An overview of some recent results on the Euler system of isentropic gas dynamics

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Abstract

This overview is concerned with the well-posedness problem for the isentropic compressible Euler equations of gas dynamics. The results we present are in line with the program of investigating the efficiency of different selection criteria proposed in the literature in order to weed out non-physical solutions to more-dimensional systems of conservation laws and they build upon the method of convex integration developed by De Lellis and Székelyhidi for the incompressible Euler equations. Mainly following [5], we investigate the role of the maximal dissipation criterion proposed by Dafermos in [6]: we prove how, for specific pressure laws, some non-standard (i.e. constructed via convex integration methods) solutions to the Riemann problem for the isentropic Euler system in two space dimensions have greater energy dissipation rate than the classical self-similar solution emanating from the same Riemann data. We therefore show that the maximal dissipation criterion proposed by Dafermos does not favour in general the self-similar solutions.

1 Introduction

We consider the isentropic compressible Euler system of gas dynamics in 2 space dimensions (cf. [8] or [15] or [1]). It is obtained as a simplification of the full compressible Euler equations, by assuming the entropy to be constant. The state of the gas is described through the state vector

$$V = (\rho, v)$$

whose components are the density ρ and the velocity v. The system consists of 3 equations which correspond to balance statements for mass and linear momentum. The corresponding Cauchy problems reads as

(1.1)
$$\begin{cases} \partial_t \rho + \operatorname{div}_x(\rho v) = 0 \\ \partial_t(\rho v) + \operatorname{div}_x(\rho v \otimes v) + \nabla_x[p(\rho)] = 0 \\ \rho(\cdot, 0) = \rho^0 \\ v(\cdot, 0) = v^0, \end{cases}$$

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with $t \in \mathbb{R}^+$, $x \in \mathbb{R}^2$. The pressure p is a function of the density ρ determined from the constitutive thermodynamic relations of the gas under consideration and it is assumed to satisfy p' > 0 (under this assumption the system is hyperbolic). We will work with pressure laws $p(\rho) = \rho^{\gamma}$ with constant $\gamma \geq 1$.

Our aim is to discuss the issue of uniqueness of weak solutions to the Cauchy problem (1.1). The theory of the Cauchy problem for hyperbolic systems of conservation laws is typically confronted with two major challenges. First, it is well-known that classical solutions develop discontinuities, even starting out from smooth initial data. In the literature this behaviour is known as breakdown of classical solutions. Therefore, it becomes imperative to introduce the notion of weak solution. However, weak solutions fail to be unique. In order to restore uniqueness restrictions need to be imposed in hope of singling out a unique physical solution. When dealing with systems coming from Physics, as in our case, the second law of Thermodynamics naturally induces such restrictions, such admissibility criteria by stipulating that weak solutions are admissible/entropy solutions if they satisfy some entropy inequalities (see (2.3) for the specific case of the compressible Euler system). Finally, a third important challenge then arises: do entropy inequalities really serve as selection criteria? Are admissible solutions unique? Or at least, do there exist efficient criteria restoring uniqueness? This is a central problem to set down a complete theory for the Cauchy problem. It has deserved a lot of attention, but positive answers were found only for scalar conservation laws or for systems in one space dimensions (under smallness assumptions on the initial data): for a complete account of the existing literature we refer the reader to [8] and [15]. When dealing with systems of conservation laws in more than one space dimension, it is still an intriguing mathematical problem to develop a theory of well-posedness for the Cauchy problem which includes the formation and evolution of shock waves.

In 2006, Elling [12] studied numerically a particular case of initial data for the two dimensional non-isentropic Euler equations. His results show that the numerical method does not always converge to the physical solution. Moreover, they suggest that entropy solutions (in the weak entropy inequality sense) to the multi-dimensional Euler equations are not always unique.

In a groundbreaking paper [10], De Lellis-Székelyhidi give an example in favour of the conjecture that entropy/admissible solutions to the multi-dimensional compressible Euler equations are in general not unique. The non-uniqueness result by De Lellis-Székelyhidi is a byproduct of their analysis of the incompressible Euler equations based on its formulation as a differential inclusion (see [9] and [11]) combined with convex integration methods: they exploit the result for the incompressible Euler equations to exhibit bounded initial density and bounded compactly supported initial velocity for which admissible solutions of (1.1) are not unique (in more than one space dimension). However the initial data constructed in [10] are very irregular. The result by De Lellis and Székelyhidi is improved by the author in [2] where it is proven that non-uniqueness still holds in the case of regular initial density (see also [4] for further generalizations). Non-unique solutions constructed via convex integration are referred to as non-standard or oscillatory solutions. Moreover in [3], using the Riemann problem as a building block, the authors show that, in the two dimensional case, the entropy inequality (see (2.3)) does not single out unique weak solutions even under very strong assumptions on the initial data $((\rho^0, v^0) \in W^{1,\infty}(\mathbb{R}^2))$:

Theorem 1.1 (Chiodaroli, De Lellis, Kreml) There are Lipschitz initial data (ρ^0, v^0)

for which there are infinitely many bounded admissible solutions (ρ, v) of (1.1) on $\mathbb{R}^2 \times [0, \infty[$ with $\inf \rho > 0$. These solutions are all locally Lipschitz on a finite interval on which they all coincide with the unique classical solution.

This is proven by constructing infinitely many entropy weak solutions in forward time to a Riemann problem for (1.1) whose Riemann data can be generated, backwards in time, by a classical compression wave: the Lipschitz initial data of Theorem 1.1 will be provided by the values of the compression wave at some finite negative time. It is clear now that the infinitely many admissible solutions constructed in Theorem 1.1 all coincide with the unique classical solution (compression wave) on a finite time interval whereas non–uniqueness arises after the first blow–up time.

This series of negative results concerning the entropy inequality as selection criterion for system (1.1) motivated the authors to explore other admissibility criteria which could work in favour of uniqueness, in particular we investigated an alternative criterion which has been proposed by Dafermos in [6] under the name *entropy rate admissibility criterion*. The ideas developed in [3] enabled us to prove in [5] the following theorem:

Theorem 1.2 (Chiodaroli, Kreml) Let $p(\rho) = \rho^{\gamma}$, $1 \leq \gamma < 3$. There exist Riemann data for which the self-similar solution to (1.1) emanating from these data is not entropy rate admissible.

This result does not exclude that the entropy rate admissibility criterion could still select a unique solution, but surely prevents the self–similar solution to be the selected one. Moreover, since Theorem 1.2 is proven using non–standard solutions as competitors, with respect to Dafermos' criterion, for the self–similar solution, we can affirm that the entropy rate criterion cannot, at least in our setting, rule out oscillatory solutions obtained via convex integration. In the rest of the paper we will further explain this result.

2 Entropy criteria

2.1 Entropy inequality

We recall here the usual definitions of weak and admissible solutions to (1.1) in the twodimensional case.

Definition 2.1 By a weak solution of (1.1) on $\mathbb{R}^2 \times [0, \infty[$ we mean a pair $(\rho, v) \in L^{\infty}(\mathbb{R}^2 \times [0, \infty[)$ such that the following identities hold for every test functions $\psi \in C_c^{\infty}(\mathbb{R}^2 \times [0, \infty[), \phi \in C_c^{\infty}(\mathbb{R}^2 \times [0, \infty[)))$:

(2.1)
$$\int_0^\infty \int_{\mathbb{R}^2} \left[\rho \partial_t \psi + \rho v \cdot \nabla_x \psi \right] dx dt + \int_{\mathbb{R}^2} \rho^0(x) \psi(x, 0) dx = 0$$

$$(2.2) \qquad \int_0^\infty \int_{\mathbb{R}^2} \left[\rho v \cdot \partial_t \phi + \rho v \otimes v : D_x \phi + p(\rho) \operatorname{div}_x \phi \right] + \int_{\mathbb{R}^2} \rho^0(x) v^0(x) \cdot \phi(x, 0) dx \ = \ 0.$$

Definition 2.2 A bounded weak solution (ρ, v) of (1.1) is admissible if it satisfies the following inequality for every nonnegative test function $\varphi \in C_c^{\infty}(\mathbb{R}^2 \times [0, \infty[):$

$$\int_{0}^{\infty} \int_{\mathbb{R}^{2}} \left[\left(\rho \varepsilon(\rho) + \rho \frac{|v|^{2}}{2} \right) \partial_{t} \varphi + \left(\rho \varepsilon(\rho) + \rho \frac{|v|^{2}}{2} + p(\rho) \right) v \cdot \nabla_{x} \varphi \right]
+ \int_{\mathbb{R}^{2}} \left(\rho^{0}(x) \varepsilon(\rho^{0}(x)) + \rho^{0}(x) \frac{|v^{0}(x)|^{2}}{2} \right) \varphi(x, 0) dx \ge 0.$$

Note that (2.3) is rather a weak form of energy balance.

2.2 Entropy rate admissibility criterion

An alternative criterion to the entropy inequality has been proposed by Dafermos in [6] under the name of entropy rate admissibility criterion. In order to formulate this criterion for the specific system (1.1) we define the total energy of the solutions (ρ, v) to (1.1) as

(2.4)
$$E[\rho, v](t) = \int_{\mathbb{R}^2} \left(\rho \varepsilon(\rho) + \rho \frac{|v|^2}{2} \right) dx.$$

Let us remark that in Dafermos' terminology $E[\rho, v](t)$ is called "total entropy" (see [6]). However, since in the context of system (1.1) the physical energy plays the role of the mathematical entropy, it is more convenient to call $E[\rho, v](t)$ total energy. The right derivative of $E[\rho, v](t)$ defines the energy dissipation rate of (ρ, v) at time t:

(2.5)
$$D[\rho, v](t) = \frac{\mathrm{d}_+ E[\rho, v](t)}{\mathrm{d}t}.$$

Since our solutions will have piecewise constant values of ρ and $|v|^2$ and it is easy to see that the total energy of any solution we construct is infinite, we shall restrict the infinite domain \mathbb{R}^2 to a finite box $(-L, L)^2$ and denote

(2.6)
$$E_L[\rho, v](t) = \int_{(-L, L)^2} \left(\rho \varepsilon(\rho) + \rho \frac{|v|^2}{2} \right) dx$$

(2.7)
$$D_L[\rho, v](t) = \frac{\mathrm{d}_+ E_L[\rho, v](t)}{\mathrm{d}t}.$$

The problem of infinite energy of solutions may be solved also by restricting to a periodic domain and constructing (locally in time) periodic solutions. This procedure is carefully described in [5].

Using the concept of energy dissipation rate, Dafermos in [6] introduces a new selection criterion for weak solutions which goes under the name of *entropy rate admissibility criterion*. We recall here the definition of *entropy rate admissible solutions*.

Definition 2.3 (Entropy rate admissible solution) A weak solution (ρ, v) of (1.1) is entropy rate admissible if there exists $L^* > 0$ such that there is no other weak solution $(\overline{\rho}, \overline{v})$ with the property that for some $\tau \geq 0$, $(\overline{\rho}, \overline{v})(x, t) = (\rho, v)(x, t)$ on $\mathbb{R}^2 \times [0, \tau]$ and $D_L[\overline{\rho}, \overline{v}](\tau) < D_L[\rho, v](\tau)$ for all $L \geq L^*$.

In other words, we call entropy rate admissible the solution(s) dissipating most total energy.

3 Background literature and main results

The investigation of the entropy rate admissibility criterion initiated with the paper [6] of Dafermos where he puts it forward and moreover proves that for a single equation the entropy rate criterion is equivalent to the viscosity criterion in the class of piecewise smooth solutions. Later on, following the approach of Dafermos, Hsiao in [14] proves, in the class of piecewise smooth solutions, the equivalence of the entropy rate criterion and the viscosity criterion for the one-dimensional system of equations of nonisentropic gas dynamics in lagrangian formulation with pressure laws $p(\rho) = \rho^{\gamma}$ for $\gamma \geq 5/3$ while the same equivalence is disproved for $\gamma < 5/3$. For further analysis on the relation between entropy rate minimization and admissibility of solutions for a more general class of evolutionary equations we refer to [7]. However, to our knowledge, up to some time ago the entropy rate criterion had not been tested in the case of several space variables and on broader class of solutions than the piecewise smooth ones.

Recently Feireisl in [13] extended the result of Chiodaroli [2] and obtained infinitely many global admissible weak solutions of (1.1) none of which is entropy rate admissible: this results seems in favour of the effectiveness of the entropy rate criterion to rule out non-standard solutions (i.e. constructed by the method of De Lellis and Székelyhidi). In [5] we have actually shown that for specific initial data, and in the two-dimensional case, the oscillatory (non-standard) solutions dissipate more energy than the self-similar solution which may be believed to be the physical one. The results obtained in [3] hinge upon some of the ideas devised in [3] combined with novel developments to deal with the entropy rate criterion.

We refer also to the work [16], where Székelyhidi constructed irregular solutions of the incompressible Euler equations with vortex-sheet initial data and computed their dissipation rate.

We focus on the Riemann problem for the system (1.1),(2.3) in two-space dimensions. Hence, we denote the space variable as $x = (x_1, x_2) \in \mathbb{R}^2$ and consider initial data in the form

(3.1)
$$(\rho^{0}(x), v^{0}(x)) := \begin{cases} (\rho_{-}, v_{-}) & \text{if } x_{2} < 0 \\ (\rho_{+}, v_{+}) & \text{if } x_{2} > 0, \end{cases}$$

where $\rho_{\pm}, v_{\pm 1}, v_{\pm 2}$ are constants. Our concern has been to compare the energy dissipation rate of standard self-similar solutions associated to the Riemann problem (1.1), (2.3), (3.1) with the energy dissipation rate of non-standard solutions for the same problem obtained by the method developed in [3].

We obtained the following results.

Theorem 3.1 Let $p(\rho) = \rho^{\gamma}$ with $\gamma \geq 1$. For every Riemann data (3.1) such that the self-similar solution to the Riemann problem (1.1)–(2.3), (3.1) consists of an admissible 1–shock and an admissible 3–shock, i.e. $v_{-1} = v_{+1}$ and

(3.2)
$$v_{+2} - v_{-2} < -\sqrt{\frac{(\rho_{-} - \rho_{+})(p(\rho_{-}) - p(\rho_{+}))}{\rho_{-}\rho_{+}}},$$

there exist infinitely many admissible solutions to (1.1)–(2.3), (3.1).

Theorem 3.1 can be viewed as an extension of the results obtained together with De Lellis in [3]. As a consequence of Theorem 3.1 and by a suitable choice of initial data, we can prove the following main theorem.

Theorem 3.2 Let $p(\rho) = \rho^{\gamma}$, $1 \leq \gamma < 3$. There exist Riemann data (3.1) for which the self-similar solution to (1.1),(2.3) emanating from these data is not entropy rate admissible.

Theorem 3.2 ensures that for $1 \le \gamma < 3$ there exist initial Riemann data (3.1) for which some of the infinitely many nonstandard solutions constructed as in Theorem 3.1 dissipate more energy than the self-similar solution, suggesting in particular that the Dafermos entropy rate admissibility criterion would not pick the self-similar solution as the admissible one.

4 Strategies of proof

In this section we explain how to prove Theorem 3.1 and 3.2. For the complete proofs we refer the reader to [5].

Both theorems stem from the framework introduced in [3] where the Riemann problem constitutes the starting point for constructing non–unique admissible non–standard solutions. In particular, in [3], the authors jointly with Camillo De Lellis developed a method which allows to obtain infinitely many entropy (oscillatory) solutions to a Riemann problem provided a suitable admissible subsolution exists.

4.1 From subsolutions to solutions

We shall introduce the notion of fan subsolution and admissible fan subsolution as in [3, Section 3].

Definition 4.1 (Fan partition) A fan partition of $\mathbb{R}^2 \times (0, \infty)$ consists of three open sets P_-, P_1, P_+ of the following form

$$(4.1) P_{-} = \{(x,t) : t > 0 \quad and \quad x_2 < \nu_{-}t\}$$

$$(4.2) P_1 = \{(x,t) : t > 0 \quad and \quad \nu_- t < x_2 < \nu_+ t\}$$

$$(4.3) P_{+} = \{(x,t) : t > 0 \quad and \quad x_2 > \nu_{+}t\},$$

where $\nu_{-} < \nu_{+}$ is an arbitrary couple of real numbers.

We denote by $\mathcal{S}_0^{2\times 2}$ the set of all symmetric 2×2 matrices with zero trace.

Definition 4.2 (Fan subsolution) A fan subsolution to the compressible Euler equations (1.1) with initial data (3.1) is a triple $(\overline{\rho}, \overline{v}, \overline{u}) : \mathbb{R}^2 \times (0, \infty) \to (\mathbb{R}^+, \mathbb{R}^2, \mathcal{S}_0^{2 \times 2})$ of piecewise constant functions satisfying the following requirements.

(i) There is a fan partition P_-, P_1, P_+ of $\mathbb{R}^2 \times (0, \infty)$ such that

$$(\overline{\rho}, \overline{v}, \overline{u}) = (\rho_-, v_-, u_-) \mathbf{1}_{P_-} + (\rho_1, v_1, u_1) \mathbf{1}_{P_1} + (\rho_+, v_+, u_+) \mathbf{1}_{P_-}$$

where ρ_1, v_1, u_1 are constants with $\rho_1 > 0$ and $u_{\pm} = v_{\pm} \otimes v_{\pm} - \frac{1}{2} |v_{\pm}|^2 \mathrm{Id}$;

(ii) There exists a positive constant C such that

$$(4.4) v_1 \otimes v_1 - u_1 < \frac{C}{2} \operatorname{Id};$$

(iii) The triple $(\overline{\rho}, \overline{v}, \overline{u})$ solves the following system in the sense of distributions:

$$(4.5) \partial_t \overline{\rho} + \operatorname{div}_r(\overline{\rho} \, \overline{v}) = 0$$

$$(4.6) \partial_t(\overline{\rho}\,\overline{v}) + \operatorname{div}_x(\overline{\rho}\,\overline{u}) + \nabla_x\left(p(\overline{\rho}) + \frac{1}{2}\left(C\rho_1\mathbf{1}_{P_1} + \overline{\rho}|\overline{v}|^2\mathbf{1}_{P_+\cup P_-}\right)\right) = 0.$$

Definition 4.3 (Admissible fan subsolution) A fan subsolution $(\overline{\rho}, \overline{v}, \overline{u})$ is said to be admissible if it satisfies the following inequality in the sense of distributions

$$\partial_{t} \left(\overline{\rho} \varepsilon(\overline{\rho}) \right) + \operatorname{div}_{x} \left[\left(\overline{\rho} \varepsilon(\overline{\rho}) + p(\overline{\rho}) \right) \overline{v} \right] + \partial_{t} \left(\overline{\rho} \frac{|\overline{v}|^{2}}{2} \mathbf{1}_{P_{+} \cup P_{-}} \right) + \operatorname{div}_{x} \left(\overline{\rho} \frac{|\overline{v}|^{2}}{2} \overline{v} \mathbf{1}_{P_{+} \cup P_{-}} \right) \\
+ \left[\partial_{t} \left(\rho_{1} \frac{C}{2} \mathbf{1}_{P_{1}} \right) + \operatorname{div}_{x} \left(\rho_{1} \overline{v} \frac{C}{2} \mathbf{1}_{P_{1}} \right) \right] \leq 0.$$

The strategy which lies behind Theorem 3.1, as well as behind Theorem 1.1 in [3], consists in reducing the proof of the existence of infinitely many admissible solutions to the Riemann problem for (1.1) to the proof of the existence of an admissible fan subsolution as defined in Definition 4.3. This is the content of the following Proposition which represents the key ingredient of [5] and is proven in [3].

Proposition 4.1 Let p be any C^1 function and (ρ_{\pm}, v_{\pm}) be such that there exists at least one admissible fan subsolution $(\overline{\rho}, \overline{v}, \overline{u})$ of (1.1) with initial data (3.1). Then there are infinitely many bounded admissible solutions (ρ, v) to (1.1)-(2.3), (3.1) such that $\rho = \overline{\rho}$ and $|v|^2 \mathbf{1}_{P_1} = C$.

Roughly speaking, the infinitely many bounded admissible solutions (ρ, v) of Proposition 4.1 are constructed by adding to the subsolution solutions to the linearized pressureless incompressible Euler equations supported in P_1 . Such solutions are given by the following Lemma, cf. [3, Lemma 3.7].

Lemma 4.2 Let $(\tilde{v}, \tilde{u}) \in \mathbb{R}^2 \times \mathcal{S}_0^{2 \times 2}$ and $C_0 > 0$ be such that $\tilde{v} \otimes \tilde{v} - \tilde{u} < \frac{C_0}{2}$ Id. For any open set $\Omega \subset \mathbb{R}^2 \times \mathbb{R}$ there are infinitely many maps $(\underline{v}, \underline{u}) \in L^{\infty}(\mathbb{R}^2 \times \mathbb{R}, \mathbb{R}^2 \times \mathcal{S}_0^{2 \times 2})$ with the following property

- (i) \underline{v} and \underline{u} vanish identically outside Ω ;
- (ii) $\operatorname{div}_{x}v = 0$ and $\partial_{t}v + \operatorname{div}_{x}u = 0$;

(iii)
$$(\tilde{v} + \underline{v}) \otimes (\tilde{v} + \underline{v}) - (\tilde{u} + \underline{u}) = \frac{C_0}{2} \text{Id } a.e. \text{ on } \Omega.$$

Proposition 4.1 is then proved by applying Lemma 4.2 with $\Omega = P_1$, $(\tilde{v}, \tilde{u}) = (v_1, u_1)$ and $C_0 = C$. It is then a matter of easy computation to check that each couple $(\overline{\rho}, \overline{v} + \underline{v})$ is indeed an admissible weak solution to (1.1)–(2.3) with initial data (3.1), for details see [3, Section 3.3]. The whole proof of Lemma 4.2 can be found in [3, Section 4].

4.2 Concluding arguments

Thanks to Proposition 4.1, Theorem 3.1 amounts to showing the existence of a fan admissible subsolution with appropriate initial data under the hypothesis that (3.1) is such that the self-similar solution to the Riemann problem (1.1), (2.3), (3.1) consists of an admissible 1—shock and an admissible 3—shock

Indeed, a fan admissible sunsolution with initial data (3.1) is defined through a the set of identities and inequalities which we recall here (see also [3, Section 5]).

We introduce the real numbers $\alpha, \beta, \gamma_1, \gamma_2, v_{-1}, v_{-2}, v_{+1}, v_{+2}$ such that

$$(4.8) v_1 = (\alpha, \beta),$$

$$(4.9) v_{-} = (v_{-1}, v_{-2})$$

$$(4.10) v_{+} = (v_{+1}, v_{+2})$$

$$(4.11) u_1 = \begin{pmatrix} \gamma_1 & \gamma_2 \\ \gamma_2 & -\gamma_1 \end{pmatrix}.$$

Then, Proposition 4.1 translates into the following set of algebraic identities and inequalities.

Proposition 4.3 Let P_-, P_1, P_+ be a fan partition as in Definition 4.1.

The constants $v_1, v_-, v_+, u_1, \rho_-, \rho_+, \rho_1$ as in (4.8)-(4.11) define an admissible fan subsolution as in Definitions 4.2-4.3 if and only if the following identities and inequalities hold:

• Rankine-Hugoniot conditions on the left interface:

$$(4.12) \nu_{-}(\rho_{-} - \rho_{1}) = \rho_{-}v_{-2} - \rho_{1}\beta$$

$$(4.13) \nu_{-}(\rho_{-}v_{-1} - \rho_{1}\alpha) = \rho_{-}v_{-1}v_{-2} - \rho_{1}\gamma_{2}$$

(4.14)
$$\nu_{-}(\rho_{-}v_{-2}-\rho_{1}\beta) = \rho_{-}v_{-2}^{2}+\rho_{1}\gamma_{1}+p(\rho_{-})-p(\rho_{1})-\rho_{1}\frac{C}{2};$$

• Rankine-Hugoniot conditions on the right interface:

$$(4.15) \nu_{+}(\rho_{1} - \rho_{+}) = \rho_{1}\beta - \rho_{+}v_{+2}$$

$$(4.16) \nu_{+}(\rho_{1}\alpha - \rho_{+}v_{+1}) = \rho_{1}\gamma_{2} - \rho_{+}v_{+1}v_{+2}$$

(4.17)
$$\nu_{+}(\rho_{1}\beta - \rho_{+}v_{+2}) = -\rho_{1}\gamma_{1} - \rho_{+}v_{+2}^{2} + p(\rho_{1}) - p(\rho_{+}) + \rho_{1}\frac{C}{2};$$

• Subsolution condition:

$$(4.18) \alpha^2 + \beta^2 < C$$

(4.19)
$$\left(\frac{C}{2} - \alpha^2 + \gamma_1\right) \left(\frac{C}{2} - \beta^2 - \gamma_1\right) - (\gamma_2 - \alpha\beta)^2 > 0;$$

• Admissibility condition on the left interface:

$$\nu_{-}(\rho_{-}\varepsilon(\rho_{-}) - \rho_{1}\varepsilon(\rho_{1})) + \nu_{-}\left(\rho_{-}\frac{|v_{-}|^{2}}{2} - \rho_{1}\frac{C}{2}\right)
(4.20) \leq \left[(\rho_{-}\varepsilon(\rho_{-}) + p(\rho_{-}))v_{-2} - (\rho_{1}\varepsilon(\rho_{1}) + p(\rho_{1}))\beta\right] + \left(\rho_{-}v_{-2}\frac{|v_{-}|^{2}}{2} - \rho_{1}\beta\frac{C}{2}\right);$$

• Admissibility condition on the right interface:

$$\nu_{+}(\rho_{1}\varepsilon(\rho_{1}) - \rho_{+}\varepsilon(\rho_{+})) + \nu_{+}\left(\rho_{1}\frac{C}{2} - \rho_{+}\frac{|v_{+}|^{2}}{2}\right)$$

$$(4.21) \qquad \leq \left[(\rho_{1}\varepsilon(\rho_{1}) + p(\rho_{1}))\beta - (\rho_{+}\varepsilon(\rho_{+}) + p(\rho_{+}))v_{+2}\right] + \left(\rho_{1}\beta\frac{C}{2} - \rho_{+}v_{+2}\frac{|v_{+}|^{2}}{2}\right).$$

Theorem 3.2 is then a corollary of the following theorem, proven in [5] and showing the existence of an admissible fan subsolution, combined with Proposition 4.1.

Theorem 4.1 Let $p(\rho) = \rho^{\gamma}$ with $\gamma \geq 1$. For every Riemann data (3.1) such that $v_{-1} = v_{+1}$ and

(4.22)
$$v_{+2} - v_{-2} < -\sqrt{\frac{(\rho_{-} - \rho_{+})(p(\rho_{-}) - p(\rho_{+}))}{\rho_{-}\rho_{+}}},$$

there exist $\nu_-, \nu_+, v_1, u_1, \rho_1, C$ such that (4.12)-(4.21) hold.

It remains to prove Theorem 3.2. This is obtained by showing that among the infinitely many admissible solutions provided by Theorem 3.1 one has lower energy dissipation rate than the self–similar solution emanating from the same Riemann data, thus contradicting Definition 2.3. Let us remark that the Riemann data allowing for the result of Theorem 3.2 are of the same type as the ones of Theorem 3.1, i.e. they admit a forward in time self–similar solution consisting of two shocks. We also underline that the self–similar solution depends in fact only on one variable, specifically on x_2 .

Assume from now on for simplicity that $v_{-1} = v_{+1} = 0$ in (3.1). Let us denote the self-similar solution emanating from the Riemann data as in Theorem 3.1. The value of the dissipation rate $D_L[\rho_S, v_S](t)$ has a specific form for the solution (ρ_S, v_S) consisting (by assumption) of two shocks of speeds ν_1 and ν_2 . Denoting the middle state $(\rho_m, v_m = (0, \overline{v}))$ and introducing the notation

(4.23)
$$E_{\pm} := \rho_{\pm} \varepsilon(\rho_{\pm}) + \rho_{\pm} \frac{v_{\pm}^2}{2}$$

(4.24)
$$E_m := \rho_m \varepsilon(\rho_m) + \rho_m \frac{\overline{v}^2}{2}$$

we have

(4.25)
$$D_L[\rho_S, v_S] = -2L\left(\nu_1(E_- - E_m) + \nu_2(E_m - E_+)\right).$$

Now let us consider a solution (ρ, v) with the same initial data (3.1) but constructed by the method of convex integration starting from an admissible fan subsolution using Proposition 4.1. We also assume, that the fan admissible subsolution (which exists by Theorem 4.1) has an underlying fan partition defined by the speeds ν_- and ν_+ . Although v is not constant in P_1 we still have, by construction (see Proposition 4.1), that $|v|^2 \mathbf{1}_{P_1} = C$, in particular $E_1 := \rho_1 \varepsilon(\rho_1) + \rho_1 \frac{C}{2}$ is constant in P_1 . The dissipation rate for all solutions constructed from a given subsolution hence depends only on the underlying subsolution and is given by

(4.26)
$$D_L[\rho, v] = -2L\left(\nu_-(E_- - E_1) + \nu_+(E_1 - E_+)\right).$$

If, for a moment, we assume that the speeds of the self-similar solution and of the subsolution coincide, i.e. $\nu_{-} = \nu_{1}$ and $\nu_{+} = \nu_{2}$, it is clearly enough to achieve $E_{1} > E_{m}$ in order to prove Theorem 3.2. Of course, as one can see from [5, Section 4] this is not the case; nonetheless the proof works along the same line: we can prove that there is still some freedom in the choice of the parameters defining the underlying subsolution for (ρ, v) which allows to tune them in such a way that Theorem 3.2 holds. For a complete and rigorous proof we refer to [5, Section 5].

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