic, detrital ^{230}Th fraction.

214 In order to highlight the higher and lower periods of speleothems deposition rate, we calculate
215 the growth rate for all the considered speleothems from Corchia Cave and for speleothems from
216 lower altitude caves, cumulated at 1000-yr intervals. Uncertainties in the growth rates were
217 obtained using the standard errors propagation.

218 It is worth estimating potential changes in air temperature at the altitude of the different
219 infiltration areas of the four caves (we ignore BCO because the sole sampled stalagmite covers only

220 MIS 5), in order to understand causes of any leads and lags in speleothem growth phases in the
221 context of likely glacier development. The following calculations represent a first-order estimate
222 that can be used to underpin the interpretations proposed below.

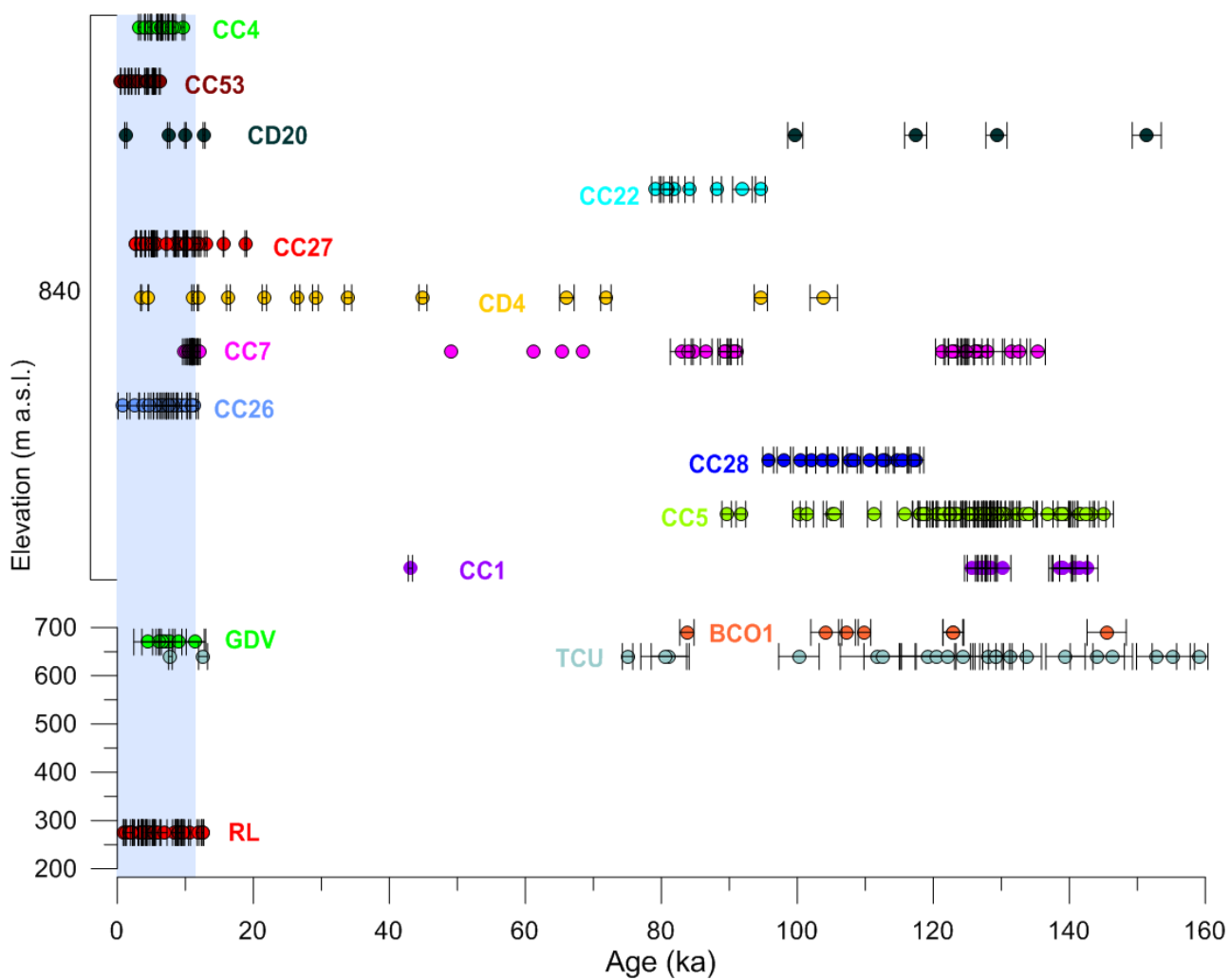
223 Speleothem growth patterns described above were compared with sea-surface temperature (SST)
224 series from three Mediterranean cores: ODP 161-977 (Martrat et al., 2014) spanning all over the
225 considered period; PRGL1 (Cortina et al., 2015), the nearest core to the Apuan Alps; and MD
226 90917 (Siani et al 2013), a high-resolution dataset covering the last ~25 kyr (cores locations are
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228 unsaturation indices (Müller et al., 1998; Conte et al., 2006) and planktonic foraminifera
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230 in the Alboran Sea (ODP191-977) are more related to spring-summer conditions (Martrat et al.,
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232 signal showing a colder imprint than other more southern SST records (Rigual-Hernández et al.,
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234 productive period during the spring and the synchronous bloom of *G. bulloides* in the
235 Mediterranean Sea (Pujol and Vergnaud-Grazzini, 1995; Siani et al., 2013). In performing the
236 ocean-cave comparisons, we make several assumptions. First, we assume that the difference in
237 mean annual SST (MASST) between the selected marine cores and the MASST offshore of the
238 Apuan Alps (Locarnini et al., 2010) was constant through time; second, that the difference between
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240 Apuan Alps (<http://www.autorita.bacinoserchio.it/>) was constant through time; third, we adopt an
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243 elevations of the water recharge area feeding the caves, two values are considered: the elevation of
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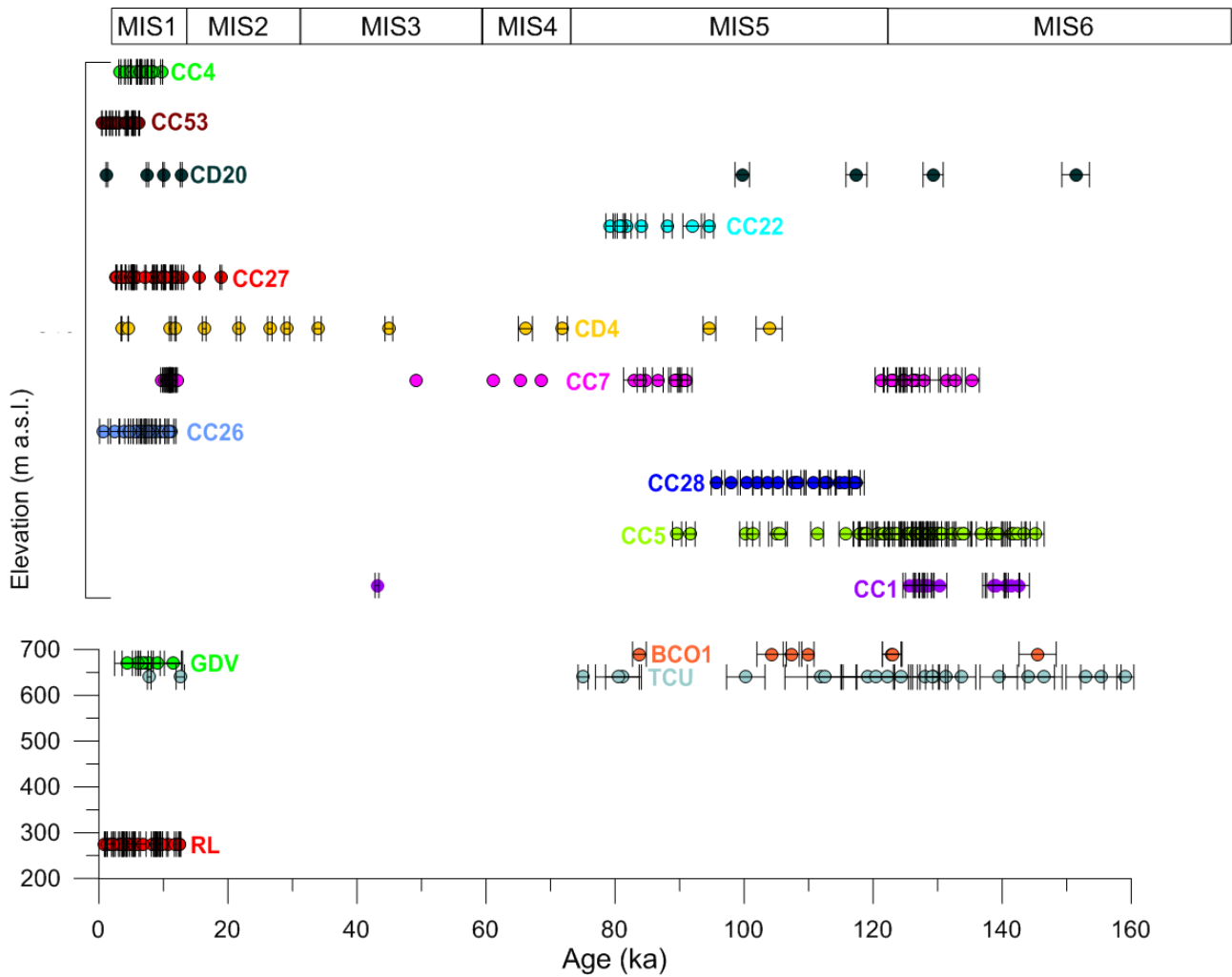
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248 whole hydrographic basin upvalley the cave. A more refined estimate of MAAT could incorporate
249 also the effects of sea-level lowering during glaciations and tectonic uplift (estimates to be 0.6
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251 negligible temperature variations. Indeed, the maximum increase in relative elevation of the
252 infiltration area above the contemporary sea level is during the LGM (e.g. ca. 100-120 m for the
253 LGM, Lambeck and Bard 2000; Lambeck et al., 2014), corresponding to a total temperature
254 variation (ΔT) of about -0.6°C , while the lowest, during MIS6, corresponds to ΔT of about $+0.1^{\circ}\text{C}$.

255

256 **5. Results**

257 For the 19 speleothems considered, the dataset comprises 293 U/Th ages (supplementary
258 material Table 1). In Figure 2, the ages are plotted vs the elevation of the sampling sites.





260

261 Fig. 2. $^{293}\text{U}/\text{Th}$ ages from 19 Apuan speleothems over the past 160 ka (2σ error bars) vs the
 262 elevation of sampling sites. Details on dating can be found in the original papers (Drysedale et al.,
 263 2004; 2009; Regattieri et al., 2012, 2014a, 2014b; Piccini et al, 2003; Drysdale et al., 2006;
 264 Zhornyak et al 2011, Isola et al., 2018). ~~The Holocene is shown in light blue shading.~~

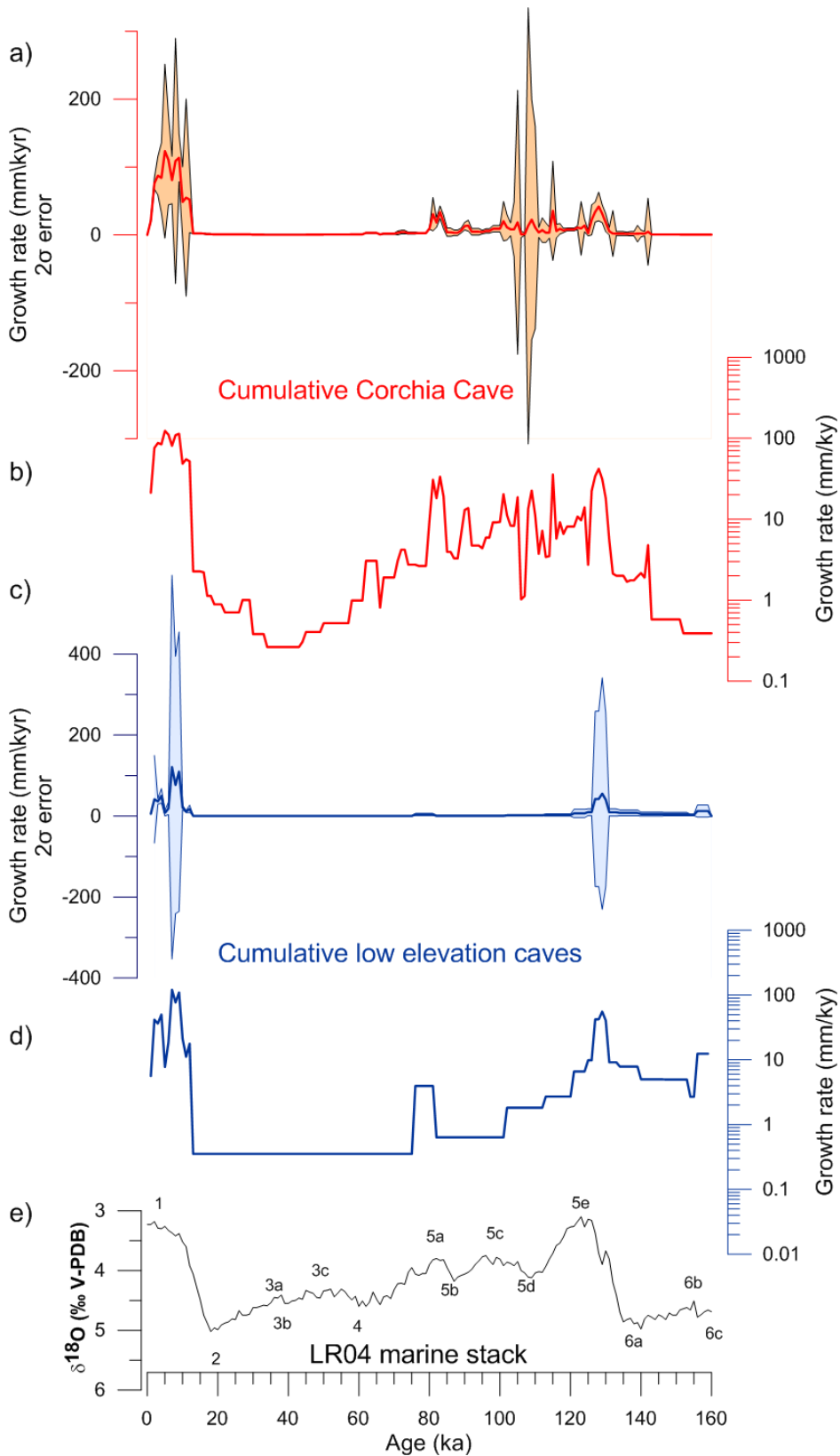
265

266 Despite some differences, there are many common features. All caves show a major episodes of
 267 calcite deposition during the Holocene and the latest Late Glacial. Available samples from RL and
 268 GDV do not show phases of growth older than this period and the studied concretions cover, or are
 269 stratigraphically located over, thick fluvial deposits. TCU shows continuous deposition between ca.
 270 160 and 121 ka (Regattieri et al., 2014a), followed by a discontinuous deposition to ca. 75 ka, after
 271 which there is a long hiatus from ca. 75 to 12.5 ka. The most relevant record is represented by

272 Corchia Cave, showing almost continuous calcite deposition from ca. 145 to 79 ka. A long lag of
273 growth, with only very reduced phases of calcite deposition is evident during the Pleniglacial. The
274 stalagmite from BCO has an intermittent growth phase throughout MIS 5, confirming a
275 speleothems deposition phase during this period in the Apuan Alps. No information about other
276 periods including the Holocene are available for this site until now, but we cannot exclude further
277 growing phases, in fact, a single speleothem cannot be considered representative of the whole
278 speleothems growth history of this cave.

279 More interesting than the simple ages distribution is the cumulative growth curve which takes
280 into account the propagation of age errors for Corchia cave speleothems and for the other studied
281 caves located at lower altitudes (Fig. 3). It is evident that speleothem growth at Corchia, after a
282 period of low values, becomes significant between ca. 132 and 80 ka (curve b in figure 3), after
283 which it decreases abruptly until ca. 60 ka, and then almost stops for ca. 44 kyr. Continuous growth
284 resumes at ca. 15 ka and increases later at ca. 12 ka.

285 The growth of speleothems belonging to the lower-altitude caves evidently decreases after 130
286 ka, reaching a local minimum at ca. 100 ka. An increase of composite growth rate is then observed
287 around 80 ka, followed by an interruption in calcite deposition up to 13 ka.

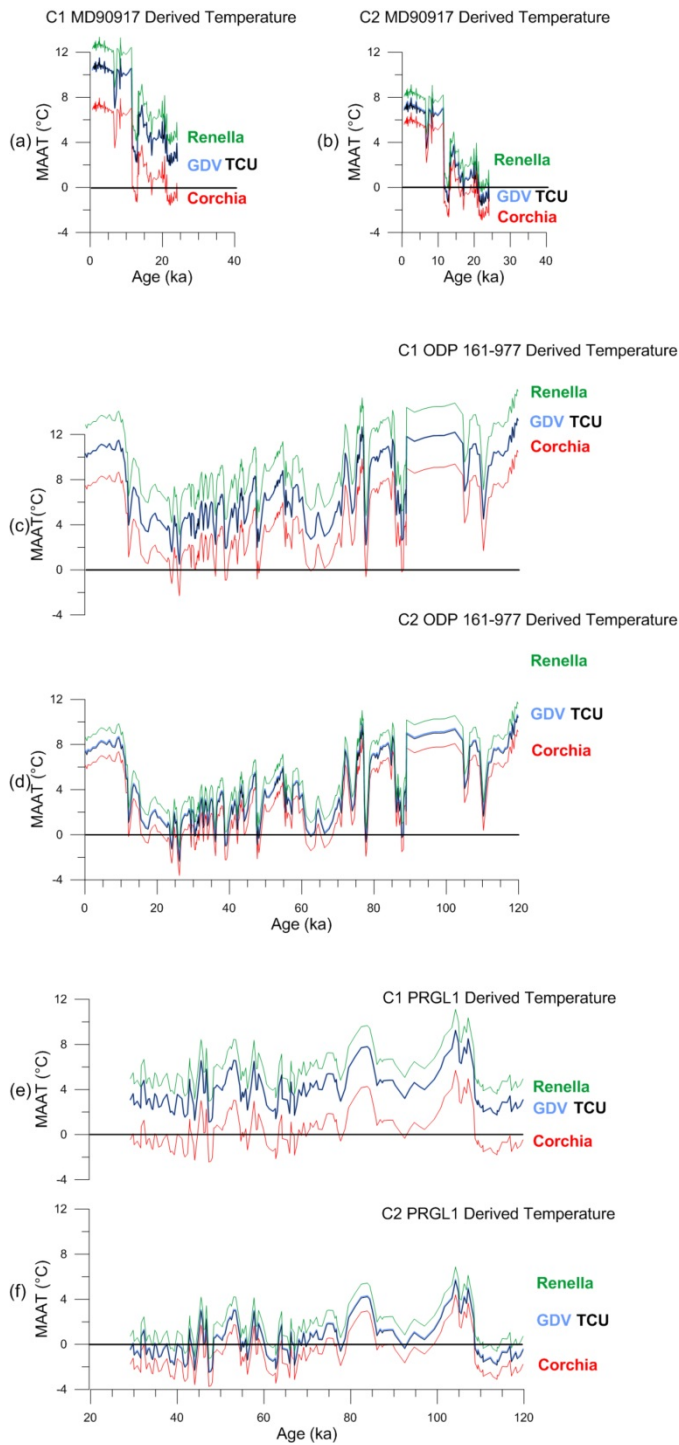


288

289 Fig. 3. Cumulative growth curves (i.e. the total speleothem growth calculated for all the considered
 290 speleothems and binned into 1000-yr intervals) for Corchia Cave speleothems (red lines a and b)
 291 and for the caves located at lower altitude (blue lines, c and d), the associated errors in shaded

292 strips. Note the linear (b and d) and the logarithmic (a and c) scales in the vertical axes. Shown in
293 black (e) is the LR04 stack (Lisiecki and Raymo 2005) from benthic foraminifera $\delta^{18}\text{O}$ data.
294 Numbers indicate Marine Isotope Substages.

295 Figure 4 shows the estimation of MAAT as described above. Due to the very similar geographic
296 location (fig. 1), GDV and TCU temperature estimates are almost indistinguishable in all the
297 reconstructions (blue and black lines in Fig. 4). The MAAT curves obtained for each cave using
298 MD90917 data (~ 0.6 - 24 ka interval), including the two different infiltrating elevations (C1 and
299 C2), are shown respectively in Figure 4a and b. The MAAT curves obtained using ODP 161-977
300 data (~ 0.06 - 120 ka interval) are shown respectively in figure 4c and d. The MAAT curves
301 obtained using PRGL1 data (~ 29 - 120 ka interval) are shown respectively in Figure 4e and f. In
302 the C1 cases (Fig. 4a, c, e) only the CC series displays excursions where MAAT is $<0^{\circ}\text{C}$. In the C2
303 scenario (Fig. 4b, d, f), GDV and TCU areas also show periods when MAAT is $<0^{\circ}\text{C}$. As expected,
304 the temperature above RL is the mildest.



305

306 Fig. 4. MAAT curves derived for all the caves from MD90917: (a) C1 (see the text for explanation),
 307 (b) C2; from ODP 161-977, (c) C1 and (d) C2 ; and from PRGL1, (e) C1 and (f) C2. Black
 308 horizontal lines indicate the 0°C MAAT.

309

310 6. Discussion

311

312 According to the presented data, there is a long interval during the Pleniglacial where calcite
313 deposition at CC and TCU caves slows down or is absent (fig. 3). This interval roughly corresponds
314 to an extensive loess deposition in the Po Plain (Fig. 1 and 5; Cremaschi et al., 2015b; Zarboni et
315 al., 2015). This marks a wide phase of aridity in northern Italy that might have involved also the
316 Apuan Alps, but probably with different effects. Nowadays the mean annual precipitation in Po
317 Plain does not exceed 1000 mm (www.arpalombardia.it), while the Apuan Alps are characterized
318 by higher precipitations exceeding 2500mm/yr (Rapetti and Vittorini, 1994; Piccini et al., 2008).
319 The different geographic position underpinning the actual rainfall regime, certainly played a role in
320 the past too, mitigating the drought effect in the Apuan Alps. Moreover, it is very difficult to
321 explain in a dry regime, the strong fluvial activity suggested, in the low elevation caves, by the
322 presence of thick fluvial deposits, stratigraphically older than the phase of speleothems regrowth
323 (about 12.5 ka). The resumption of calcite deposition is ~~instead~~, almost coeval in all caves, with
324 evidence of sustained growth slightly preceding the beginning of the Holocene, ~~instead~~, the period
325 of deposition decrease, at the end of MIS5, is not coeval between TCU and CC. However, the CC
326 dataset is more extensive because the number of analyzed speleothems.

327 Moreover, this cave is very close to areas of the catchment occupied by glaciers in the Last
328 Glacial (Fig.1). Specifically, the recharge basin of the “Galleria delle Stalattiti”, mainly consisting
329 of the higher slopes of the northern side of Monte Corchia, was covered by a glacier according to
330 Braschi et al. (1987). The most obvious process responsible for the strong reduction of speleothem
331 growth at the “Galleria delle Stalattiti” would thus be a persistent ice cover. This long interval of
332 extremely reduced or absent calcite deposition covers most of MIS4-MIS2, a time span more
333 extensive than the classical attribution of Apuan glacial activity to the only Last Glacial Maximum
334 (Federici, 2005; Jaurand 1998). This hypothesis is supported by evidence of moraines formed
335 during MIS3 and MIS4 glacier expansions in the Central Apennines (Giraudi et al. 2011; Giraudi
336 and Giaccio, 2015, 2017). In this regard, the Corchia speleothems cumulative growth curve delimits
337 for the first time the temporal interval of possible glacier existence over the Apuan Alps or

338 temperature lower enough to inhibit speleothem growth at higher altitudes. Although the
339 chronological constraints are indirect because they do not refer to moraine ages, there is
340 considerable robustness in the speleothem-age dataset. However, cessation of speleothem growth
341 can be related to combined effect of low air temperature, soil-CO₂ deficiency and reduction of
342 dripping. These conditions might have been present without invoking continuous ice covering of the
343 slopes, i.e. where permafrost inhibits water infiltration and following dryer cave conditions. In the
344 nearby Apennine chain, landforms diagnostic of permanently frozen ground are rare and essentially
345 limited to the southern latitudes, where higher elevations are reached (Kotarba et al., 2001; Giraudi
346 et al., 2011; Oliva et al., 2018).

347 The absence of speleothem growth in the other caves during the glacial period cannot be directly
348 related to the reconstructed ice cover (Braschi et al. 1987) because of their lower altitudes. A recent
349 geomorphology synthesis of the Apuan Alps (Baroni et al., 2015) does not report glacial deposits
350 indicative of a glacial cover of the lower caves.

351 At RL and GDV, the presence of thick fluvial deposits stratigraphically older than the phase of
352 speleothem growth (about 12.5 ka) suggests strong fluvial activity during the glacial period, which
353 would inhibit speleothem growth (Zhornyak et al., 2011). This strong activity might be related to
354 seasonal snow/ice melting and abundant availability of loose material coming from physical
355 weathering. A similar explanation is also valid for the long deposition hiatus at TCU (about 12.6-75
356 ka). A continuous glacier cover nearby is unlikely throughout this period.

357 ~~Speleothem growth, which was possible for most of MIS5, decreases or ceases during MIS4~~
358 ~~(Fig. 3)~~

359 It is noteworthy that both CC and TCU show speleothem deposition during part of MIS6,
360 suggesting that temperatures over the Apuan Alps during this period were not as low as those for
361 MIS2-4. ~~The transition between the penultimate glacial and the last interglacial, is marked by the~~
362 ~~abrupt increase of speleothems growth about 132 ka, as already described for CC by Drysdale et al.,~~
363 ~~(2009) and for TCU by Regattieri et al. (2014a) and in agreement with other proxies like the more~~

364 [umid-temperate vegetation in Southern Europe \(Sanchez –Goñi et al., 1999\)](#), or the high level in
365 [Ioannina Lake \(Wilson et al., 2015\)](#). Then, speleothem growth, which was possible for most of
366 [MIS5, decreases or ceases during the interval MIS2-4 \(Fig. 3\)](#).

367 The almost coeval resumption of calcite deposition in all the investigated caves requires further
368 discussion. At CC, some U/Th ages indicate that a small growth phase also occurs at the beginning
369 of the Late Glacial, but significant growth did not occur before ca. 12 ka. However, given the
370 southern latitude and the relatively low altitude of the Apuan Alps, it is hard to argue that glaciers
371 still existed during the middle part of the Lateglacial (Hughes ~~et al~~ and Gibbard, 2015). Based on
372 deglacial trends in CC speleothem carbon isotope composition, Drysdale et al. (2004) and Zanchetta
373 et al. (2007) have suggested that soil development at CC is delayed after glacial periods due to
374 intense erosive processes active at high altitudes during the first phases of climatic amelioration,
375 under conditions of increasing rainfall. After the definitive ice melting at the end of MIS2, cold but
376 ice-free conditions and delayed soil development would have retarded significant speleothem
377 growth, which did not accelerate until ca. 12 ka. The fundamental elements of this hypothesis are
378 supported by Bajo et al. (2017), who also invoked near-closed-system conditions and sulphuric-acid
379 dissolution as additional processes to carbonate-acid dissolution (from soil-sourced CO₂) to explain
380 the unusually high carbon isotope composition of CC26 stalagmite during late Late Glacial and the
381 beginning of Holocene. However, an implication of this study is that growth was unsuitable before
382 ca. 12 ka, in spite of the possibility of sulphuric-acid dissolution substituting for carbonic-acid
383 dissolution, thus suggesting that speleothem growth was inhibited by cold climate conditions. From
384 the same stalagmite, Regattieri et al. (2014b) concluded, based on the presence of an aragonite layer
385 at the base of the CC26 succession and on trace element/Ca ratios (Ba, Sr, Mg) changes, that the
386 start of growth of CC26 was related to a re-opening of a dripping point due to the replenishing of
387 the cave plumbing system previously interrupted by reduced recharge.

388 Whilst we cannot detail the point at which resumption of speleothem growth occurred in CC
389 during the Late Glacial, conditions generally favoring speleothem growth seem to have occurred

390 when the plumbing system was newly recharged significantly even if ice cover should have already
391 disappeared. It is probably not obvious that these conditions did not concomitantly occur in upper
392 and lower altitude caves. Studying the transition between MIS6 and MIS5 at TCU Regattieri et al.
393 (2014a) suggested that the delay in soil development as inferred for CC is not supported for TCU,
394 due to its lower altitude. This may suggest that an earlier resumption of speleothem growth would
395 be anticipated at lower altitude compared to CC. Data collected until now agree with this hypothesis
396 (Fig. 3). We suggest that significant resumption of calcite deposition at CC can occur only when
397 soil development and drip recharge are restored to a certain level.

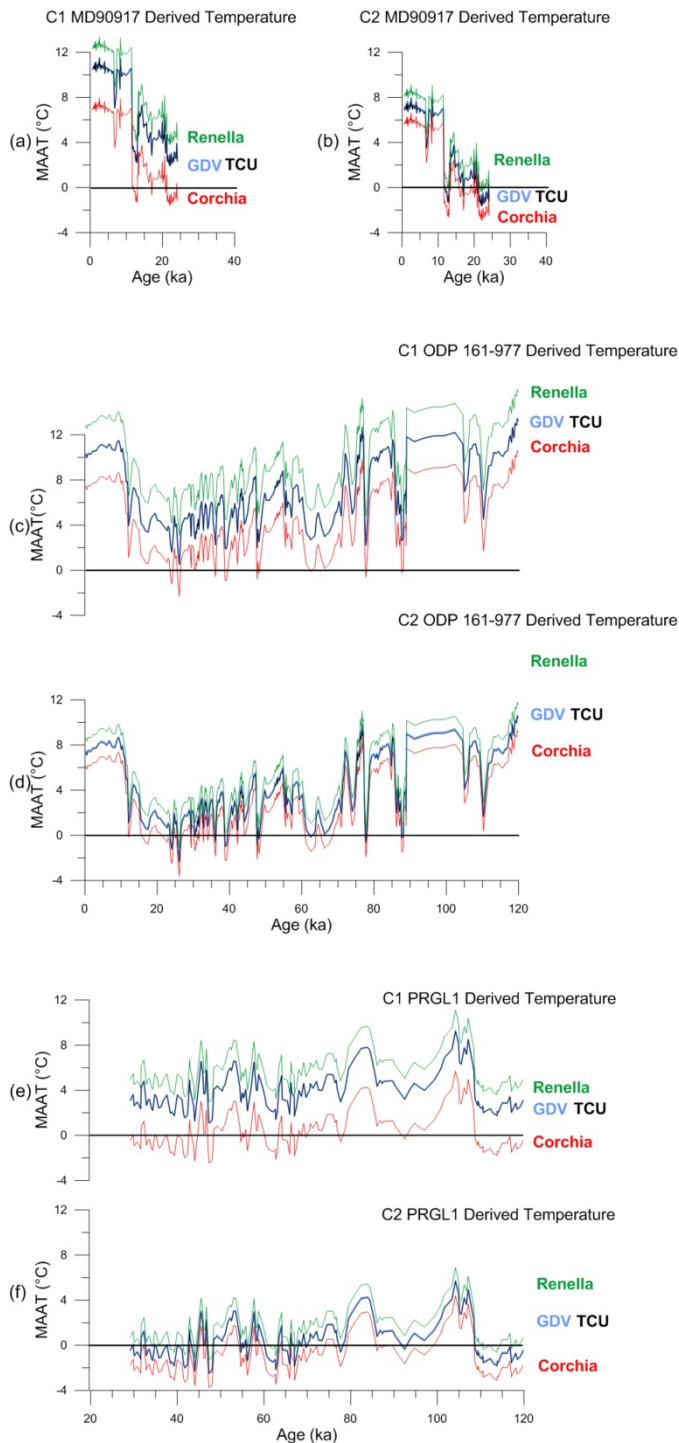
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431 negligible temperature variations. Indeed, the maximum increase in relative elevation of the
432 infiltration area above the contemporary sea level is during the LGM (e.g. ca. $100\text{--}120\text{ m}$ for the
433 LGM, Lambeck and Bard 2000; Lambeck et al., 2014), corresponding to a total temperature
434 variation (ΔT) of about -0.6°C , while the lowest, during MIS6, corresponds to ΔT of about $+0.1^{\circ}\text{C}$.

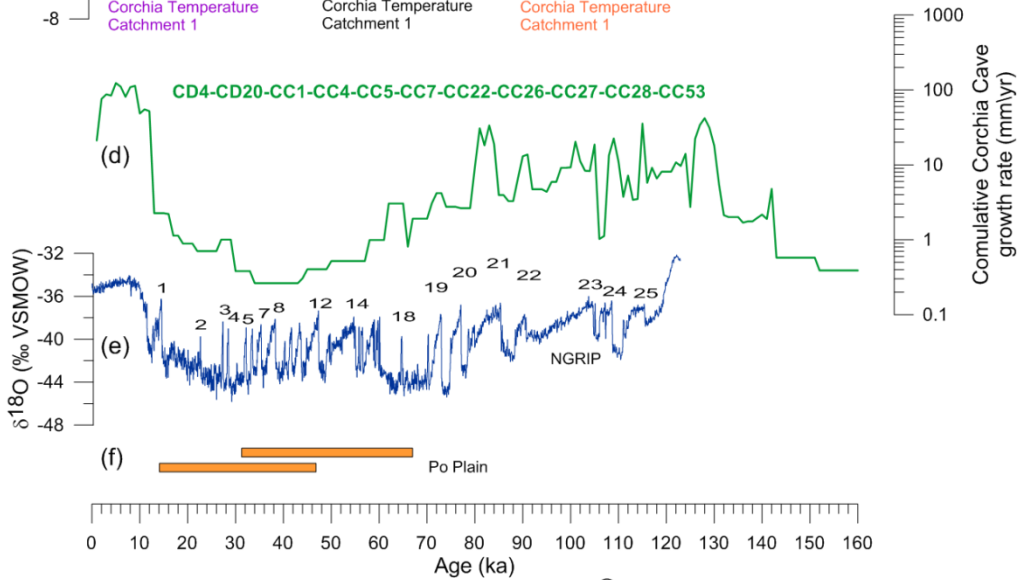
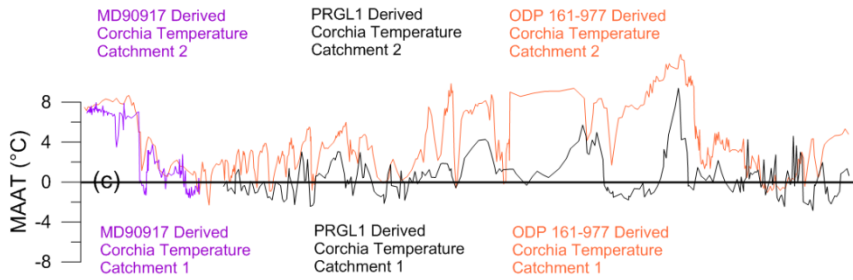
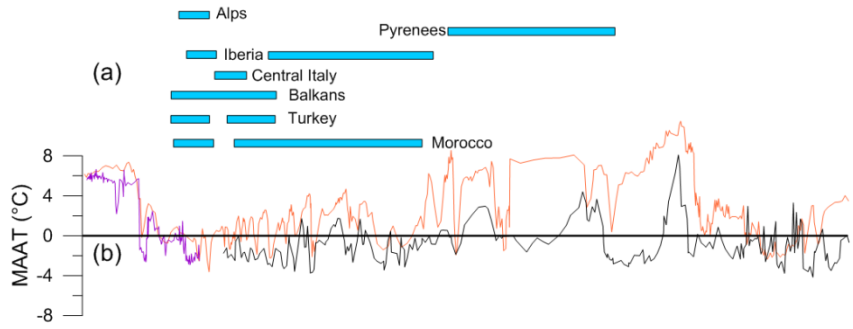
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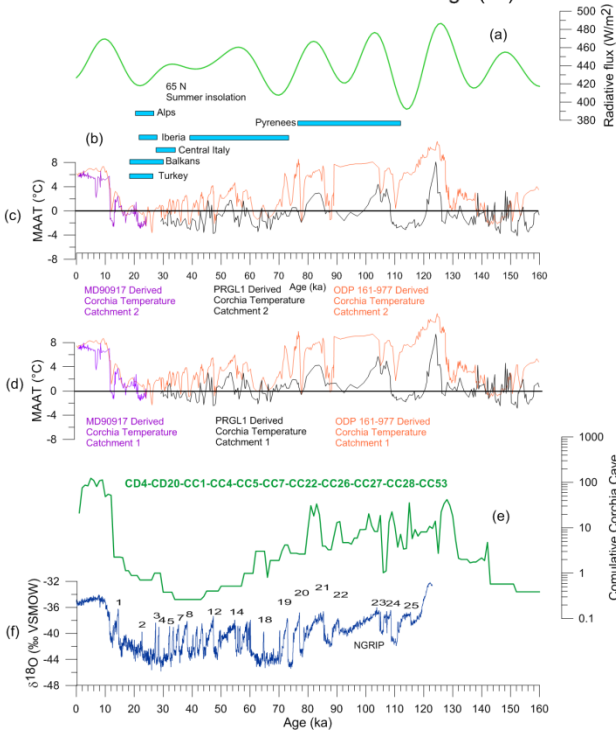


445
 446 Fig. 4. MAAT curves derived for all the caves from MD90917: (a) C1 (see the text for explanation);
 447 (b) C2; from ODP 161-977, (c) C1 and (d) C2; and from PRGL1, (e) C1 and (f) C2. Black
 448 horizontal lines indicate the 0°C MAAT.

450 According to the timing of speleothem growth and MAAT estimations, it can be seen that above CC
451 both the C1 and C2 MAAT reconstructions experienced a number of long-lasting sub-zero periods.
452 Indeed, the Equilibrium Line Altitude (ELA) of 1200-1300 m a.s.l. in the Apuan Alps during the
453 MIS 2 (Merciai, 1912; Braschi et al., 1987; Jaurand, 1998) supports the hypothesis that in the CC
454 catchments (both C1 1240 m a.s.l. and C2 1460m a.s.l.) glaciers were existing. These unfavorable
455 conditions for speleothem deposition (MAAT subzero or near 0 °C) persisted, discontinuously,
456 above CC not only during the MIS 2, but also during MIS 4 (Fig. 4). This is consistent with
457 evidence of a phase of glaciation during MIS 4 from the Alps and other peri-Mediterranean
458 mountain ranges (Hughes et al. 2013; 2018; Sarikya et al., 2014; Pope et al., 2017) (blue bars in
459 Fig. 5ba). There are evidences of very small growth phases during MIS4-MIS2 (Fig. 2, 3, 5ed) that
460 may have occurred during warming reversals, most likely associated with the millennia-scale
461 interstadials observed in the Greenland ice-core records (i.e. Dansgaard-Oeschger events,
462 Dansgaard et al., 1993; Fig. 5fe), but also indicated by SST changes in the western Mediterranean
463 (Cacho et al., 1999).



464



465

466 Fig. 5. Northern Hemisphere proxies for the Pleniglaciation: ~~a) summer insolation curve at 65°N~~
467 ~~(Berger and Loutre, 1991);~~ **ba**) the approximate timing of maximum extent of glaciers in the
468 Mediterranean mountains (after Hughes et al., 2013; 2018; Sarikya et al., 2014; Pope et al., 2017);
469 **eb**) the derived MAAT curves in the C2 case above the Galleria delle Stalattiti: red curve is derived
470 from ODP 161-977, purple curve from MD90917 data, black curve is derived from PRGL1 data.
471 Horizontal black line indicates the 0°C MAAT; **ec**) the derived MAAT curves in the C1 case above
472 the Galleria delle Stalattiti: red curve is derived from ODP 161-977, purple curve from MD90917
473 data, black curve is derived from PRGL1 data and horizontal black line indicates the 0°C MAAT;
474 **ed**) Cumulative growth curve calculated for interval of 1000 yr for Corchia cave speleothems (green
475 line); **fe**) NGRIP $\delta^{18}\text{O}$ values from North Greenland Ice Core (North Greenland Ice Core Project
476 members, 2004). Numbers over the graph mark the position of the Dansgaard-Oeschger
477 interstadials (Dansgaard et al., 1993); **f**) **approximate interval of extensive loess accumulation in Po**
478 **Plain (after Cremaschi et al., 2015b; Zerboni et al., 2015).**

479

480 In general, pollen data from the western Mediterranean show that cold SSTs characteristic of
481 Greenland stadials were contemporaneous with the expansion of semi-desert or steppic vegetation.
482 On the contrary, Greenland interstadials were synchronous with the expansion of open forests
483 (Sánchez-Goñi et al., 2008). Presumably, phases of forest expansion prompted the development of
484 deeper soils. However, the very limited speleothem growth and the available ages prevent a precise
485 correlation with the warmer intervals. On the other hand, according to Atkinson (1983) and Spötl et
486 al. (2007), speleothems may form, although with a lower rate of growth, in caves overlain by
487 temperate glaciers with available glacial meltwater infiltrating the karst system. The very low
488 partial pressure of CO_2 in such water due to the absence of soil can be balanced by pyrite oxidation
489 during the flow paths through fractures, providing the dissolved carbonate necessary for speleothem
490 deposition. This process is demonstrably present at Corchia (Piccini et al., 2008; Bajo et al., 2017)
491 and can be responsible of the very thin deposition layers included between two hiatus.

492 Summing up, growth histories and temperature estimation support the hypothesis that
493 interruptions or strong reductions in speleothem growth at Corchia could have been due to a
494 combination of glaciation and/or ice-free but cold temperatures with low or absent soil activity.

495 Glacial phases occurring in this period in the mountain regions around the Mediterranean are
496 confirmed by multi-proxy analyses carried out on the glacial lacustrine sequences, as evidenced by
497 a maximum of glacier expansion of Pyrenean glaciers before 30 ka (Garcia-Riuz et al., 2003) and in
498 Picos de Europa dated at about 40 ka (Moreno et al. 2010; Serrano et al., 2012). Consistently in the
499 central Apennines, tephra layers in lacustrine sediments bracket the maximum glacier extent
500 between 33-27 ka, not a period of maximum cold conditions but one that was humid (Giraudi,
501 2012). Regarding the Alps there is a controversial discussion about glaciers extent before MIS 2
502 when the maximum expansion was reached (Ivy-Ochs et al. 2008; Preusser et al. 2011; Hughes et
503 al. 2013). The uncertainties mainly derive from the fact that the majority of terrestrial records (i.e.
504 glacial deposit, pro-glacial outwash, river terrace, lake deposit) were eroded, or at least remoulded
505 or reworked, when MIS 2 glaciers invaded the valleys and in many cases spread out in piedmont
506 lowlands. Indeed, while in the Eastern Alps no evidence for the presence of pre-MIS 2 glaciers are
507 clearly found (Ivy-Ochs et al., 2008), an important glaciation reaching the lowlands of the Western
508 Alps during MIS 4 is documented by OSL dating of pro-glacial outwash (Preusser et al. 2007) and
509 lake sediments (Link and Preusser 2006).

510 ~~This~~ The above described asynchronicity is not contradictory a priori, because it may be
511 explained by the short reaction times of small glaciers (like those of most of the Mediterranean
512 mountains) to changes in precipitation and/or temperature. Further sampling may reveal sporadic or
513 very slow episodes of growth during this time interval, allowing better and more precise
514 interpretations of MIS 3 climatic oscillations.

515

516

517 **7. Conclusions**

518

519 Speleothems in the Apuan Alps grew near-continuously during most of MIS5 (from ca. 130 ka to
520 about 75 ka), then experienced a long period of absent to very low deposition covering the MIS4 to
521 MIS2 interval, until ~12.5 ka. This growth history can be explained as the result of development of
522 glaciers at high elevations, not necessarily continuously existent for the whole considered period,
523 and/or low soil development and near-zero temperatures. The speleothem growth curves may
524 suggest that glacial expansion in the Apuan Alps was not only limited to the LGM (MIS2) but also
525 to MIS3 and MIS4, analogous to other peri-Mediterranean mountains.

526 During spring and summer ice-melting periods, high water discharges occurred and caves
527 located at lower altitudes were flooded, resulting in phreatic conditions, inhibiting speleothem
528 deposition and infilling caves with thick fluvial deposits.

529 There is no evidence of significant growth during interstadials within MIS3, which supports the
530 fact that no particularly warm conditions occurred in the Apuan Alps during those climatic events.

531 Sustained speleothem growth resumed close to the beginning of the Holocene both at lower and
532 high-altitude caves. This can be interpreted as the result of a definitive melting of ice cover and the
533 consequent progressive soil development at high altitude, and the stabilisation of loose glacial
534 debris and reduced flooding for lower altitude caves.

535 These data indicate the potential for cumulative speleothems growth curve to chronologically
536 constrain large-scale geomorphological processes, like the timing of glaciers presence. However,
537 other processes can locally bias the speleothem growth (i.e. cold-non glacial conditions,
538 absence/limited soils), suggesting that this approach is restricted to a first-order approximation
539 which should be followed by a direct dating of glacial features. Moreover, the adoption of
540 speleothem growth curves needs to be verified in other mountain ranges, where U/Th dating can be
541 used for constraining glacial events.

542

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544

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556

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962 **List of tables (supplementary materials)**

963 Tab. 1 U/Th data used for the age distribution

964 Tab. 2 Elevation used to to estimate the MAAT in the speleothem catchment areas (C1, C2)

965

Highlights

speleothem growth curves used to better constrain the timing of glacial features

293 U\Th ages from 19 speleothems collected in five caves at different elevations

glacial expansion in the Apuan Alps was not only limited to the MIS2

1 **Speleothem U/Th age constraints for the Last Glacial conditions in the Apuan Alps,**
2 **northwestern Italy.**

3
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20
21 **Abstract**

22
23 During the Quaternary several glaciations occurred in the mountain regions around the
24 Mediterranean and, in recent years, new ages have better constrained their timing. However, this is
25 not the case for the Apuan Alps, a high-rainfall mountain chain adjacent to the Mediterranean Sea.
26 Here, in spite of the widespread evidence for glaciers, the complete lack of geochronological

27 information hinders our understanding of glaciation history. In this paper, we utilize speleothem
28 ages to better constrain the timing of these glacial features. We re-examine 293 uranium-thorium
29 ages from 19 speleothems collected in five caves at different elevations. After a period of very low
30 growth between 160 and 132 ka, the analysed speleothems grew almost continuously to ~75 ka, this
31 period was followed by intermittent growth with lower deposition rate and presence of hiatuses
32 until ~12.5 /12 ka. This is consistent with an ice coverage persisting over the Apuan Alps, inhibiting
33 or interrupting the growth of speleothems via the limited availability of groundwater and the
34 scarcity/absence of soils. This interval is much greater than the time interval that has previously
35 been attributed to the existence of glaciers on the Apuan Alps, which has been assumed to be
36 restricted to Marine Isotope Stage (MIS) 2. Instead, ice cover probably also appeared in the Apuan
37 Alps during MIS 4. The phase of restarting of growth, which may implies the definitive or
38 substantial glacier melts seem to predate the Holocene.

39 **Keywords:** Glacier; Pleniglacial; MIS2; MIS3; MIS4; Italy

40

41 **1. Introduction**

42

43 Speleothems (i.e. cave carbonate deposits) are multi-proxy paleoclimate archives, which can be
44 accurately dated back to several hundred thousand years before present thanks to U-Th isotope
45 systematic (Edwards et al. 1987; Hellstrom, 2006). Most of the paleoclimate research conducted
46 using speleothems focused on the interpretation of stable isotope and trace element proxy records
47 (e.g. Dykoski et al., 2005; Bar-Matthews et al., 1996; Wang et al., 2008; Regattieri et al., 2016).
48 However, speleothems need certain conditions to grow continuously, such as the presence of liquid
49 water, CO₂ and high Ca concentrations in drip waters (e.g. Atkinson et al., 1986; Gascoyne 1992;
50 Genty et al., 2001). This generally implies relatively wet climate conditions, surface temperature
51 above 0 °C and a well-developed soil above the cave (Gordon et al., 1989; Genty et al., 2001,2005).
52 For instance, in desert environments, phases of speleothem growth are important indicators of past

53 humid periods (Burns et al., 2001; Fleitmann et al., 2011; Vaks et al., 2006, 2007, 2010), as well as
54 tufa, which occurrence significantly correlates with wetter phases (Smith et al., 2004; Cremaschi et
55 al., 2015a). Similarly, the growth of speleothems in permafrost or glacier-dominated environments,
56 often marks interruptions of cold conditions (i.e. Berstad et al., 2002; Spötl and Mangini, 2006;
57 Pons-Branchu et al., 2010; Vaks et al., 2013). This is potentially consistent with degradation of
58 permafrost and the formation of large ice-free surfaces, favoring water infiltration and soil
59 formation (Varks et al., 2013). Growth hiatuses within a single or multiple speleothems have been
60 interpreted as evidences of particularly dry and/or cold conditions (e.g. Hodges et al., 2008; Genty
61 et al., 2003, Baldini, 2002; Moreno et al., 2010; Mayer et al. 2012, Stoll et al., 2013), even if the
62 nature and duration of hiatuses in speleothems cannot be unambiguously correlated to climatic
63 conditions. However, stacking multiple speleothem records can give more robust information about
64 the relationship between growth cessations and climate (e.g. Stoll et al., 2013).

65 During glacial periods, depressed external air temperatures at high latitudes and/or high altitudes
66 sites may cause extensive freezing at the surface above a cave, strongly limiting water infiltration
67 (Berstad et al., 2002; Genty et al., 2003,2005; Ayalon et al., 2013) and totally inhibiting soil activity
68 (McDermott, 2004). This implies a strong reduction in the CO₂ transfer to the epikarst, a
69 prerequisite for bedrock dissolution and speleothem growth. Therefore, development of glacial
70 conditions (i.e. the presence of glaciers) in the water-infiltrating area over a cave (catchment area
71 from here onwards) would be a factor for reduced or ceased speleothem growth. In such conditions,
72 the only possibility for the deposition of calcite can be related to sulphide oxidation if liquid water
73 is present (Atkinson 1983; Gascoyne and Nelson, 1983; Spötl and Mangini, 2007).

74 In the Mediterranean region there are numerous mountains where geomorphological and
75 geological evidence for the presence of glaciers has been dated (Baroni et al., 2018; Finsinger and
76 Ribolini 2001; Perez-Alberti et al., 2004; Federici et al., 2008; 2012; 2017; Kuhlemann et al., 2008;
77 Hughes et al., 2004, 2011, 2013; Giraudi et al., 2011 Giraudi and Giaccio, 2017; Ribolini al., 2011,
78 2018; Serrano et al. 2012; Hughes and Woodward, 2017 and references therein, Gromig et al.,

79 2018; Hannah et al., 2017; Akçar et al., 2017; Çiner et al., 2017; Sarikaya and Çiner, 2017).
80 However, for some regions, it is still challenging to identify the glacial cycles to which this
81 evidence belongs. This is particularly true for the Apuan Alps, a mid-latitude (44°N) mountain
82 chain in northwestern Italy where, despite the recognition of numerous glacial features, no precise
83 chronological data exist so far.

84 In this paper, we use periods of speleothem deposition as an indicator of ice-free conditions in
85 cave catchment areas, and periods of very low or absent stalagmite growth as proxy for constraining
86 the chronology of glacier presence in the Apuan Alps. To achieve this, we use previously published
87 and new U/Th ages of speleothems from five different caves located at different altitudes (Fig. 1).

88

89 **2. Site description**

90

91 The Apuan Alps is a NW-SE-oriented mountain range rising abruptly to about 2000 m a.s.l. from
92 the narrow coastal plain bordered by the Ligurian Sea (Fig. 1). The atmospheric circulation is
93 dominated by the important cyclogenesis centre of the Gulf of Genoa (Trigo et al., 2002) and by the
94 humid westerly air masses of North Atlantic provenance (e.g. Reale and Lionello, 2013; Fig. 1).
95 These sources of moisture impact on the Apuan Alps mountain chain, which acts as a natural barrier
96 by forcing an adiabatic rise of air masses, resulting in high precipitation (> 2,500 mm/yr Rapetti and
97 Vittorini, 1994; Piccini et al., 2008). As with most of the Apennine chain, winter rainfall amount,
98 which is the main period of recharge of Apuan caves (Piccini et al., 2008), is strongly regulated by
99 the North Atlantic Oscillation (NAO), with a negative correlations observed between NAO index
100 and winter precipitation (López-Moreno et al., 2011).

101 Two main tectono-metamorphic regional events have determined the structural setting of the
102 Apuan Alps (e.g. Carmignani and Kligfield, 1990; Molli and Vaselli, 2006), where extensive areas
103 of carbonate lithologies outcrop (Mesozoic marbles and metadolostones). The actions of glacial,

104 fluvial and karst processes led to an Alpine-like landscape, where exokarst landforms and
105 Quaternary cave systems are particularly developed (Piccini, 1998, 2003, 2011).

106 The Apuan Alps is one of the first places amongst the larger Apennine chain where glacial
107 landforms have been described (Cocchi, 1872; Stopani, 1872). Following these pioneering studies,
108 in the second half of nineteenth and throughout the twentieth century, further qualitative evidence
109 of glacial landforms were reported (De Stefani, 1874,1875,1890; Merciai, 1912; Paci, 1935;
110 Zaccagna, 1937; Tongiorgi and Trevisan, 1940; Beneo, 1945; Masini, 1970; Federici, 1981; 2005;
111 Braschi al., 1987; Putzolu, 1995, Baroni et al. 2013; 2015). Most of these glacial features are now
112 under risk of destruction due to marble quarrying (Bini, 2005). Despite only traces of small glaciers
113 have been reported on the seaward side of the mountain chain (i.e. erosional landforms and a few
114 scattered glacial deposits, Federici, 1981), features associated with at least nine glaciers have been
115 identified on the northern side. These include lateral and terminal moraines, cirques and polished
116 surfaces. Different phases of expansion have been described and, in some cases, the glacier termini
117 reached exceptionally low elevations (~600 a.s.l.) considering the latitude (Braschi et al., 1987;
118 Jaurand, 1998). The best-preserved glacial features, with relatively unweathered, incoherent detrital
119 materials, are traditionally attributed to the Last Glacial Maximum (LGM) by all previous authors
120 on the basis of their fresh appearance and by analogy with the Apennines and Alpine landforms,
121 The temporal definition of LGM itself is not univocal, depending on the considered references, e.g.
122 Clark at al. (2009) defined the LGM as the duration of sea-level lowstand during the interval 26.5-
123 20/19 ka, while Hughes and Gibbard (2015) as the event between the top (end) of Greenland
124 Interstadial 3 and the base (onset) of Greenland Interstadial 2, spanning the interval 27.540-23.340
125 ka (Greenland Stadial 3). In both cases within MIS2.

126 Inner terminal moraines have been interpreted as the Late-Glacial readvance/stillstand phases
127 (Braschi et al., 1987; Federici, 2005). Cemented glacial deposits, rarely outcropping, have been
128 attributed to a pre-LGM glacial event according to their appearance, altitude and stratigraphic
129 position (Braschi et al. 1987). The presence of these older, cemented deposits, attributed to “Riss

130 glaciation” (i.e. penultimate glaciation), has long been debated but no conclusive results have been
131 forthcoming due to the absence of chronological constraints (Federici, 2005), moreover, Kotarba et
132 al. (2001) have dated in the Apennines, similar types of older cemented moraines to the Riss, using
133 U-series.

134

135 **3. Cave descriptions**

136

137 In the Apuan Alps karst massif, about 2000 cave entrances are now mapped with a composite
138 length of cave development of ~500 km (<http://www.speleotoscana.it>). About 80 caves are carved
139 below the areas potentially covered by glaciers during the Last Glacial expansion, as reconstructed
140 by Braschi et al. (1987). Despite most of the Apuan caves are mainly vertical with active drainage
141 systems, many kilometres of passages are relict and nowadays host carbonate (usually calcite)
142 concretions. Here we focus on speleothems from five caves located at different elevations (Fig. 1).

143 The largest and highest cave is Antro del Corchia (CC), a large complex cave system (~60 km
144 long ~1200 m deep) mainly developed in the Upper Triassic-Liassic marble, dolomitic marble and
145 metadolostone of the Mt Corchia syncline confined by the non-karstifiable, low-permeable rocks of
146 the Paleozoic basement (Piccini et al., 2008).

147 Tana che Urla (TCU) is a small resurgence cave with a permanent stream (~600 m of total length
148 of which almost half is submerged; +45 m of total difference in height), developed at the contact
149 between Paleozoic schist basement and Triassic metadolostone (Regattieri et al., 2012). The
150 entrance is located at ~620 m a.s.l. on the south-eastern side of Panie Massif and functions as an
151 overflow spring during intense rainfall and snowmelt.

152 Grotta Del Vento (GDV) is an almost fossil phreatic cave (~4.5 km total length and over 450 m
153 of elevation change) developed at the contact between the Paleozoic schist basement and Triassic
154 metadolostone. The entrance is located at ~630 m a.s.l. on the south-eastern side of the Panie Massif
155 (Piccini et al., 2003).

156 Buca della Renella (RL) is a small, shallow cave at the confluence of Canale Regolo and the
157 Frigido River (Drysdale et al. 2006; Zhornyak et al., 2011). The entrance is located at ~275 m a.s.l.,
158 only a few metres above the stream confluence. The cave, predominantly horizontal (~200 m
159 length), is carved in Triassic metadolostone close to the contact with the Paleozoic phyllite.

160 Buca Cava dell'Onice (BCO) is a fossil small cave (85 m long and 65 m deep), carved in Lower
161 Jurassic cherty metalimestone. The entrance is located in the northwest side of Mt. Castagnolo at
162 ~700 m a.s.l.

163

164 **4. Methods**

165

166 We consider 19 speleothems (ages in supplementary material, Table 1), collected from the five
167 caves described above. Nine stalagmites and two cores drilled from flowstones have been collected
168 from the “Galleria delle Stalattiti” of the Corchia Cave. This part of the cave is a near-horizontal
169 chamber carved at ~840 m a.s.l., vertically overlain by ~400 m of rock. The chamber lies hundreds
170 of metres above the present groundwater table and is characterised by a mean annual air
171 temperature of 7.8 °C (Piccini et al., 2008). Five speleothems have already been described in
172 previous papers (CC1, CC5, CC7, CC26, CC27 and CC28, Drysdale et al., 2004, 2005, 2007, 2009;
173 Zanchetta et al., 2007; Bajo et al., 2017; Isola et al., 2018). We present new ages for stalagmites
174 CC4, CC7, CC22, CC27, CC53, and two flowstone cores CD4 and CD20.

175 At TCU two cores have been drilled from the same flowstone in the main gallery, about 100 m
176 from the entrance. The mean air temperature is 10.2 °C (Regattieri et al., 2014a). Speleothem
177 stratigraphy, chronology and geochemistry have been discussed by Regattieri et al. (2012, 2014a,
178 2016).

179 Two cores have been drilled from different flowstones located in the middle level of GDV
180 (Piccini et al., 2003).

181 At RL, three samples have been collected in different stratigraphic levels from a fan-like
182 flowstone deposited at the margin of an epiphreatic passage in the inner part of the cave (average air
183 temperature ca. 12 °C) and their stratigraphy and chronology have been discussed by Drysdale et al.
184 (2006) and Zhornyak et al. (2011).

185 Only one stalagmite (BCO1) has been collected in BCO, in a fossil gently dipping gallery, about
186 30 m from the entrance.

187 All ages from Corchia, Tana Che Urla and Renella have been obtained following the same
188 procedure at the University of Melbourne (Victoria, Australia). Briefly, uranium and thorium
189 isotopes were analysed using a Nu Instruments Multicollector Inductively Coupled Plasma–Mass
190 Spectrometer (MC-ICP-MS), according to the analytical method and corrections (where necessary)
191 described by Hellstrom (2003, 2006).

192 Age-depth models and speleothems growth rates were obtained from a Monte Carlo-derived
193 method (described in Drysdale et al. 2005 and Scholz et al. 2012). Dating methods for Grotta del
194 Vento are discussed by Piccini et al. (2003).

195 BCO1 $^{230}\text{Th}/^{234}\text{U}$ ages were determined at LSCE from 9 subsamples collected along the growth
196 axis of the stalagmite. Subsamples were crushed in an agate pestle and aliquots of typically ~400
197 mg were rinsed twice with MilliQ, dissolved in diluted HCl (~10%) and equilibrated with a mixed
198 ^{236}U - ^{233}U - ^{229}Th spike that was calibrated against a Harwell Uraninite solution (HU-1) assumed to
199 be at secular equilibrium. Uranium and thorium were purified and separated using Eichrom
200 UTEVA® and pre-filter resins in nitric media following a procedure modified from Pons-Branchu
201 et al. (2005) and Douville et al. (2010). The isotopes of uranium and thorium were analysed by
202 solution multi-collector ICPMS using a ThermoScientific Neptune^{Plus} hosted at the Laboratoire des
203 Sciences du Climat et de l'Environnement (LSCE, Gif-sur-Yvette, France) following the procedure
204 described by Pons-Branchu et al. (2014). Based on the measured atomic ratios, the $^{230}\text{Th}/^{234}\text{U}$ ages
205 were calculated through iterative age estimation (Ludwig and Titterton, 1994) using the ^{230}Th ,
206 ^{234}U and ^{238}U decay constants of Cheng et al. (2013) and Jaffey et al. (1971). An initial $^{230}\text{Th}/^{232}\text{Th}$

207 activity ratio of $0.9 \pm 50\%$ was used for correction of the non-radiogenic, detrital ^{230}Th fraction.

208 In order to highlight the higher and lower periods of speleothems deposition rate, we calculate
209 the growth rate for all the considered speleothems from Corchia Cave and for speleothems from
210 lower altitude caves, cumulated at 1000-yr intervals. Uncertainties in the growth rates were
211 obtained using the standard errors propagation.

212 It is worth estimating potential changes in air temperature at the altitude of the different
213 infiltration areas of the four caves (we ignore BCO because the sole sampled stalagmite covers only
214 MIS 5), in order to understand causes of any leads and lags in speleothem growth phases in the
215 context of likely glacier development. The following calculations represent a first-order estimate
216 that can be used to underpin the interpretations proposed below.

217 Speleothem growth patterns described above were compared with sea-surface temperature (SST)
218 series from three Mediterranean cores: ODP 161-977 (Martrat et al., 2014) spanning all over the
219 considered period; PRGL1 (Cortina et al., 2015), the nearest core to the Apuan Alps; and MD
220 90917 (Siani et al 2013), a high-resolution dataset covering the last ~25 kyr (cores locations are
221 shown in Fig. 1). As these cores (ODP 161-977, PRGL1) provide SST data based on Alkenone
222 unsaturation indices (Müller et al., 1998; Conte et al., 2006) and planktonic foraminifera
223 assemblages (Siani et al 2013), they do not reflect exactly the same temperature signal. SST records
224 in the Alboran Sea (ODP191-977) are more related to spring-summer conditions (Martrat et al.,
225 2004, 2014), while the SST records from the Gulf of Lyon reflects a winter-spring temperature
226 signal showing a colder imprint than other more southern SST records (Rigual-Hernández et al.,
227 2013, Cortina et al., 2015). Finally, planktonic foraminifera assemblages coincide with the most
228 productive period during the spring and the synchronous bloom of *G. bulloides* in the
229 Mediterranean Sea (Pujol and Vergnaud-Grazzini, 1995; Siani et al., 2013). In performing the
230 ocean-cave comparisons, we make several assumptions. First, we assume that the difference in
231 mean annual SST (MASST) between the selected marine cores and the MASST offshore of the

232 Apuan Alps (Locarnini et al., 2010) was constant through time; second, that the difference between
233 the MASST and the Mean Annual Air Temperature (MAAT) measured along the coast facing the
234 Apuan Alps (<http://www.autorita.bacinoserchio.it/>) was constant through time; third, we adopt an
235 air temperature lapse rate of 0.6 °C/100 m (as calculated for the area by Rapetti and Vittorini, 2012)
236 to estimate the MAAT in the speleothem catchment areas of four of the studied caves. For the
237 elevations of the water recharge area feeding the caves, two values are considered: the elevation of
238 the surface above the cave on the vertical of the sampling sites (C1, supplementary material Table
239 2) and the average between this point and the highest point of the drainage basin (C2,
240 supplementary material 2). The first is the most conservative value representing the minimum
241 elevation of recharge area while the second is probably more realistic because takes into account the
242 whole hydrographic basin upvalley the cave. A more refined estimate of MAAT could incorporate
243 also the effects of sea-level lowering during glaciations and tectonic uplift (estimates to be 0.6
244 mm/yr; Fellin et al. 2007; Piccini 2011) of the Apuan Alps. However, both effects introduce
245 negligible temperature variations. Indeed, the maximum increase in relative elevation of the
246 infiltration area above the contemporary sea level is during the LGM (e.g. ca. 100-120 m for the
247 LGM, Lambeck and Bard 2000; Lambeck et al., 2014), corresponding to a total temperature
248 variation (ΔT) of about -0.6°C, while the lowest, during MIS6, corresponds to ΔT of about +0.1°C.

249

250 **5. Results**

251 For the 19 speleothems considered, the dataset comprises 293 U/Th ages (supplementary
252 material Table 1). In Figure 2, the ages are plotted vs the elevation of the sampling sites.

253 Despite some differences, there are many common features. All caves show a major episodes of
254 calcite deposition during the Holocene and the latest Late Glacial. Available samples from RL and
255 GDV do not show phases of growth older than this period and the studied concretions cover, or are
256 stratigraphically located over, thick fluvial deposits. TCU shows continuous deposition between ca.

257 160 and 121 ka (Regattieri et al., 2014a), followed by a discontinuous deposition to ca. 75 ka, after
258 which there is a long hiatus from ca. 75 to 12.5 ka. The most relevant record is represented by
259 Corchia Cave, showing almost continuous calcite deposition from ca. 145 to 79 ka. A long lag of
260 growth, with only very reduced phases of calcite deposition is evident during the Pleniglacial. The
261 stalagmite from BCO has an intermittent growth phase throughout MIS 5, confirming a
262 speleothems deposition phase during this period in the Apuan Alps. No information about other
263 periods including the Holocene are available for this site until now, but we cannot exclude further
264 growing phases, in fact, a single speleothem cannot be considered representative of the whole
265 speleothems growth history of this cave.

266 More interesting than the simple ages distribution is the cumulative growth curve which takes
267 into account the propagation of age errors for Corchia cave speleothems and for the other studied
268 caves located at lower altitudes (Fig. 3). It is evident that speleothem growth at Corchia, after a
269 period of low values, becomes significant between ca. 132 and 80 ka (curve b in figure 3), after
270 which it decreases abruptly until ca. 60 ka, and then almost stops for ca. 44 kyr. Continuous growth
271 resumes at ca. 15 ka and increases later at ca. 12 ka.

272 The growth of speleothems belonging to the lower-altitude caves evidently decreases after 130
273 ka, reaching a local minimum at ca. 100 ka. An increase of composite growth rate is then observed
274 around 80 ka, followed by an interruption in calcite deposition up to 13 ka.

275 Figure 4 shows the estimation of MAAT as described above. Due to the very similar geographic
276 location (fig. 1), GDV and TCU temperature estimates are almost indistinguishable in all the
277 reconstructions (blue and black lines in Fig. 4). The MAAT curves obtained for each cave using
278 MD90917 data (~ 0.6 - 24 ka interval), including the two different infiltrating elevations (C1 and
279 C2), are shown respectively in Figure 4a and b. The MAAT curves obtained using ODP 161-977
280 data (~ 0.06 - 120 ka interval) are shown respectively in figure 4c and d. The MAAT curves
281 obtained using PRGL1 data (~ 29 - 120 ka interval) are shown respectively in Figure 4e and f. In
282 the C1 cases (Fig. 4a, c, e) only the CC series displays excursions where MAAT is $<0^{\circ}\text{C}$. In the C2

283 scenario (Fig. 4b, d, f), GDV and TCU areas also show periods when MAAT is $<0^{\circ}\text{C}$. As expected,
284 the temperature above RL is the mildest.

285

286 **6. Discussion**

287

288 According to the presented data, there is a long interval during the Pleniglacial where calcite
289 deposition at CC and TCU caves slows down or is absent (fig. 3). This interval roughly corresponds
290 to an extensive loess deposition in the Po Plain (Fig. 1 and 5; Cremaschi et al., 2015b; Zarboni et
291 al., 2015). This marks a wide phase of aridity in northern Italy that might have involved also the
292 Apuan Alps, but probably with different effects. Nowadays the mean annual precipitation in Po
293 Plain does not exceed 1000 mm (www.arpalombardia.it), while the Apuan Alps are characterized
294 by higher precipitations exceeding 2500mm/yr (Rapetti and Vittorini, 1994; Piccini et al., 2008).
295 The different geographic position underpinning the actual rainfall regime, certainly played a role in
296 the past too, mitigating the drought effect in the Apuan Alps. Moreover, it is very difficult to
297 explain in a dry regime, the strong fluvial activity suggested, in the low elevation caves, by the
298 presence of thick fluvial deposits, stratigraphically older than the phase of speleothems regrowth
299 (about 12.5 ka). The resumption of calcite deposition is almost coeval in all caves, with evidence of
300 sustained growth slightly preceding the beginning of the Holocene, instead, the period of deposition
301 decrease, at the end of MIS5, is not coeval between TCU and CC. However, the CC dataset is more
302 extensive because the number of analyzed speleothems.

303 Moreover, this cave is very close to areas of the catchment occupied by glaciers in the Last
304 Glacial (Fig.1). Specifically, the recharge basin of the “Galleria delle Stalattiti”, mainly consisting
305 of the higher slopes of the northern side of Monte Corchia, was covered by a glacier according to
306 Braschi et al. (1987). The most obvious process responsible for the strong reduction of speleothem
307 growth at the “Galleria delle Stalattiti” would thus be a persistent ice cover. This long interval of
308 extremely reduced or absent calcite deposition covers most of MIS4-MIS2, a time span more

309 extensive than the classical attribution of Apuan glacial activity to the only Last Glacial Maximum
310 (Federici, 2005; Jaurand 1998). This hypothesis is supported by evidence of moraines formed
311 during MIS3 and MIS4 glacier expansions in the Central Apennines (Giraudi et al. 2011; Giraudi
312 and Giaccio, 2015, 2017). In this regard, the Corchia speleothems cumulative growth curve delimits
313 for the first time the temporal interval of possible glacier existence over the Apuan Alps or
314 temperature lower enough to inhibit speleothem growth at higher altitudes. Although the
315 chronological constraints are indirect because they do not refer to moraine ages, there is
316 considerable robustness in the speleothem-age dataset. However, cessation of speleothem growth
317 can be related to combined effect of low air temperature, soil-CO₂ deficiency and reduction of
318 dripping. These conditions might have been present without invoking continuous ice covering of the
319 slopes, i.e. where permafrost inhibits water infiltration and following dryer cave conditions. In the
320 nearby Apennine chain, landforms diagnostic of permanently frozen ground are rare and essentially
321 limited to the southern latitudes, where higher elevations are reached (Kotarba et al., 2001; Giraudi
322 et al., 2011; Oliva et al., 2018).

323 The absence of speleothem growth in the other caves during the glacial period cannot be directly
324 related to the reconstructed ice cover (Braschi et al. 1987) because of their lower altitudes. A recent
325 geomorphology synthesis of the Apuan Alps (Baroni et al., 2015) does not report glacial deposits
326 indicative of a glacial cover of the lower caves.

327 At RL and GDV, the presence of thick fluvial deposits stratigraphically older than the phase of
328 speleothem growth (about 12.5 ka) suggests strong fluvial activity during the glacial period, which
329 would inhibit speleothem growth (Zhornyak et al., 2011). This strong activity might be related to
330 seasonal snow/ice melting and abundant availability of loose material coming from physical
331 weathering. A similar explanation is also valid for the long deposition hiatus at TCU (about 12.6-75
332 ka). A continuous glacier cover nearby is unlikely throughout this period.

333 It is noteworthy that both CC and TCU show speleothem deposition during part of MIS6,
334 suggesting that temperatures over the Apuan Alps during this period were not as low as those for

335 MIS2-4. The transition between the penultimate glacial and the last interglacial, is marked by the
336 abrupt increase of speleothems growth about 132 ka, as already described for CC by Drysdale et al.,
337 (2009) and for TCU by Regattieri et al. (2014a) and in agreement with other proxies like the more
338 umid-temperate vegetation in Southern Europe (Sanchez –Goñi et al., 1999), or the high level in
339 Ioannina Lake (Wilson et al., 2015). Then, speleothem growth, which was possible for most of
340 MIS5, decreases or ceases during the interval MIS2-4 (Fig. 3).

341 The almost coeval resumption of calcite deposition in all the investigated caves requires further
342 discussion. At CC, some U/Th ages indicate that a small growth phase also occurs at the beginning
343 of the Late Glacial, but significant growth did not occur before ca. 12 ka. However, given the
344 southern latitude and the relatively low altitude of the Apuan Alps, it is hard to argue that glaciers
345 still existed during the middle part of the Lateglacial (Hughes and Gibbard, 2015). Based on
346 deglacial trends in CC speleothem carbon isotope composition, Drysdale et al. (2004) and Zanchetta
347 et al. (2007) have suggested that soil development at CC is delayed after glacial periods due to
348 intense erosive processes active at high altitudes during the first phases of climatic amelioration,
349 under conditions of increasing rainfall. After the definitive ice melting at the end of MIS2, cold but
350 ice-free conditions and delayed soil development would have retarded significant speleothem
351 growth, which did not accelerate until ca. 12 ka. The fundamental elements of this hypothesis are
352 supported by Bajo et al. (2017), who also invoked near-closed-system conditions and sulphuric-acid
353 dissolution as additional processes to carbonate-acid dissolution (from soil-sourced CO₂) to explain
354 the unusually high carbon isotope composition of CC26 stalagmite during late Late Glacial and the
355 beginning of Holocene. However, an implication of this study is that growth was unsuitable before
356 ca. 12 ka, in spite of the possibility of sulphuric-acid dissolution substituting for carbonic-acid
357 dissolution, thus suggesting that speleothem growth was inhibited by cold climate conditions. From
358 the same stalagmite, Regattieri et al. (2014b) concluded, based on the presence of an aragonite layer
359 at the base of the CC26 succession and on trace element/Ca ratios (Ba, Sr, Mg) changes, that the

360 start of growth of CC26 was related to a re-opening of a dripping point due to the replenishing of
361 the cave plumbing system previously interrupted by reduced recharge.

362 Whilst we cannot detail the point at which resumption of speleothem growth occurred in CC
363 during the Late Glacial, conditions generally favoring speleothem growth seem to have occurred
364 when the plumbing system was newly recharged significantly even if ice cover should have already
365 disappeared. It is probably not obvious that these conditions did not concomitantly occur in upper
366 and lower altitude caves. Studying the transition between MIS6 and MIS5 at TCU Regattieri et al.
367 (2014a) suggested that the delay in soil development as inferred for CC is not supported for TCU,
368 due to its lower altitude. This may suggest that an earlier resumption of speleothem growth would
369 be anticipated at lower altitude compared to CC. Data collected until now agree with this hypothesis
370 (Fig. 3). We suggest that significant resumption of calcite deposition at CC can occur only when
371 soil development and drip recharge are restored to a certain level.

372

373

374 According to the timing of speleothem growth and MAAT estimations, it can be seen that above CC
375 both the C1 and C2 MAAT reconstructions experienced a number of long-lasting sub-zero periods.
376 Indeed, the Equilibrium Line Altitude (ELA) of 1200-1300 m a.s.l. in the Apuan Alps during the
377 MIS 2 (Merciai, 1912; Braschi et al., 1987; Jaurand, 1998) supports the hypothesis that in the CC
378 catchments (both C1 1240 m a.s.l. and C2 1460m a.s.l.) glaciers were existing. These unfavorable
379 conditions for speleothem deposition (MAAT subzero or near 0 °C) persisted, discontinuously,
380 above CC not only during the MIS 2, but also during MIS 4 (Fig. 4). This is consistent with
381 evidence of a phase of glaciation during MIS 4 from the Alps and other peri-Mediterranean
382 mountain ranges (Hughes et al. 2013; 2018; Sarikya et al., 2014; Pope et al., 2017) (blue bars in
383 Fig. 5a). There are evidences of very small growth phases during MIS4-MIS2 (Fig. 2, 3, 5d) that
384 may have occurred during warming reversals, most likely associated with the millennia-scale
385 interstadials observed in the Greenland ice-core records (i.e. Dansgaard-Oeschger events,

386 Dansgaard et al., 1993; Fig. 5e), but also indicated by SST changes in the western Mediterranean
387 (Cacho et al., 1999).

388 In general, pollen data from the western Mediterranean show that cold SSTs characteristic of
389 Greenland stadials were contemporaneous with the expansion of semi-desert or steppic vegetation.
390 On the contrary, Greenland interstadials were synchronous with the expansion of open forests
391 (Sánchez-Goñi et al., 2008). Presumably, phases of forest expansion prompted the development of
392 deeper soils. However, the very limited speleothem growth and the available ages prevent a precise
393 correlation with the warmer intervals. On the other hand, according to Atkinson (1983) and Spötl et
394 al. (2007), speleothems may form, although with a lower rate of growth, in caves overlain by
395 temperate glaciers with available glacial meltwater infiltrating the karst system. The very low
396 partial pressure of CO₂ in such water due to the absence of soil can be balanced by pyrite oxidation
397 during the flow paths through fractures, providing the dissolved carbonate necessary for speleothem
398 deposition. This process is demonstrably present at Corchia (Piccini et al., 2008; Bajo et al., 2017)
399 and can be responsible of the very thin deposition layers included between two hiatus.

400 Summing up, growth histories and temperature estimation support the hypothesis that
401 interruptions or strong reductions in speleothem growth at Corchia could have been due to a
402 combination of glaciation and/or ice-free but cold temperatures with low or absent soil activity.

403 Glacial phases occurring in this period in the mountain regions around the Mediterranean are
404 confirmed by multi-proxy analyses carried out on the glacial lacustrine sequences, as evidenced by
405 a maximum of glacier expansion of Pyrenean glaciers before 30 ka (Garcia-Riuz et al., 2003) and in
406 Picos de Europa dated at about 40 ka (Moreno et al. 2010; Serrano et al., 2012). Consistently in the
407 central Apennines, tephra layers in lacustrine sediments bracket the maximum glacier extent
408 between 33-27 ka, not a period of maximum cold conditions but one that was humid (Giraudi,
409 2012). Regarding the Alps there is a controversial discussion about glaciers extent before MIS 2
410 when the maximum expansion was reached (Ivy-Ochs et al. 2008; Preusser et al. 2011; Hughes et
411 al. 2013). The uncertainties mainly derive from the fact that the majority of terrestrial records (i.e.

412 glacial deposit, pro-glacial outwash, river terrace, lake deposit) were eroded, or at least remoulded
413 or reworked, when MIS 2 glaciers invaded the valleys and in many cases spread out in piedmont
414 lowlands. Indeed, while in the Eastern Alps no evidence for the presence of pre-MIS 2 glaciers are
415 clearly found (Ivy-Ochs et al., 2008), an important glaciation reaching the lowlands of the Western
416 Alps during MIS 4 is documented by OSL dating of pro-glacial outwash (Preusser et al. 2007) and
417 lake sediments (Link and Preusser 2006).

418 The above described asynchronicity is not contradictory a priori, because it may be explained by
419 the short reaction times of small glaciers (like those of most of the Mediterranean mountains) to
420 changes in precipitation and/or temperature. Further sampling may reveal sporadic or very slow
421 episodes of growth during this time interval, allowing better and more precise interpretations of
422 MIS 3 climatic oscillations.

423

424

425 **7. Conclusions**

426

427 Speleothems in the Apuan Alps grew near-continuously during most of MIS5 (from ca. 130 ka to
428 about 75 ka), then experienced a long period of absent to very low deposition covering the MIS4 to
429 MIS2 interval, until ~12.5 ka. This growth history can be explained as the result of development of
430 glaciers at high elevations, not necessarily continuously existent for the whole considered period,
431 and/or low soil development and near-zero temperatures. The speleothem growth curves may
432 suggest that glacial expansion in the Apuan Alps was not only limited to the LGM (MIS2) but also
433 to MIS3 and MIS4, analogous to other peri-Mediterranean mountains.

434 During spring and summer ice-melting periods, high water discharges occurred and caves
435 located at lower altitudes were flooded, resulting in phreatic conditions, inhibiting speleothem
436 deposition and infilling caves with thick fluvial deposits.

437 There is no evidence of significant growth during interstadials within MIS3, which supports the
438 fact that no particularly warm conditions occurred in the Apuan Alps during those climatic events.

439 Sustained speleothem growth resumed close to the beginning of the Holocene both at lower and
440 high-altitude caves. This can be interpreted as the result of a definitive melting of ice cover and the
441 consequent progressive soil development at high altitude, and the stabilisation of loose glacial
442 debris and reduced flooding for lower altitude caves.

443 These data indicate the potential for cumulative speleothems growth curve to chronologically
444 constrain large-scale geomorphological processes, like the timing of glaciers presence. However,
445 other processes can locally bias the speleothem growth (i.e. cold-non glacial conditions,
446 absence/limited soils), suggesting that this approach is restricted to a first-order approximation
447 which should be followed by a direct dating of glacial features. Moreover, the adoption of
448 speleothem growth curves needs to be verified in other mountain ranges, where U/Th dating can be
449 used for constraining glacial events.

450

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452

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464

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869

870 **Captions**

871 Figure 1. Glacier extent (light blue) during the LGM according to the reconstruction of Braschi et
872 al. (2001). Sampled sites are shown in yellow.

873 Figure 2. ²⁹³U/Th ages from 19 Apuan speleothems over the past 160 ka (2 σ error bars) vs the
874 elevation of sampling sites. Details on dating can be found in the original papers (Drysdale et al.,
875 2004; 2009; Regattieri et al., 2012, 2014a, 2014b; Piccini et al, 2003; Drysdale et al., 2006;
876 Zhornyak et al 2011, Isola et al., 2018).

877

878 Figure 3. Cumulative growth curves (i.e. the total speleothem growth calculated for all the
879 considered speleothems and binned into 1000-yr intervals) for Corchia Cave speleothems (red lines
880 a and b) and for the caves located at lower altitude (blue lines, c and d), the associated errors in
881 shaded strips. Note the linear (b and d) and the logarithmic (a and c) scales in the vertical axes.
882 Shown in black (e) is the LR04 stack (Lisiecki and Raymo 2005) from benthic foraminifera $\delta^{18}\text{O}$
883 data. Numbers indicate Marine Isotope Substages.

884

885 Figure 4. MAAT curves derived for all the caves from MD90917: (a) C1 (see the text for
886 explanation), (b) C2; from ODP 161-977, (c) C1 and (d) C2 ; and from PRGL1, (e) C1 and (f) C2.
887 Black horizontal lines indicate the 0°C MAAT.

888

889 Figure 5. Northern Hemisphere proxies for the Pleniglaciation: a) the approximate timing of
890 maximum extent of glaciers in the Mediterranean mountains (after Hughes et al., 2013; 2018;
891 Sarikya et al., 2014; Pope et al., 2017); b) the derived MAAT curves in the C2 case above the
892 Galleria delle Stalattiti: red curve is derived from ODP 161-977, purple curve from MD90917 data,
893 black curve is derived from PRGL1 data. Horizontal black line indicates the 0°C MAAT; c) the
894 derived MAAT curves in the C1 case above the Galleria delle Stalattiti: red curve is derived from
895 ODP 161-977, purple curve from MD90917 data, black curve is derived from PRGL1 data and
896 horizontal black line indicates the 0°C MAAT; d) Cumulative growth curve calculated for interval
897 of 1000 yr for Corchia cave speleothems (green line); e) NGRIP $\delta^{18}\text{O}$ values from North Greenland
898 Ice Core (North Greenland Ice Core Project members, 2004). Numbers over the graph mark the
899 position of the Dansgaard-Oeschger interstadials (Dansgaard et al., 1993); f) approximate interval
900 of extensive loess accumulation in Po Plain (after Cremaschi et al., 2015b; Zerboni et al., 2015).

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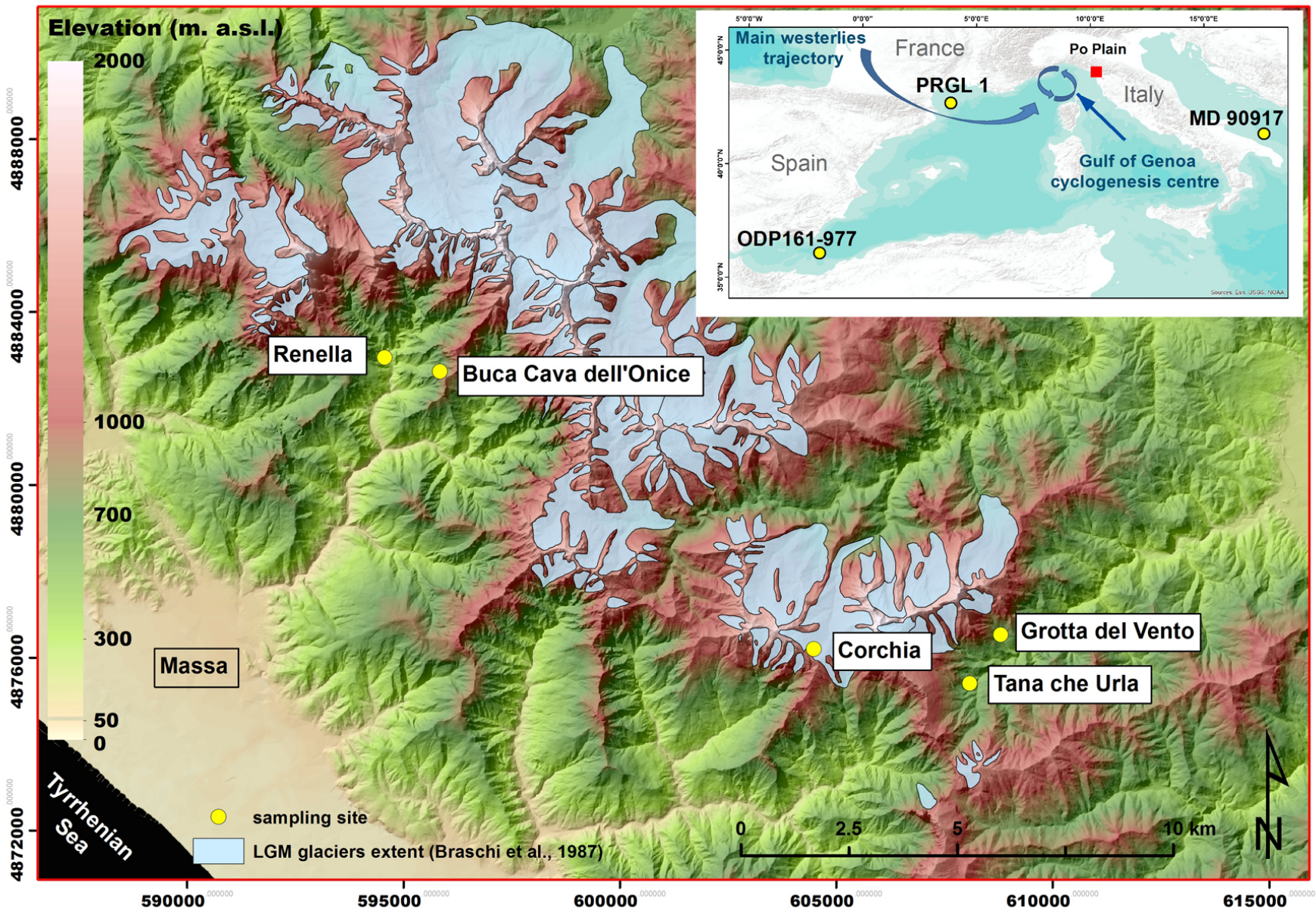
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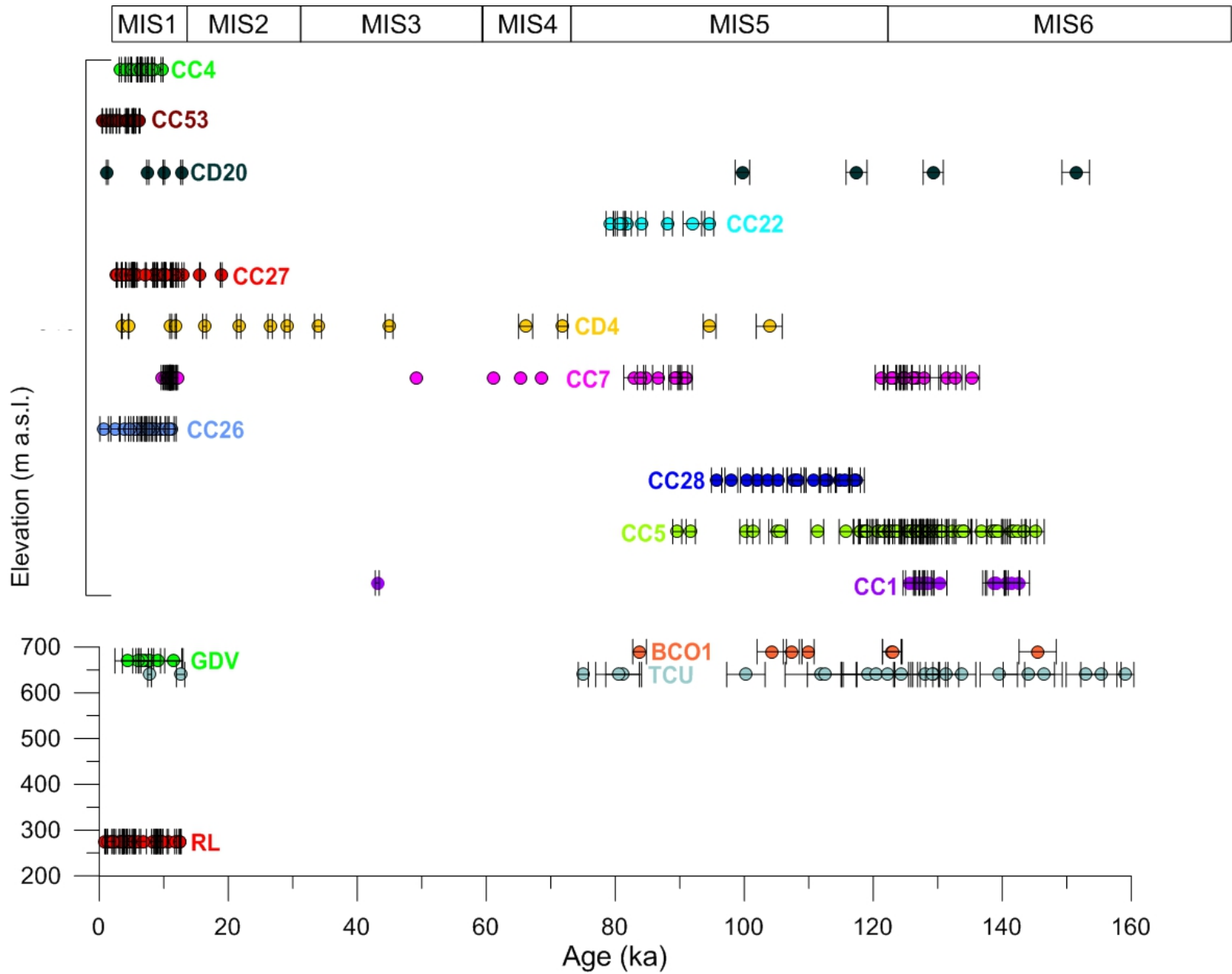
903 **List of tables (supplementary materials)**

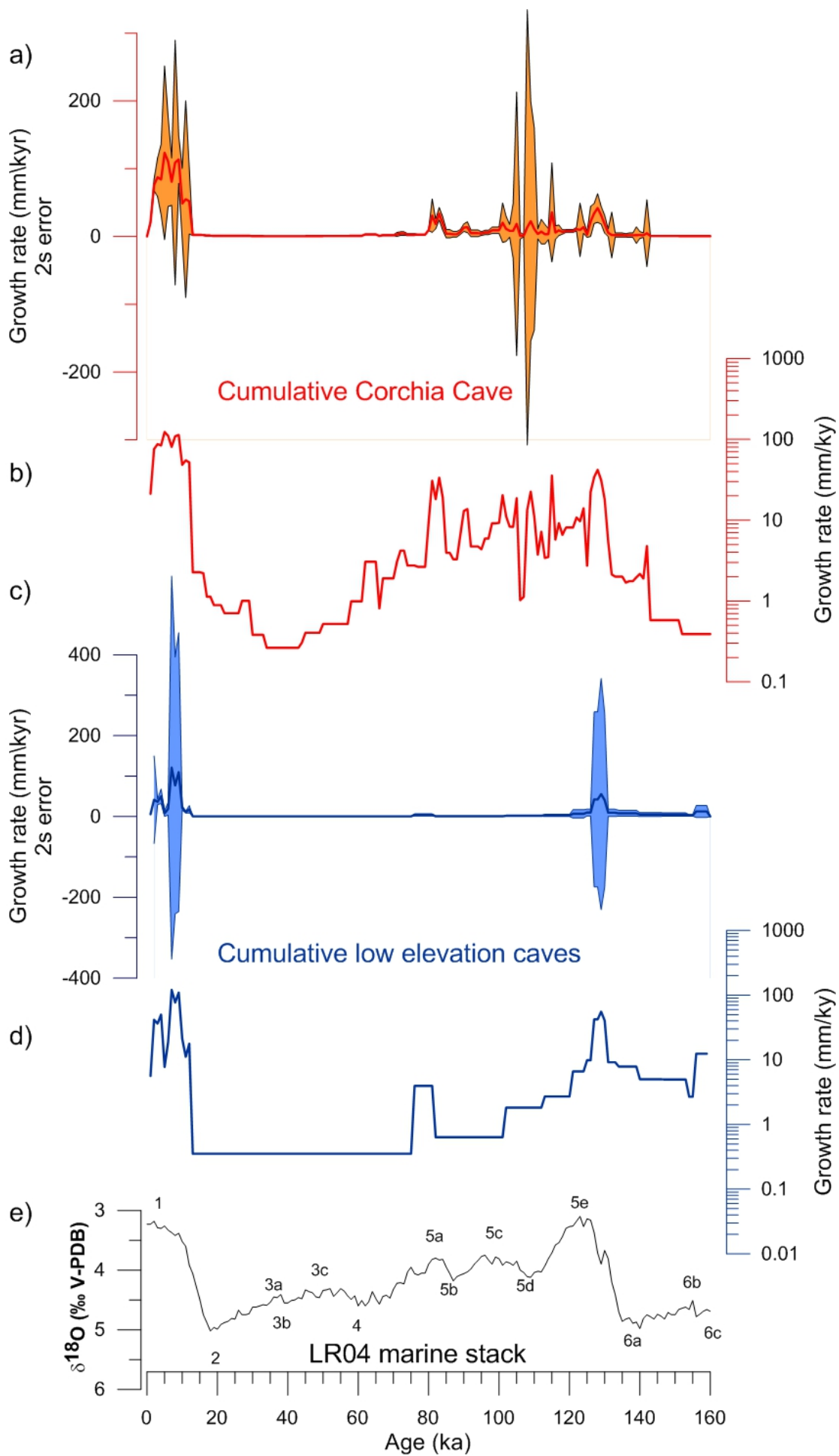
904 Tab. 1 U/Th data used for the age distribution

905 Tab. 2 Elevation used to to estimate the MAAT in the speleothem catchment areas (C1, C2)

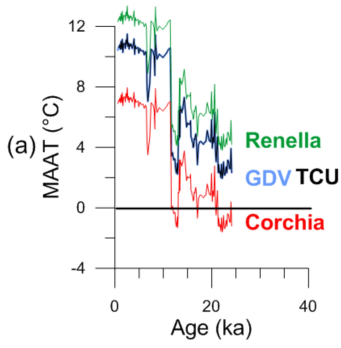
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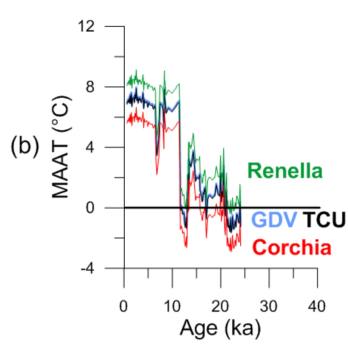




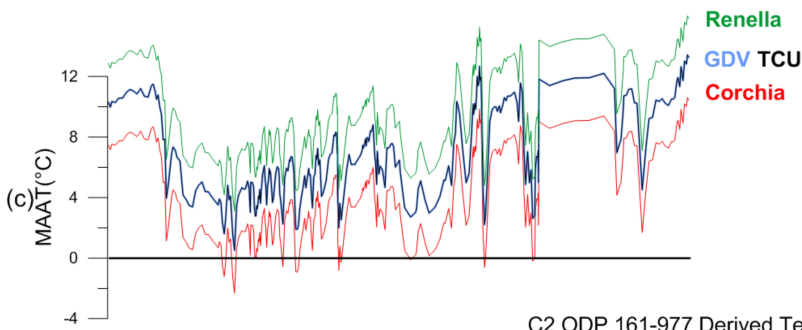
C1 MD90917 Derived Temperature



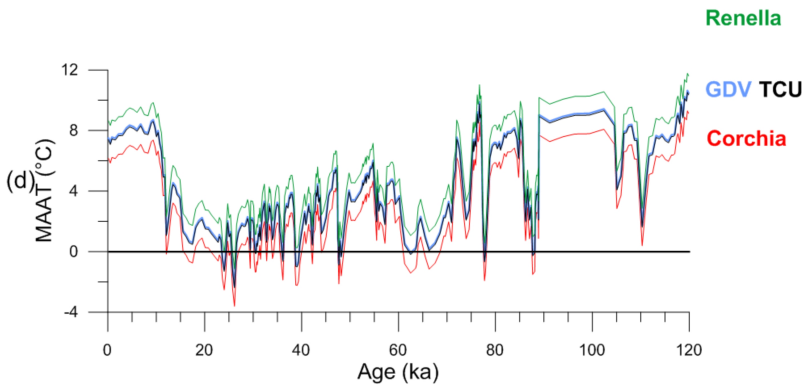
C2 MD90917 Derived Temperature



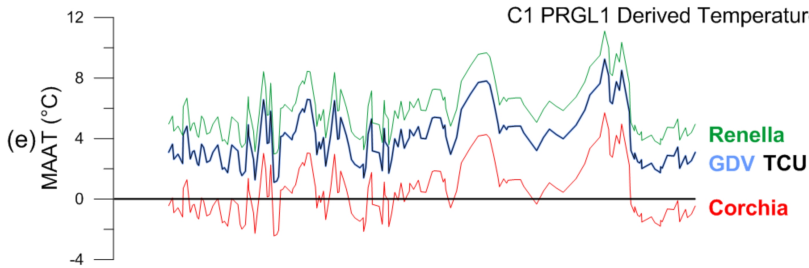
C1 ODP 161-977 Derived Temperature



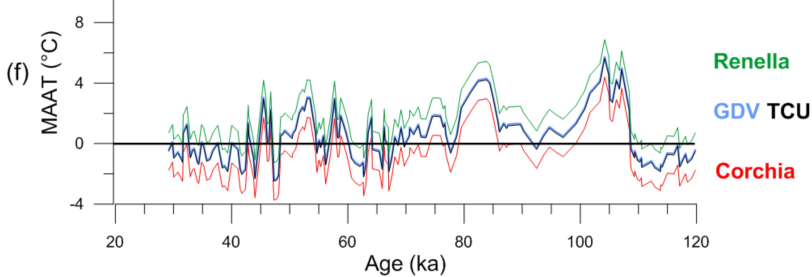
C2 ODP 161-977 Derived Temperature



C1 PRGL1 Derived Temperature



C2 PRGL1 Derived Temperature



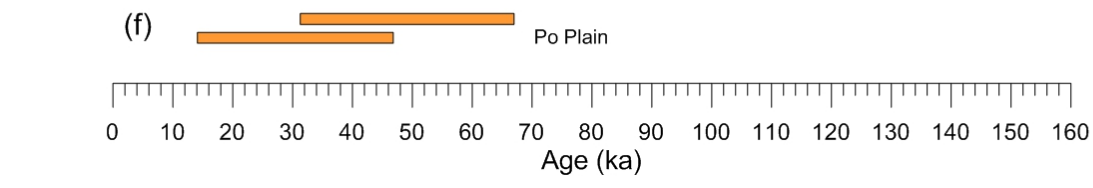
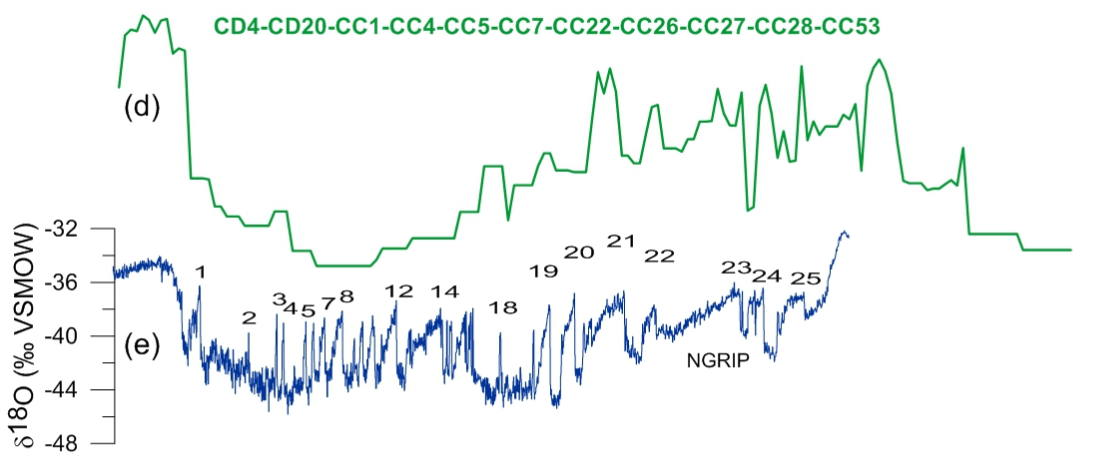
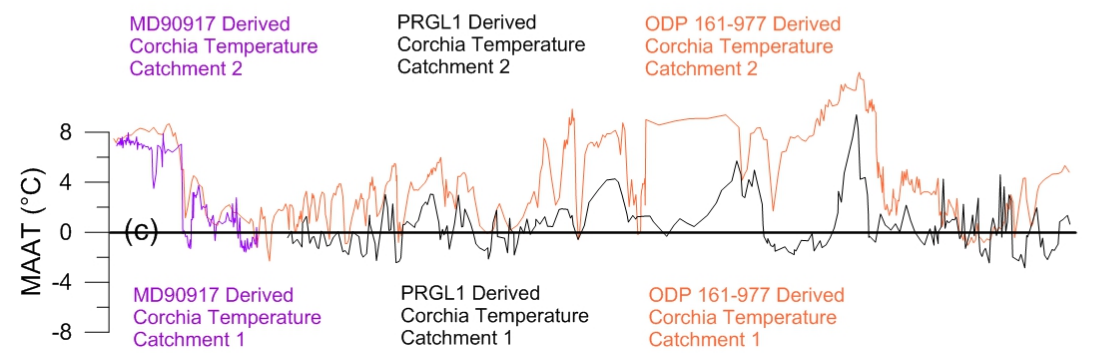
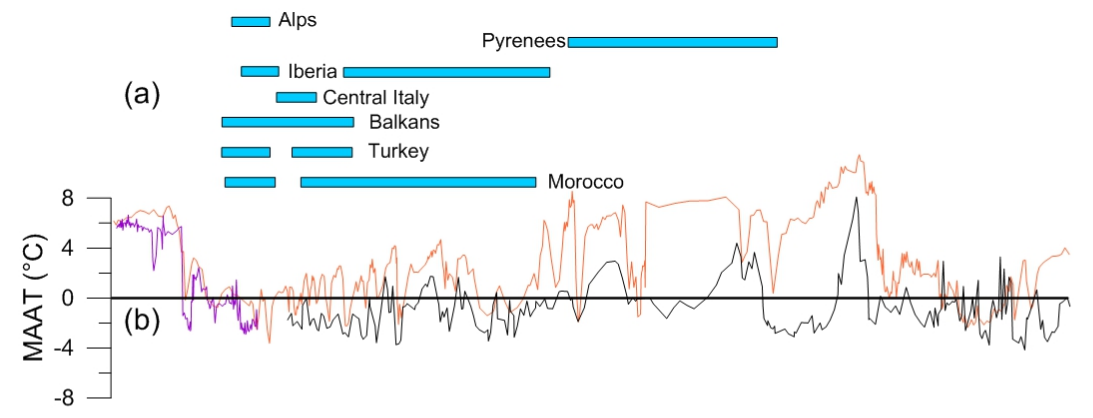


Table 1 U/Th data used for the age distribution

Cave	Sample	Age (ka)	2 σ	Sample site elevation (m a.s.l.)	Sampling depth (mm from top)	References
Corchia	CC110	139	1.4	840	40	drysdale et al., 2004
Corchia	Cor4	126.4	1.4	840	13	drysdale et al., 2004
Corchia	CC117	130.2	1.2	840	35	drysdale et al., 2004
Corchia	CC116	141.5	1.1	840	43	drysdale et al., 2004
Corchia	CC119	125.6	1	840	11	drysdale et al., 2004
Corchia	CC118	127.2	0.8	840	23	drysdale et al., 2004
Corchia	Cor12	126.9	0.7	840	19	drysdale et al., 2004
Corchia	Cor11	127.8	0.6	840	23	drysdale et al., 2004
Corchia	Cor3	43.1	0.3	840	8	drysdale et al., 2004
Corchia	CC5-133	139.4	1.8	850	247	drysdale et al., 2005
Corchia	CC5-5	122.4	1.7	850	129	drysdale et al., 2005
Corchia	CC5-4	120.7	1.6	850	118	drysdale et al., 2005
Corchia	CC5-13	129.9	1.5	850	239	drysdale et al., 2005
Corchia	CC5-127	145.1	1.4	850	253	drysdale et al., 2005
Corchia	CC5-4R	120.7	1.4	850	118	drysdale et al., 2005
Corchia	CC5-136	138.9	1.3	850	244	drysdale et al., 2005
Corchia	CC5-2	105.1	1.3	850	60	drysdale et al., 2005
Corchia	CC5-328	126.6	1.3	850	144	drysdale et al., 2005
Corchia	CC5-6	128.7	1.3	850	169	drysdale et al., 2005
Corchia	CC5-3	105.5	1.2	850	63	drysdale et al., 2005
Corchia	CC5-7	134	1.2	850	242	drysdale et al., 2005
Corchia	CC5-271	115.8	1.1	850	84	drysdale et al., 2005
Corchia	CC5-32	129.5	1.1	850	231	drysdale et al., 2005
Corchia	CC5-227	100.3	1	850	36	drysdale et al., 2005
Corchia	CC5-239	101.4	1	850	50	drysdale et al., 2005
Corchia	CC5-253	111.3	1	850	67	drysdale et al., 2005
Corchia	CC5-288	118.8	1	850	103	drysdale et al., 2005
Corchia	CC5-30	128.5	1	850	234	drysdale et al., 2005
Corchia	CC5-31	128.6	0.9	850	228	drysdale et al., 2005
Corchia	CC5-333	127.6	0.9	850	148	drysdale et al., 2005
Corchia	CC5-12	89.6	0.7	850	16	drysdale et al., 2005
Corchia	CC5-217	91.7	0.7	850	24	drysdale et al., 2005
Corchia	CC28-829	114.7	1.7	860	76	drysdale et al., 2007
Corchia	CC28-597	108	1.3	860	53	drysdale et al., 2007
Corchia	CC28-1020	117.4	1.2	860	107	drysdale et al., 2007
Corchia	CC28-801-803	112.9	1.2	860	73	drysdale et al., 2007
Corchia	CC28-883	115.5	1.2	860	81	drysdale et al., 2007
Corchia	CC28-473	107.7	1.1	860	40	drysdale et al., 2007
Corchia	CC28-667	110.7	1.1	860	60	drysdale et al., 2007
Corchia	CC28-135	100.4	1	860	13	drysdale et al., 2007
Corchia	CC28-635	108.3	1	860	56	drysdale et al., 2007
Corchia	CC28-81	98	1	860	8	drysdale et al., 2007
Corchia	CC28-1144	117.1	0.9	860	94	drysdale et al., 2007
Corchia	CC28-334	103.6	0.9	860	33	drysdale et al., 2007
Corchia	CC28-732-734	112.5	0.9	860	66	drysdale et al., 2007
Corchia	CC28-2	95.7	0.8	860	1	drysdale et al., 2007
Corchia	CC28-439-440	105.2	0.8	860	37	drysdale et al., 2007
Corchia	CC28-281	102	0.7	860	28	drysdale et al., 2007

Corchia	CC1-632	147.6	3.5	840	39	drysdale et al., 2009
Corchia	CC5696+697	138.4	2.4	850	241	drysdale et al., 2009
Corchia	CC5-734	132.3	2.4	850	237	drysdale et al., 2009
Corchia	CC1-587	136.3	2.2	840	35	drysdale et al., 2009
Corchia	CC1-609	139.7	2.2	840	37	drysdale et al., 2009
Corchia	CC1615-616	140.6	2.1	840	38	drysdale et al., 2009
Corchia	CC5-653	143.3	2.1	850	245	drysdale et al., 2009
Corchia	CC5-766	130.5	2.1	850	234	drysdale et al., 2009
Corchia	CC7-325	126.8	2	900	140	drysdale et al., 2009
Corchia	CC1-570	130.3	1.9	840	33	drysdale et al., 2009
Corchia	CC5638-636	142.4	1.9	870	247	drysdale et al., 2009
Corchia	CC1-578	132.1	1.8	840	34	drysdale et al., 2009
Corchia	CC5-685	141.7	1.8	850	242	drysdale et al., 2009
Corchia	CC1-549	128.5	1.7	840	31	drysdale et al., 2009
Corchia	CC1593-594	138.6	1.7	840	35	drysdale et al., 2009
Corchia	CC5-722	133.4	1.7	850	239	drysdale et al., 2009
Corchia	CC5-749	130.6	1.7	850	236	drysdale et al., 2009
Corchia	CC7331-332	126.3	1.7	900	139	drysdale et al., 2009
Corchia	CC1-503	128.5	1.6	840	22	drysdale et al., 2009
Corchia	CC1623-624	142.6	1.6	840	38	drysdale et al., 2009
Corchia	CC1-508	129	1.5	840	26	drysdale et al., 2009
Corchia	CC1605-606	139	1.5	840	37	drysdale et al., 2009
Corchia	CC5673-671	141.5	1.5	850	244	drysdale et al., 2009
Corchia	CC5-707	136.8	1.5	850	240	drysdale et al., 2009
Corchia	*CC1-528	127.6	1.4	840	29	drysdale et al., 2009
Corchia	CC5-136	138.9	1.3	850	242	drysdale et al., 2009
Corchia	CC5-328	126.6	1.3	850	144	drysdale et al., 2009
Corchia	CC5-4R	120.7	1.3	850	120	drysdale et al., 2009
Corchia	CC5-6	128.5	1.3	850	169	drysdale et al., 2009
Corchia	CC5784-783	130.4	1.3	850	232	drysdale et al., 2009
Corchia	CC5-877	127.5	1.3	850	223	drysdale et al., 2009
Corchia	CC5a467-469	121.7	1.3	850	128	drysdale et al., 2009
Corchia	*CC5834-836	130	1.2	850	227	drysdale et al., 2009
Corchia	CC5-32	129.4	1.2	850	232	drysdale et al., 2009
Corchia	CC5-7R	134	1.2	850	239	drysdale et al., 2009
Corchia	CC5c210-212	123.6	1.2	850	131	drysdale et al., 2009
Corchia	CC5f17-19	127.2	1.2	850	134	drysdale et al., 2009
Corchia	CC7-11	124.7	1.2	900	138	drysdale et al., 2009
Corchia	CC7728-732	122.9	1.2	900	125	drysdale et al., 2009
Corchia	CC7804-806	124.8	1.2	900	138	drysdale et al., 2009
Corchia	CC1773	128.3	1.1	840	19	drysdale et al., 2009
Corchia	CC5-30	128.3	1.1	850	233	drysdale et al., 2009
Corchia	CC5a504-506	125.5	1.1	850	144	drysdale et al., 2009
Corchia	CC5c203+205	122.9	1.1	850	132	drysdale et al., 2009
Corchia	CC5f4+6	126	1.1	850	133	drysdale et al., 2009
Corchia	*CC7688-692	123.4	1	850	121	drysdale et al., 2009
Corchia	CC5-288	118.8	1	850	101	drysdale et al., 2009
Corchia	CC5-31	128.4	1	850	230	drysdale et al., 2009
Corchia	CC5c4+6	119	1	850	102	drysdale et al., 2009
Corchia	CC5d73	118	1	850	93	drysdale et al., 2009
Corchia	CC7635-639	121.5	1	850	115	drysdale et al., 2009
Corchia	CC5a327-331	126.8	0.9	850	147	drysdale et al., 2009
Corchia	CC5a572-575	125.6	0.9	850	137	drysdale et al., 2009
Corchia	CC5d61-65	118	0.9	850	94	drysdale et al., 2009
Corchia	CC5d76-80	118.6	0.9	850	92	drysdale et al., 2009

Corchia	CC5f9-11	125.2	0.9	850	133	drysdale et al., 2009
Corchia	CC1511-513	128.5	0.8	840	27	drysdale et al., 2009
Corchia	CC5-333	127.6	0.8	850	148	drysdale et al., 2009
Corchia	CC5a437-439	122.5	0.8	850	125	drysdale et al., 2009
Corchia	CC26-10	7.16	0.67	890	96	Zanchetta et al., 2007
Corchia	CC26-15	9.54	0.67	890	133	Zanchetta et al., 2007
Corchia	CC26-9	6.52	0.67	890	86	Zanchetta et al., 2007
Corchia	CC26-14	8.8	0.67	890	123	Zanchetta et al., 2007
Corchia	CC26-12	7.98	0.67	890	108	Zanchetta et al., 2007
Corchia	CC26-13	8.2	0.67	890	112	Zanchetta et al., 2007
Corchia	CC26-2	6	0.67	890	77	Zanchetta et al., 2007
Corchia	CC26-8	6.54	0.67	890	85	Zanchetta et al., 2007
Corchia	CC26-16	10.14	0.67	890	142	Zanchetta et al., 2007
Corchia	CC26-11	7.62	0.66	890	102	Zanchetta et al., 2007
Corchia	CC26-17	11.26	0.66	890	152	Zanchetta et al., 2007
Corchia	CC26-7	5.57	0.66	890	68	Zanchetta et al., 2007
Corchia	CC26-4	0.75	0.66	890	4	Zanchetta et al., 2007
Corchia	CC26-1	10.93	0.66	890	151	Zanchetta et al., 2007
Corchia	CC26-5	2.48	0.66	890	27	Zanchetta et al., 2007
Corchia	CC26-3	3.9	0.66	890	45	Zanchetta et al., 2007
Corchia	CC26-6	4.65	0.66	890	54	Zanchetta et al., 2007
Corchia	CC7-1	9.736	0.15	900	2	this work
Corchia	CC74.5	10.15	0.07	900	5	this work
Corchia	H	10.460	0.118	900	5	this work
Corchia	G	10.735	0.093	900	15	this work
Corchia	F	10.931	0.059	900	22	this work
Corchia	CC727.5	10.674	0.117	900	28	this work
Corchia	E	11.142	0.095	900	30	this work
Corchia	D	11.109	0.078	900	37	this work
Corchia	CC739.5	10.992	0.053	900	40	this work
Corchia	C	11.436	0.098	900	48	this work
Corchia	CC754.1	11.212	0.098	900	54	this work
Corchia	B	11.615	0.124	900	59	this work
Corchia	A	11.862	0.097	900	68	this work
Corchia	CC776.1	12.080	0.110	900	76	this work
Corchia	CD4A-2	3.504	0.069	910	0.4	this work
Corchia	CD4A-212	11.079	0.149	910	42.2	this work
Corchia	CD4A-244	11.853	0.095	910	48.0	this work
Corchia	CD4A-286	16.322	0.273	910	56.0	this work
Corchia	CD4A-335	33.929	0.544	910		this work
Corchia	CD4A-340	44.938	0.595	910	65.6	this work
Corchia	CD4A-360	66.097	1.091	910	66.6	this work
Corchia	CD4A-43.44	4.556	0.036	910	70.6	this work
Corchia	CD4T-15	29.129	0.425	910	8.6-8.8	this work
Corchia	CD4T-26	26.482	0.345	910	64.6	this work
Corchia	CD4T-40	21.619	0.328	910	62.4	this work
Corchia	CD4-3	71.839	0.763	910	59.8	this work
Corchia	CD4-4	94.611	0.988	910	75.0	this work
Corchia	CD4-5	103.872	2.011	910	131.0	this work
Corchia	CC27 1005	15.61	0.06	920	165.0	this work
Corchia	CC27 1008,1009,1010	12.977	0.174	920	202	this work
Corchia	CC27 1018,1019	12.152	0.135	920	204	this work
Corchia	CC27 1035,1036	11.682	0.165	920	21	this work
Corchia	CC27 1041,1042	11.155	0.153	920	208	this work
Corchia	CC27 1071,1072,1073	10.239	0.099	920	214	this work

Corchia	CC27 1080	10.08	0.05	920	189	this work
Corchia	CC27 1160	9.72	0.06	920	173	this work
Corchia	CC27 1164	8.99	0.05	920	172	this work
Corchia	CC27 1183	8.44	0.07	920	168	this work
Corchia	CC27 140	3.53	0.03	920	28	this work
Corchia	CC27 275	4.09	0.02	920	55	this work
Corchia	CC27 31,32,33	2.761	0.071	920	6	Isola et al., 2018
Corchia	CC27 410	4.99	0.03	920	82	this work
Corchia	CC27 437	5.189	0.072	920	87	Isola et al., 2018
Corchia	CC27 497	5.267	0.072	920	99	Isola et al., 2018
Corchia	CC27 555	5.46	0.02	920	111	this work
Corchia	CC27 844,845	8.426	0.149	920	169	this work
Corchia	CC27 857,858	8.794	0.131	920	171	this work
Corchia	CC27 868	9.946	0.168	920	174	this work
Corchia	CC27 895	10.181	0.114	920	179	this work
Corchia	CC27 9	2.67	0.02	920	2	this work
Corchia	CC27-1000-1002	18.897	0.130	920	200	this work
Corchia	CC27-1045-1046	11.292	0.118	920	209	this work
Corchia	CC27-1135-1138	10.178	0.081	920	227	this work
Corchia	CC27-118	3.451	0.053	920	24	Isola et al., 2018
Corchia	CC27-220	4.125	0.035	920	44	Isola et al., 2018
Corchia	CC27-492	5.393	0.039	920	98	Isola et al., 2018
Corchia	CC27-608	5.809	0.054	920	122	Isola et al., 2018
Corchia	CC7-303/305	4.928	0.301	920	61	this work
Corchia	CC27-502/503	5.324	0.21	920	100	Isola et al., 2018
Corchia	CC27-809	7.200	0.112	920	122	Isola et al., 2018
Corchia	CC22 97-101	79.139	0.557	930	10	this work
Corchia	CC22 722	81.862	0.613	930	62	this work
Corchia	CC22 464	80.944	0.634	930	37	this work
Corchia	CC22 207-211	80.753	0.774	930	21	this work
Corchia	CC22 1325-1328	94.574	0.709	930	116	this work
Corchia	CC22 1280-1282	91.933	1.436	930	112	this work
Corchia	CC22 1157	88.177	0.667	930	99	this work
Corchia	CC22 1136-1139	84.109	0.637	930	97	this work
Corchia	CD20-1	1.233	0.147	940	2.00	this work
Corchia	CD20-2	7.530	0.163	940	2.7	this work
Corchia	CD20-3	10.002	0.103	940	3.8	this work
Corchia	CD20-4	12.755	0.145	940	4.3	this work
Corchia	CD20-5	99.702	1.112	940	5.0	this work
Corchia	CD20-6	117.398	1.631	940	5.4	this work
Corchia	CD20-7	129.286	1.566	940	7.0	this work
Corchia	CD20-8	151.382	2.126	940	7.9	this work
Corchia	CC53-A	0.478	0.052	950	1	this work
Corchia	CC53-6	1.106	0.054	950	6	this work
Corchia	CC53-15	1.668	0.07	950	15	this work
Corchia	CC53-B	2.092	0.035	950	40	this work
Corchia	CC53-65	2.665	0.036	950	65	this work
Corchia	CC53-C	3.147	0.035	950	87	this work
Corchia	CC53-112	4.133	0.125	950	112	this work
Corchia	CC53-114	4.306	0.055	950	114	this work
Corchia	CC53-123	4.544	0.054	950	123	this work
Corchia	CC53-154	5.175	0.108	950	154	this work
Corchia	CC53-156	5.466	0.127	950	156	this work
Corchia	CC53-D	5.249	0.055	950	160	this work
Corchia	CC53-170	5.557	0.114	950	170	this work

Corchia	CC53-186	6.229	0.095	950	186	this work
Corchia	CC4-1	5.971	0.057	960		this work
Corchia	CC4-A	9.731	0.156	960	319	this work
Corchia	CC4-B	8.421	0.154	960	202	this work
Corchia	CC4-C	7.341	0.226	960	113	this work
Corchia	CC4-D	6.577	0.128	960	91	this work
Corchia	CC4-E	4.706	0.199	960	41	this work
Corchia	CC4-F	3.257	0.165	960	8	this work
Corchia	CC4-23	4.040	0.071	960	23	this work
Corchia	CC4-39	4.960	0.077	960	39	this work
Corchia	CC4-60	5.864	0.070	960	60	this work
Corchia	CC4-78	6.370	0.055	960	78	this work
Corchia	CC4-84	6.516	0.072	960	84	this work
Corchia	CC4-139	7.484	0.074	960	139	this work
Corchia	CC4-186	8.108	0.058	960	186	this work
Grotta del vento	GV1-83_C	6.66	0.96	670	83	Piccini et al., 2003
Grotta del vento	GV1-142_B	7.15	0.96	670	142	Piccini et al., 2006
Grotta del vento	GV1-169_C	7.78	1.72	670	169	Piccini et al., 2007
Grotta del vento	GV1-264_C	9.08	3.9	670	264	Piccini et al., 2009
Grotta del vento	GV3-35_C	6.03	2.42	670	39	Piccini et al., 2010
Grotta del vento	GV3-95_C	4.45	2.02	670	107	Piccini et al., 2011
Grotta del vento	GV3-Base_C	11.46	1.32	670	133	Piccini et al., 2012
Tana che Urla	TCUD3E	111.826	5.5	640	313	Regattieri et al., 2014
Tana che Urla	TCUD4-102.5A	120.44	5.36	640	103	Regattieri et al., 2014
Tana che Urla	TCUD4-102.5B	122.15	4.68	640	102	Regattieri et al., 2014
Tana che Urla	TCUD3C	119.164	4.16	640	459	Regattieri et al., 2014
Tana che Urla	TCU.D3B_e	144.1	3.98	640	510	Regattieri et al., 2014
Tana che Urla	TCUD4-77.5	80.52	3.56	640	78	Regattieri et al., 2014
Tana che Urla	TCUD4-2	155.3	3.12	640	288	Regattieri et al., 2014
Tana che Urla	TCUD3Abis.b	100.25	2.97	640	281	Regattieri et al., 2014
Tana che Urla	TCUD4-1	152.83	2.94	640	310	Regattieri et al., 2014
Tana che Urla	TCUD4-263.5	146.38	2.91	640	264	Regattieri et al., 2014
Tana che Urla	TCUD4-244.5	139.44	2.89	640	245	Regattieri et al., 2014
Tana che Urla	TCUD4-90.5	112.55	2.77	640	91	Regattieri et al., 2014
Tana che Urla	TCUD4-70.5	81.12	2.61	640	71	Regattieri et al., 2014
Tana che Urla	TCUD4-212.5	133.73	2.17	640	213	Regattieri et al., 2014
Tana che Urla	TCUD4-3	131.23	2.01	640	226	Regattieri et al., 2014
Tana che Urla	TCUD4-181.5	129.24	1.98	640	182	Regattieri et al., 2014
Tana che Urla	TCUD4-4	128.1	1.98	640	157	Regattieri et al., 2014
Tana che Urla	TCUD4-A	159.08	1.3	640	335	Regattieri et al., 2014
Tana che Urla	TCUD4-C	124.3	1.16	640	124	Regattieri et al., 2014
Tana che Urla	TCUD4-B	129.21	1.1	640	208	Regattieri et al., 2014
Tana che Urla	TCUD4-46.5	75.03	0.78	640	47	Regattieri et al., 2014
Tana che Urla	TCUD4-24.5	12.61	0.65	640	25	Regattieri et al., 2014
Tana che Urla	TCUD4-D	7.79	0.28	640	9	Regattieri et al., 2014
Renella	RL4-2	1.77	0.77	275	15	Zhornyak et al., 2011
Renella	RL4-13	4.66	0.44	275	112	Zhornyak et al., 2011
Renella	RL4-8	6.87	0.43	275	153	Zhornyak et al., 2011
Renella	RL18-6	8.96	0.32	275	23	Zhornyak et al., 2011
Renella	RL13-8	9.13	0.3	275	39	Zhornyak et al., 2011
Renella	RL4-7	5.94	0.28	275	136	Zhornyak et al., 2011
Renella	RL13-2	9.26	0.19	275	43	Zhornyak et al., 2011
Renella	RL18-5	8.29	0.18	275	2	Zhornyak et al., 2011
Renella	RL13-7	8.68	0.17	275	36	Zhornyak et al., 2011
Renella	RL18-8	12.53	0.17	275	69	Zhornyak et al., 2011

Renella	RL13-3	4.57	0.16	275	34	Zhornyak et al., 2011
Renella	RL4-9	2.09	0.16	275	26	Zhornyak et al., 2011
Renella	RL18-4	11.85	0.15	275	60	Zhornyak et al., 2011
Renella	RL18-7	9.7	0.15	275	42	Zhornyak et al., 2011
Renella	RL13-6	9.72	0.14	275	41	Zhornyak et al., 2011
Renella	RL13-1	10.61	0.12	275	57	Zhornyak et al., 2011
Renella	RL4-6	5.44	0.11	275	120	Zhornyak et al., 2011
Renella	RL13-4	3.8	0.1	275	20	Zhornyak et al., 2011
Renella	RL18-2	12.54	0.1	275	84	Zhornyak et al., 2011
Renella	RL4-11	3.63	0.1	275	77	Zhornyak et al., 2011
Renella	RL4-5	5.35	0.1	275	117	Zhornyak et al., 2011
Renella	RL4-10	3.11	0.09	275	56	Zhornyak et al., 2011
Renella	RL4-12	4.21	0.09	275	104	Zhornyak et al., 2011
Renella	RL18-3	8.86	0.08	275	10	Zhornyak et al., 2011
Renella	RL4-4	3.62	0.07	275	86	Zhornyak et al., 2011
Renella	RL13-5	0.9	0.05	275	3	Zhornyak et al., 2011
Renella	RL4-1	1.27	0.05	275	2	Zhornyak et al., 2011
Renella	RL4-3	2.28	0.05	275	31	Zhornyak et al., 2011
Buca Cava dell'Onice	A	83.76	1.06	700	14	this work
Buca Cava dell'Onice	B	104.25	2.28	700	23	this work
Buca Cava dell'Onice	C	107.29	1.25	700	28	this work
Buca Cava dell'Onice	J	109.89	0.93	700	69	this work
Buca Cava dell'Onice	D	122.88	1.45	700	78	this work
Buca Cava dell'Onice	E	122.93	1.53	700	99	this work
Buca Cava dell'Onice	F	145.48	2.88	700	110	this work

The activity ratios are standardized to the HU-1 secular equilibrium, and ages calculated using decay constants of 9.195_{-10}^{-6} (230Th) and 2.835_{-10}^{-6} (234U). 2σ errors (column F) are the 95% uncertainties. See text and or the original papers for further information, references in the last column.

Table 2 Elevation used to to estimate the MAAT in the speleothem catchment areas

CAVE	Elevation (m, a.s.l.)		
	Sampling Site	C1	C2
Corchia Cave (CC)	840	1240	1460
Grotta Del Vento (GDV)	670	650	1230
Tana Che Urla (TCU)	640	650	1250
Renella (REN)	275	340	1045

Elevation of sampling sites in the cave (column B); the elevations of surface above the sampling site (C1); mean elevation values between the surface above sampling site and the maximum elevation of the basin (C2). All values are in metres a.s.l.