

1 **Revisiting the Y-3 tephrostratigraphic marker: a new diagnostic**
2 **glass geochemistry, age estimate, and details on its**
3 **climatostratigraphical context**

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29 **Abstract**

30 The 'Y-3' tephra is a crucial stratigraphic marker within the central Mediterranean region that
31 falls close to the Marine Isotope Stage 3/2 transition and a cooling event proposed as a
32 correlative of the North Atlantic Heinrich Stadial 3 (HS3). Consequently, this tephra offers
33 great potential to assess any leads and lags in environmental responses to this abrupt
34 climatic transition. New grain-specific glass analysis (EMPA and LA-ICP-MS) of the type
35 locality Y-3 tephra recorded in the Ionian Sea confirms its origin from Campi Flegrei (CF) but
36 reveals that it is compositionally different from the previously suggested proximal equivalent
37 the VRa eruptive unit (Verdolino Valley, CF). Consequently, the $^{40}\text{Ar}/^{39}\text{Ar}$ age of the VRa
38 should not be exported distally to the Y-3 tephra. Instead, we propose a new robust age for
39 the Y-3 tephra following its identification in the Tenaghi Philippon sedimentary record, NE
40 Greece. A Bayesian-based ^{14}C age model from Tenaghi Philippon provides a distal age of
41 28,680-29,420 cal yrs BP for the Y-3 tephra. The identification of this tephra in NE Greece
42 markedly extends its known eastern dispersal. Whilst its stratigraphic position falls within the
43 latter part of a period of low tree pollen percentages related to dry stadial conditions. This
44 new age and environmental context suggest that this marker postdates the onset of HS3 in
45 the eastern Mediterranean region by ~2,300 years.

46 **Keywords**

47 Y-3 tephra, Ionian Sea, Mediterranean tephrochronology, Tenaghi Philippon, Stadial
48 conditions, Bayesian age modelling.

49 **1. Introduction**

50 Northern hemisphere palaeoenvironmental archives indicate that the last glacial period was
51 punctuated by abrupt and high-amplitude climatic variability (Dansgaard et al., 1993; NGRIP
52 Members, 2004; Rasmussen et al., 2006), a variability that is observed in Mediterranean
53 climate archives (Allen et al., 1999; Sanchez Goni et al., 2000; Tzedakis et al., 2002, 2004;
54 Kotthoff et al., 2011; Müller et al., 2011). However detailed assessment of the spatial and

55 temporal patterning of this extreme climate variability, even at a centennial resolution, is
56 limited by the precision and accuracy of the available chronological information. Radiocarbon
57 dating (^{14}C) remains the most frequently adopted geochronological tool for the last 50 ka BP,
58 but its precision is often insufficient to assess the exact timing of climatic change between
59 different archives and is further complicated by inherent uncertainties associated with
60 temporal variations in marine reservoir offsets (Siani et al., 2001).

61 Volcanic ash (< 2mm) or tephra associated with explosive volcanism can be used to
62 synchronise palaeoenvironmental archives (i.e. tephrostratigraphy) owing to its widespread
63 and synchronous deposition. Furthermore, where the age of a tephra can be determined it
64 provides a chronological marker (i.e., tephrochronology). The high frequency of explosive
65 volcanic activity in the Mediterranean region during the Late Quaternary has made tephra
66 layers particularly powerful chronological tools with marine (e.g., Keller et al., 1978; Vinci,
67 1985; Paterne et al., 1986, 1988, 2008; Calanchi et al., 1996; Vezzoli, 1991; Hardiman 1999;
68 Aksu et al., 2008; Albert et al., 2012) and terrestrial (e.g., Ramrath et al., 1999; Wulf et al.,
69 2004, 2008; Margari et al., 2007; Sulpizio et al., 2010; Wagner et al., 2008; Vogel et al.,
70 2010) archives. More recent cryptotephra (non-visible) studies have increased the
71 geographic range of many known tephra markers and have also presented a number of new
72 isochrons (Siani et al., 2004; Lowe et al., 2007; Bourne et al., 2010; Damaschke et al.,
73 2013).

74 One of these layers crucial to the central Mediterranean tephrostratigraphy is the 'Y-3'
75 tephra, a tephra first identified and labelled based on its occurrence in the Last Glacial
76 (climate zone Y) marine sediments of the Ionian Sea (Keller et al., 1978) (Fig. 1a). This K-
77 trachytic tephra was reported stratigraphically above the thicker K-trachytic Y-5 layer and
78 was thought to be associated with major explosive activity from the Campanian region
79 (Keller et al. 1978). The presence of the Y-3 tephra was later confirmed in new Ionian Sea
80 sediment cores from the *Meteor* cruise M25/4 (Keller et al., 1996; Kraml, 1997) (Fig. 1a).
81 The Ionian Sea records can therefore be considered the 'type locality' for the Y-3 distal

82 marker tephra. The thickness of this K-trachytic ash layer in the Ionian Sea records (Table
83 1), over 450 km from Campanian Volcanic Zone (CVZ) (Table 1; Fig.1a), clearly
84 demonstrates the large magnitude of this Campanian eruption. Stratigraphically this tephra
85 layer is very important due to its close association with the marine isotope stage 3/2
86 transition or Heinrich Stadial 3 (HS3) (Negri et al. 1999). Subsequently this layer has been
87 readily identified in numerous other Mediterranean archives and thus offers a crucial central
88 Mediterranean regional marker layer (Fig 1a) (Table 1; Munno and Petrosino, 2004, 2007;
89 Wulf et al., 2004; Wagner et al., 2008; Zanchetta et al., 2008; Bourne et al., 2010; Caron et
90 al., 2010; Vogel et al., 2010; Damaschke et al., 2013). The timing of this eruption and its
91 widespread dispersal means that it offers significant potential to precisely synchronise
92 archives, enabling the assessment of spatial leads and/or lags associated with an important
93 environmental transition.

94 However, complexity still surrounds the use of the Y-3 tephra as a precise stratigraphic and
95 chronostratigraphic marker. The absence of detailed glass chemistry from the type locality Y-
96 3 tephra means that existing distal correlations have not been subject to the necessary
97 levels of geochemical validation. For this reason it is difficult to know which of these distal
98 ages should be adopted for this marker tephra (Table 1). Determining the precise proximal
99 counterpart of this tephra also presents a further uncertainty. Whilst it is generally regarded
100 that the Y-3 tephra originates from an eruption within the Campi Flegrei (CF) caldera,
101 southern Italy (Fig. 1) (Zanchetta et al., 2008), determining the proximal equivalent remains
102 challenging due to more recent activity and limited exposure in this heavily developed
103 region. Tephrostratigraphic investigations on deposits outside the caldera suggested that the
104 SMP1-e (Santa Maria di Pozzano 1-e) ignimbrite deposits that are exposed along Sorrentine
105 Peninsula (Fig. 1b) were the medial equivalent of the Y-3 tephra (Sulpizio et al., 2003; Di
106 Vito et al., 2008). These deposits show similar stratigraphic, lithological, compositional, and
107 chronological constraints to the distal marker (Table 1). Charcoal material from the palaeosol
108 directly beneath the SMP1-e ignimbrite unit is dated to 29,390-30,720 cal yrs BP, which is

109 older than the interpolated sapropel age of the Y-3 distal marine marker in the Ionian Sea
110 (25.3 ± 3 ka; Kraml, 1997) (Table 1). These SMP1-e tephra deposits, in turn, are correlated
111 to the 4.5 m thick intra-caldera surge and fall deposits of the VRa eruptive unit outcropping in
112 the Verdolino Valley (VR) (Di Vito et al., 2008). The VRa unit represents a single eruption
113 within the Tufi Biancastri stratigraphy (Orsi et al., 1996) - a series of CF eruptions that
114 occurred between the two caldera forming eruptions, the Campanian Ignimbrite (CI)/Y-5
115 (39.28 ± 0.11 $^{40}\text{Ar}/^{39}\text{Ar}$ ka, De Vivo et al., 2001) and Neapolitan Yellow Tuff (NYT)/C-2
116 (14,320-13,900 cal yrs BP; Blockley et al., 2008a). Pappalardo et al. (1999) dated the VRa to
117 30.3 ± 0.2 (1σ) ka ($^{40}\text{Ar}/^{39}\text{Ar}$) which is consistent with the calibrated SMP1-e age (Table 1).
118 Consequently, Zanchetta et al. (2008) suggested that the best age estimate for the Y-3
119 tephra was ca. 30-31 cal ka BP.

120 Here we present the first shard-specific major, minor and trace element data for Y-3 glass
121 shards from its type locality in the Ionian Sea (Keller et al., 1978; Kraml, 1997) to offer a
122 definitive geochemical reference for this tephra. This data is used here to: (1) test proximal
123 links to Campi Flegrei using glass data from the Tufi Biancastri units presented by Tomlinson
124 et al. (2012); (2) assess links to medial-distal extra-caldera tephra deposits recorded on the
125 Sorrentine Peninsula and within the Campanian Plain (i.e., SMP1-e tephra; Di Vito et al.,
126 2008); and (3) use the diagnostic glass geochemistry of the type locality Y-3 to verify existing
127 (see Table 1) and new distal-distal correlations that underpin the synchronisation of archives
128 throughout the Mediterranean region. This will help verify the known dispersal of the Y-3
129 tephra and improve the chronological and environmental constraints placed upon this
130 important tephrostratigraphic marker.

131 **2. Materials**

132 In this section we outline the tephra samples that have been subject to new detailed shard-
133 specific major, minor and trace element geochemical analysis within this study.

134 **2.1 Distal tephra samples**

135 2.1.1 Y-3 *Ionian Sea (M25/4-12)*

136 Reported here is the Y-3 tephra from the Ionian Sea core M25/4-12 (Fig. 1a), a 1 cm thick
137 yellow-grey visible layer that occurs at 117.5-118.5 cm below the sea floor. Core M25/4-12
138 was retrieved with a piston corer from the Calabrian Rise 37°57'98"N; 18°11'04"E in the
139 Central Ionian Sea at a 2473 m water depth *Meteor* cruise M25/4 in 1993 (Keller et al., 1996,
140 Kraml, 1997). The tephra sits close to the MIS 2/3 transition in the oxygen isotope
141 stratigraphy of the core (Negri et al., 1999). The glass shards are typically clear, with
142 occasional brown shards, and these are all between 100-200 µm in size (major axis). Shards
143 comprise of two main morphologies: (1) highly vesicular, tubular shards and; (2) less
144 vesiculated, blocky/angular shards (Fig. 2 a-f). Phenocrysts of sanidine, apatite and biotite
145 are also observed within this tephra layer.

146 2.1.2 TP 9.70 (*TP-2005*), *Tenaghi Philippon, NE Greece*

147 TP 9.70 is a cryptotephra layer reported and geochemically characterised here for the first
148 time. It was identified within the TP-2005 core from the terrestrial site of Tenaghi Philippon
149 (TP), NE Greece (Fig. 1a; see Pross et al., 2007, 2009, and Müller et al., 2011 for details on
150 the site and core). This cryptotephra was detected and extracted following the procedures
151 outlined in Blockley et al. (2005). Multiple peaks in cryptotephra concentrations were
152 detected over a 60 cm interval, the largest peak in shard concentrations is found at the base
153 of this interval (9.70 m) where concentrations were as high as 2060 shards/per gram of dry
154 sediment. At 9.40 m and 9.10 m shard concentrations are lower, with 1030 and 56 shards
155 per gram of dry sediment, respectively. For this reason, 9.70 m is defined as the depth of
156 tephra deposition. Geochemical analysis of the glass shards throughout the 60 cm confirms
157 that they are from the same eruption (supplementary information) and this indicates the
158 upwards reworking of tephra within the peat sequence (Hardiman, 2012). TP 9.70 glass
159 shards were < 80 µm (long axis) and comprised of two main morphologies; (1) highly
160 vesicular, tubular shards and; (2) less vesiculated, more blocky and angular shards.

161 **2.1.3 PRAD 1332 (PRAD 1-2)**

162 The PRAD 1332 cryptotephra was previously reported by Bourne et al. (2010) at a depth of
163 1332 cm in core PRAD 1-2 (Fig. 1a). Major element glass data was presented in Bourne et
164 al. (2010) and here we present shard-specific trace element data. PRAD 1-2 was recovered
165 from the western and upper flank of the Mid-Adriatic deep (42°40.34.7826'N;
166 14°46.13.5565'E) at a water depth of 185.5 m (Bourne et al., 2010). Further details relating
167 to this tephra layer are presented in Table 1. The tephra layer resides stratigraphically just
168 above the MIS 2/3 transition in the core's oxygen isotope stratigraphy (Piva et al., 2008).

169 **2.2 Medial (extra-caldera) tephra samples**

170 The SMP1-e extra-caldera tephra deposits outlined by Sulpizio et al. (2003) and Di Vito et al.
171 (2008) are subject to new shard-specific major, minor and trace element characterisation in
172 this study. Pumices from the SMP1-e type locality at Santa Maria di Pozzano (SMP), 30 km
173 south-east of CF caldera along the Sorrentine Peninsula, are re-investigated (Fig. 1b). This
174 deposit comprises of an ash unit with sparse light grey aphyric pumice lapilli (ZS 98262).
175 These have been interpreted as pyroclastic density current deposits (Di Vito et al., 2008)
176 (Table 1). The CE1 pumice deposits from Cervino, 32 km north-east of the CF caldera (Fig.
177 1b), are also re-analysed. These tephra comprise two, well sorted beds of light grey, aphyric
178 pumice separated by an ash unit (lower ZS 2506; upper ZS 2507). Di Vito et al., (2008)
179 suggest that these deposits were the fall component associated with the SMP1-e eruption.

180 **3. Methods**

181 The visible Ionian Sea Y-3 ash (M25/4-12) was washed, dried and handpicked under a light
182 microscope. Both cryptotephra layers were identified and extracted following the procedures
183 outlined in Blockley et al. (2005). Distal tephra shards and medial pumices were mounted in
184 Struers Epofix epoxy resin. These resin stubs were sectioned, polished and carbon coated
185 for analysis. Scanning electron and transmitted light microscopy was conducted to map the

186 stubs and identify individual clasts to ensure the coupling of major and trace element
187 analysis to a single grain.

188 **3.1 Analytical methods**

189 All new geochemical data was generated from analysis of individual juvenile clasts (volcanic
190 glass shards or pumice). Major and minor element glass data was generated using a
191 wavelength-dispersive JEOL 8600 electron micro-probe (EMP) at the Research Laboratory
192 for Archaeology and the History of Art, University of Oxford. Operating conditions are the
193 same as those used in Smith et al. (2011) and are presented along with secondary
194 standards in the supplementary information. The majority of the trace element glass data
195 was generated using Laser Ablation Inductively Coupled Plasma Mass Spectrometry (LA-
196 ICP-MS). Analyses were performed using an Agilent 7500es ICP-MS coupled to Resonetics
197 193nm ArF excimer laser ablation in the Department of Earth Sciences, Royal Holloway,
198 University of London. Operating conditions are the same as those in Tomlinson et al. (2010)
199 and are presented along with secondary standards in the supplementary information.
200 Secondary standards analysed during both EMP and LA-ICP-MS runs were from the Max
201 Plank institute (MPI-DING suite; Jochum et al., 2006). Shards that were too small to be
202 analysed by LA-ICP-MS (< 20µm spots) were analysed using Secondary Ion Mass
203 Spectrometry (SIMS) on a Cameca IMS 4f ion microprobe at the Istituto di Geoscienze e
204 Georisorse (IGG), Pavia (Italy). The operating conditions used are the same as those in
205 Schiano et al. (2001; 2004) and are presented along with the secondary standards in the
206 supplementary information.

207 **3.2 Bayesian Age-Depth modelling of TP-2005**

208 The age-depth model for Tenaghi Philippon was constructed using Bayesian deposition
209 modelling (e.g. Blockley et al., 2008b) and was undertaken in OxCal version 4.2 (Bronk
210 Ramsey, 2001, 2009) using the internationally agreed IntCal13 calibration curve (Remier et

211 al., 2013)¹. The age model was developed using the '*P_Sequence*' function in OxCal with
212 '*Boundary*' functions placed at changes in lithology in the TP-2005 stratigraphy (Müller et al.,
213 2011). The final age model comprises 20 radiocarbon dates (Müller et al., 2011), and
214 includes a newly modelled proximal age for the Cape Riva tephra (Santorini) of 21,890-
215 22,420 yrs BP (Lee et al., 2013; remodelled using IntCal13) and an ⁴⁰Ar/³⁹Ar age for the
216 Campanian Ignimbrite (39,280 ± 110 yrs BP, De Vivo et al., 2001). Both of these tephras
217 form visible layers within the TP-2005 sequence labelled (TP 7.61 and TP 12.87
218 respectively; Müller et al., 2011). In order to find an optimal '*K value*' (a poisson constraining
219 parameter) the model was initially run using a low *K value* of 0.001 (as recommended by
220 Bronk Ramsey, 2008 and Blockley et al., 2008) and slowly increased until the '*Agreement*
221 *Index*' (*A_{Overall}* and *A_{Model}*) was no lower than 60% (Bronk Ramsey, 2008). No dates had
222 to be removed from the model and the '*Date*' function was used to form realistic age
223 estimates of undated horizons.

224 **4. Results**

225 Representative glass analyses from tephra units analysed in this study are given in Table 2.
226 Full glass data sets for individual tephra deposits are presented in the supplementary
227 material. Table 3 presents diagnostic concentrations and ratios for the fingerprinting of
228 individual tephra deposits. All these new results have also been integrated, where possible,
229 with existing published glass data for proposed Y-3 tephra correlatives (Table 1).

230 **4.1 Y-3 tephra, Ionian Sea – Proximal source**

231 The Ionian Sea Y-3 (M25/4-12) glasses are trachytic in composition, with some less evolved
232 glasses falling close to the phonolite/trachyte boundary (Fig.3a). The K₂O (8.2-10.4 wt. %)
233 contents are significantly higher than the Na₂O (2.9-4.9 wt.%) contents, which is more
234 consistent with a Campi Flegrei (CF) or Somma-Vesuvius source, rather than Ischia, where

¹ Note all radiocarbon ages presented here (cal yrs BP) have been calibrated using the IntCal13 or Marine13 internationally accepted calibration curve (Reimer et al., 2013) at 2 σ unless otherwise stated. Year 0 is 1950.

235 glasses show $\text{Na}_2\text{O} \geq \text{K}_2\text{O}$ (Fig. 4a-b). CaO concentrations at a given MgO clearly illustrates
236 that the Ionian Sea Y-3 derives from Campi Flegrei and not Somma-Vesuvius (Fig. 4c).
237 Major (i.e., 2.7-3.8 wt.% FeO and 2.0-2.8 wt.% CaO) and trace (i.e., 172-383 ppm Zr and 28-
238 56 ppm Nb) element glass compositions of the Ionian Sea Y-3 tephra are heterogeneous
239 (Fig. 5), with two distinct end-members and a few analyses plotting in-between (Fig. 5).

240 Two clear K-trachytic end-members are recognised within the Ionian Sea Y-3 glasses and
241 are defined herein as; (1) the higher silica (> 62 wt. % SiO_2) and; (2) the lower silica (< 62
242 wt. % SiO_2) end-members. Zr/Sr ratios, reflective of magma evolution, clearly distinguish
243 these two end-members (Table 3). The higher silica glasses show far greater levels of
244 incompatible trace element enrichment (i.e., Th, Zr, Nb) than the lower silica glasses (Fig. 6;
245 Table 2). Light Rare Earth Element (LREE) enrichment relative to the Heavy Rare Earth
246 elements (HREE) differs between the two end-members (i.e., La/Yb), with the lower silica
247 glasses displaying higher values (Table 3). Using increasing Th as a fractionation index V,
248 Sr, Ba and Eu all show decreasing concentrations between the two end-members. The lower
249 Sr, Ba and Eu concentrations in the higher silica glasses are likely to reflect greater K-
250 feldspar (Sanidine) fractionation. Incompatible trace element ratios also differ between the
251 two end-members (Table 3). Nb/Th ratios in the higher silica Y-3 glasses confirm the
252 association of the tephra layer with the Tufi Biancastri/NYT series of CF deposits (Fig. 4d)
253 (Tomlinson et al., 2012). Furthermore, vanadium concentrations in all the Y-3 glasses are
254 consistently more elevated than observed in the Pre-CI/CI series glasses and again are
255 consistent with the tephra being associated with the Tufi Biancastri/NYT series. The levels of
256 incompatible trace element enrichment in the lower silica end-member glasses of the Y-3
257 tephra are lower than any currently characterised in the Tufi Biancastri/NYT series glasses
258 (Fig. 4d) (Tomlinson et al., 2010).

259 The VRa eruptive unit is compositionally bimodal, like the Ionian Sea Y-3 tephra, with both a
260 high and low silica trachyte end-member (Fig. 3). Both the major and trace element
261 concentrations of the higher silica end-member of the VRa are the same as the high-silica

262 trachytic glasses of the Ionian Sea Y-3 tephra (Fig. 3, 6; Table 3). However, this high silica
263 component of the Y-3 tephra is also compositionally similar to other stratigraphically younger
264 Tufi Biancastri eruptive deposits (namely the VRb and PRa units; Fig. 3, 5). Notwithstanding,
265 the major element concentrations of the lower silica end-member of the Ionian Sea Y-3
266 tephra are significantly different to that of the VRa tephra (Fig. 3). The lower silica end-
267 member of the Y-3 glasses have higher SiO₂, K₂O and lower TiO₂, FeO, CaO, MgO and
268 Na₂O contents than the lower silica end-member of the VRa glasses (Fig. 3). Thus, the lower
269 silica Y-3 glasses might appear intermediate in composition between the two end-members
270 of the VRa glasses (Fig. 3). However, trace element variability between the two end-
271 members of the VRa glasses is more restricted than that of the Ionian Sea Y-3 tephra (Fig.
272 5), thus inconsistent with their respective major element variability. Vanadium, Ba and Sr
273 concentrations in the lower silica K-trachytic VRa glasses are higher than those in the lower
274 silica Y-3 glasses. Plotting these elements against an incompatible element such as Th
275 clearly illustrates that the Ionian Sea Y-3 tephra and the proximal VRa eruptive unit glasses
276 lie on separate evolutionary trends (Fig. 6) and therefore are associated with different
277 eruptions.

278 **4.2 Y-3 Ionian Sea – Medial correlatives**

279 The SMP1-e ignimbrite deposits (ZS 98 262) from the Sorrentine Peninsula have
280 homogenous (i.e., 62.1 ± 0.5 wt.% SiO₂; 8.5 ± 0.2 wt.% K₂O) glass compositions and are
281 classified as K-trachytes (Fig. 3a). The trace element compositions of the glasses are
282 equally homogenous (i.e., 54 ± 3 ppm Nb; 29 ± 5 ppm Th; Fig. 5). The major and trace
283 element compositions of these SMP1-e glasses are indistinguishable from the higher silica
284 K-trachyte end-member of the Ionian Sea Y-3 tephra (Fig. 3; 5-6). Indeed, they share
285 consistent incompatible trace element ratios and levels of LREE enrichment (Table 3).

286 The CE1 pumice fall beds from Cervino have glass compositions that are also
287 homogeneous. The lower (ZS 2506) pumice fall glasses are fractionally more evolved (i.e.,

288 58.1 ± 0.4 wt.% SiO_2) than the upper (ZS 2507) pumice fall glasses (57.7 ± 0.5 wt. % SiO_2).
289 Both fall deposits classify as phonolites but lie close to the trachyte boundary (Fig. 3a).
290 These phonolitic deposits have glass compositions that are clearly distinguishable from both
291 end-members of the Ionian Sea Y-3 tephra and also the SMP1-e ignimbrite deposits. The
292 CE1 glasses have major element compositions that partially overlap with the lower silica
293 glasses of the proximal VRa eruptive unit (Fig. 3), but the trace element compositions of the
294 lower (i.e., 27 ± 2 ppm Th) and the upper (i.e., 25 ± 2 ppm Th) CE1 glasses are clearly
295 different from those of the VRa glasses (Fig. 6; Table 3). Most noticeable is that the CE1
296 glasses show more elevated Th concentrations than the lower silica VRa glasses (21 ± 2
297 ppm Th). The CE1 glasses also have incompatible trace element ratios that are offset from
298 the currently available Tufi Biancastri/NYT series glasses (Tomlinson et al., 2012) (Fig. 6d;
299 Table 3).

300 **4.3 Y-3 Ionian Sea - distal correlatives**

301 Glass shards from cryptotephra TP 9.70 recorded at TP have a heterogeneous major
302 element composition (60.1 to 63.1 wt. % SiO_2 ; 8.1 - 10.0 wt. % K_2O) and show the same
303 major element variability as the Ionian Sea Y-3 tephra (Fig. 3, 5). Proposed Y-3 tephra
304 correlatives (Table 1), MD90-917 920 (S. Adriatic; Zanchetta et al., 2008), OT0702-4/JO-188
305 (Lake Ohrid; Balkans; Vogel et al., 2010; Caron et al., 2010) and PT9015-05 (Lake Prespa,
306 Balkans; Damaschke et al., 2013) all show major element variability that it is largely
307 consistent with that of the Ionian Sea Y-3 and TP 9.70 Tenaghi Philippon (TP) layers (Fig. 3,
308 5). These correlatives appear to comprise predominantly of intermediate glass compositions
309 and extend towards either the higher or lower silica end-members of the Ionian Sea Y-3
310 tephra (Fig. 3, 5). The TM-15 glasses from Lago Grande di Monticchio (LGdM; Wulf et al.,
311 2004; Tomlinson et al., 2012), the S-19 tephra from San Gregorio Magno (SGM; Munno and
312 Petrosino, 2004, 2007) basin, and the Tyrrhenian Sea ash layers (C-7, B2 and A2; Paterne
313 et al., 1988, Buccheri et al., 2002a, 200b; Munno and Petrosino, 2004) (Table 1) all appear

314 to be restricted to the higher silica trachytic end-member of the Ionian Sea Y-3 tephra (Fig.
315 3a).

316 The trace element glass compositions of cryptotephra TP 9.70 are heterogeneous (168-420
317 ppm Zr; 28-56 ppm Nb) with a compositional range consistent with those of the Ionian Sea
318 Y-3 tephra (Fig. 5-6). The TP 9.70 glasses present compositions that are consistent with
319 both end-members of the Ionian Sea Y-3 and this is reflected in their comparable
320 incompatible trace element ratios (Table 3). Crucially, the TP 9.70 glasses lie upon the same
321 evolution trends as the Ionian Sea Y-3 tephra, best demonstrated by V, Sr and Ba
322 concentrations relative to Th (Fig. 6a-c). TP 9.70 glasses are dominated by trace element
323 compositions that intermediate between to the two end-members of the Ionian Sea Y-3
324 tephra. TM-15 (LGdM) glasses do not display as much trace element heterogeneity as either
325 the Ionian Sea Y-3 tephra or cryptotephra TP 9.70 and this is consistent with the absence of
326 a lower silica K-trachytic end-member at a major element level (Fig. 3). However using V, Sr
327 and Ba plotted against Th it is clear that the TM-15 glasses fall upon the same diagnostic
328 trends as the Ionian Sea Y-3 tephra, which is different from the VRa eruptive unit (Fig. 6).
329 The presence of intermediate compositions also demonstrates particularly good
330 geochemical agreement with the cryptotephra TP 9.70 (Fig. 6).

331 The major element glass compositions of the bimodal central Adriatic tephra layer PRAD
332 1332 do not correspond with either end-member of the Ionian Sea Y-3 tephra (Fig. 3). The
333 higher silica K-trachyte component of the PRAD 1332 glasses have noticeably lower K_2O
334 and slightly higher CaO than the Ionian Sea Y-3 tephra glasses at a similar silica
335 concentration (Fig. 3c; 5c). The lower silica trachyte component of the PRAD 1332 layer
336 instead corresponds to the lower silica trachytic end-member of the VRa eruptive unit (Fig.
337 3). At a trace element level the bi-modality of the PRAD 1332 tephra is best observed by the
338 clear differences in Ba, Sr and V concentrations (Table 3), the lower silica component shows
339 higher concentrations of these elements compared to the higher silica component (Fig. 6)
340 and is consequently less evolved (lower Zr/Sr ratio; Table 3). As with their major element

341 compositions, the trace element concentrations and incompatible element ratios of the lower
342 silica K-trachyte glasses in the PRAD 1332 layer are similar to the compositional field of the
343 lower silica VRa glasses (Fig. 6; Table 3). The higher silica PRAD 1332 glasses fall on the
344 evolutionary trend between the two end-members of the VRa glasses (Fig. 6).

345 **5. Discussion**

346 **5.1 Y-3 tephra correlations**

347 The new glass data presented here for the type locality Ionian Sea Y-3 tephra confirms that
348 the tephra is associated with a Campi Flegrei (CF) eruption and is consistent with an event
349 from the Tufi Biancastri/NYT series (cf. Tomlinson et al., 2012) (Fig. 4d). This information
350 coupled with its stratigraphic position (above the Y-5/ Campanian Ignimbrite) and
351 chronological constraints (25.3 ± 3 ka; Table 1) are evidence for the Ionian Sea Y-3 tephra
352 being erupted between the caldera forming CI and NYT eruptions (an eruption in the Tufi
353 Biancastri sequence). The high silica K-trachyte end-member of the Ionian Sea Y-3 clearly
354 verifies this affinity to the Tufi Biancastri Series glasses (Fig. 3, 6). Unfortunately, a high
355 silica (61.5-62.5 wt. %), K-trachyte glass chemistry is repeatedly erupted through time as it is
356 recorded in successive units within the Tufi Biancastri stratigraphy (i.e., VRa, VRb and PRa;
357 Fig. 3). Consequently, this means that this geochemical component alone is not diagnostic
358 (Tomlinson et al., 2012) or useful for precisely establishing the proximal equivalent of this
359 distal tephra. The eruption of repeat major element glass compositions at CF is not unique to
360 the Tufi Biancastri deposits, many Holocene eruptive deposits are also compositionally
361 indistinguishable (i.e., Smith et al., 2011).

362 It is the presence of the full compositional variability and in particular the identification of the
363 lower silica (ca. 60-61.5 wt.%) end-member of the Ionian Sea Y-3 tephra that is diagnostic of
364 this marker layer (Fig. 3, 5-6). However, trace element concentrations confirm that the Ionian
365 Sea Y-3 and the VRa are not from the same eruption, as the two tephra lie upon separate

366 evolutionary trends (Fig. 6). Consequently, the $^{40}\text{Ar}/^{39}\text{Ar}$ of the VRa should not be exported
367 distally to the Y-3 tephra.

368 The absence of a precise proximal equivalent of the Ionian Sea Y-3 tephra illustrates a need
369 for further grain-specific geochemical characterisation of more proximal Tufi Biancastri
370 eruptive units. However, the potential for identifying the proximal equivalent of the distal
371 tephra at CF is likely to be restricted by the complexity of the proximal volcanic stratigraphy,
372 where often only limited exposure is available, particularly given that subsequent caldera
373 collapse (NYT) has destroyed and/or buried many of the older pyroclastic units (Di Vito et al.,
374 2008). The absence of a proximal age places further emphasis on establishing precise
375 medial and distal tephra correlations in order to resolve the age of this distal marker.

376 At extra-caldera localities, the SMP1-e ignimbritic tephra recorded on the Sorrentine
377 Peninsula (Fig.1b) does not present the full diagnostic compositional heterogeneity of the
378 Ionian Sea Y-3 tephra (Fig. 3, 4-5). Only the high silica trachytic end-member of the Ionian
379 Sea Y-3 is identified and, given that this composition is repeatedly erupted in this timeframe,
380 a correlation with the Ionian Sea Y-3 marker remains inconclusive based on glass chemistry
381 alone. The proposed fall component of the SMP1-e eruption, the CE1 tephra deposits, NE of
382 Campi Flegrei (Fig. 1b), are not medial equivalent of the Y-3 tephra or the VRa eruptive unit
383 (Fig. 6). Furthermore, the new glass data raises doubt over the stratigraphic correlation of
384 both the SMP1-e (ignimbritic) and CE1 (fall) units under a single SMP1-e eruptive deposit
385 (Di Vito et al., 2008). These units have different incompatible trace element ratios (Table 3)
386 and this, combined with the absence of intermediate compositions between their respective
387 compositions, means that it is difficult to envisage them as being related to the same CF
388 eruption (Fig. 5). Consequently, the CE1 tephra fall deposits should no longer be
389 chronologically constrained by the age of the SMP1-e ignimbrite deposits (Table 1). This
390 interpretation emphasises the frequency of explosive activity at CF, and the difficulty of
391 stratigraphically correlating proximal and medial tephra deposits based only on limited
392 exposure.

393 Currently, neither the extra- or intra-caldera deposits analysed fully satisfy the diagnostic
394 compositional variability of the Ionian Sea Y-3 tephra (Fig. 3, 5-6). The high silica K-trachytic
395 end-member is characteristic of most post CI/pre-NYT CF eruptions and thus this
396 component is not diagnostic. Fortunately, the compositional range of the Ionian Sea Y-3
397 tephra, in particular the presence of the lower silica trachytic end-member, is diagnostic of
398 this marker layer. Consequently, it is important to use the full diagnostic geochemical
399 signature of the Ionian Sea Y-3 tephra when attempting to validate Y-3 tephra correlations.

400 The Ionian Sea Y-3 glass data unequivocally confirms an eastern ash dispersal associated
401 with this CF eruption (Fig. 3, 5-6). The cryptotephra TP 9.70 recorded at TP, NE Greece,
402 corresponds precisely to the Ionian Sea Y-3 tephra (Fig. 3, 5-6). This correlation extends the
403 known eastern dispersal of the Y-3 tephra to over 800 km from CF (Fig. 7) and would also
404 imply that the area effected by ash deposition from this eruption is greater than 550 000 km²,
405 which was previously suggested by Caron et al., (2010). Importantly, the identification of the
406 Y-3 at TP demonstrates the potential of this marker horizon to integrate the Italian and
407 Aegean tephrostratigraphic records. This is significant as currently only two Italian tephra
408 markers have been integrated within the Aegean tephrostratigraphic record, the P-11
409 (Pantelleria) and Campanian Ignimbrite/Y-5 layers (Sulpizio et al., 2010). Trace element
410 glass chemistry verifies the correlation of TM-15 at Lago Grande di Monticchio (LGdM) to the
411 Ionian Sea Y-3 tephra (Wulf et al., 2004). This eastern dispersal of the Y-3 eruption appears
412 to be biased towards the intermediate and most evolved glass compositions (Fig. 6), while
413 the southern dispersal (Ionian Sea) is dominated by the least and most evolved
414 compositional end-members (Fig. 6).

415 Other proposed eastern occurrences of the Y-3 tephra recorded in the southern Adriatic
416 (MD90-917 920; Zanchetta et al., 2008) and the Balkans (Fig. 1b) (Lake Ohrid; Wagner et
417 al., 2008; Caron et al., 2010; Vogel et al., 2010; and Prespa; Damaschke et al., 2013) are
418 supported by our new type locality Y-3 glass data. These tephra layers have major element
419 glass compositions that match the diagnostic lower silica end-member of the Y-3 tephra, and

420 also show similar compositional variability (Fig. 3, 5). This verification indicates that the Y-3
421 tephra is a very important isochron for synchronising crucial terrestrial Mediterranean
422 palaeoenvironmental records LGdM (Brauer et al., 2007), Lake Ohrid (Belmerchi et al.,
423 2009; Lezine et al., 2010) and TP (Fletcher et al., 2011; Müller et al., 2011). This allows the
424 synchronisation of records along an east-west transect extending from southern Italy via the
425 Balkans to north-east Greece.

426 The bi-modal K-trachytic cryptotephra PRAD 1332 recorded in the marine core PRAD 1-2,
427 was previously correlated to TM-15, and by association, the Y-3 tephra and the VRa eruptive
428 unit (Bourne et al., 2010). The VRa $^{40}\text{Ar}/^{39}\text{Ar}$ age was imported to the depth of PRAD 1332
429 (Bourne et al., 2010). Even though major and trace element glass data demonstrates that
430 PRAD 1332 is not a correlative of the Y-3 tephra (Fig. 3, 5-6) the lower silica K-trachytic
431 component of PRAD 1332 geochemically corresponds to the lower silica end-member of the
432 VRa eruption. Consequently, a correlation with the VRa eruption might be argued, and thus
433 the attribution of the $^{40}\text{Ar}/^{39}\text{Ar}$ age by Bourne et al., (2010) may still be considered sensible.
434 Regardless of whether PRAD1332 is the distal equivalent of the VRa eruption, there are
435 important implications associated with its erroneous correlation with the Y-3 tephra. Firstly,
436 the stratigraphic position of PRAD 1332 close to the MIS 2/3 transition in PRAD1-2, is
437 consistent with the position of Y-3 tephra in M25/4-12 (Negri et al., 1999), testifying to a high
438 frequency of explosive activity at CF coinciding with this important environmental transition.
439 Secondly, this reappraisal also currently limits the known north-easterly dispersal of the Y-3
440 tephra (Fig. 7).

441 Given that compositional variability is so diagnostic of the Ionian Sea Y-3 tephra, the
442 absence of shard-specific glass data makes it more difficult to reliably assess other
443 proposed correlations. Validating correlations with the Tyrrhenian Sea ash layers, the C-7
444 (Paterne et al., 1988; Zanchetta et al., 2008) and the A2/B2 (Munno and Petrosino, 2004)
445 tephra deposits is challenging. Along with the terrestrial S-19 tephra, from the San Gregorio
446 Magno Basin (Munno and Petrosino, 2007), all these layers only demonstrate the presence

447 of a higher silica K-trachytic glass component (Fig. 3a). Consequently, this glass data alone
448 merely confirms that these layers derive from a Tufi Biancastri eruption but does not allow us
449 to precisely verify an Y-3 tephra correlation.

450 **5.2 Chronology of the Y-3 tephra**

451 A Bayesian modelled age of 28,680-29,420 cal yrs BP has been obtained for the confirmed
452 Y-3 tephra at TP (Fig 8; = TP 9.70), based upon multiple radiocarbon ages from both above
453 and below the tephra. Dated material includes mollusc shells, wood and bulk peat sediment,
454 with the latter providing numerous ages above and below TP 9.70 (Müller et al., 2011). The
455 'Date' function within OxCal was used to generate a robust age for the precise depth of the
456 peak in shard concentrations (9.70 m), considered representative of the timing of deposition
457 and the eruption (Fig. 9). Confidence in the robustness of the Tenaghi Philippon
458 radiocarbon-based Bayesian age-model can be indirectly assessed via the accuracy of the
459 modelled ages for other, previously correlated visible tephra layers within the TP-2005
460 sequence. This was done by removing the visible tephra ages from the model and
461 comparing the modelled age for the tephra depth with the known published age of the
462 tephra. TP 7.60 is correlated to the Cape Riva, Santorini and TP 12.87 to the Campanian
463 Ignimbrite, CF (Müller et al., 2011). The age-depth model, with only the radiocarbon
464 information included, produced modelled ages of 20,800-22,750 and 37,690-39,910 cal yrs
465 BP, respectively, for these eruptions. This is in good agreement with available proximal ages
466 for the Cape Riva (21,890-22,420 cal yrs BP; Lee et al., 2013; remodelled using IntCal13)
467 and the CI eruption ($^{40}\text{Ar}/^{39}\text{Ar}$ 39,280 \pm 110 yrs BP, De Vivo et al., 2001). Consequently, this
468 demonstrates the integrity of the Tenaghi Philippon radiocarbon ages both above and below
469 the Y-3 tephra.

470 The precise geochemical correlation between TP 9.70 and the Ionian Sea Y-3 tephra allow
471 us to import the modelled age from TP to the Ionian Sea core M25/4-12 at a depth of 119.5
472 cm (b.s.f.). The imported TP 9.70 age (28,680-29,420 cal yrs BP) clearly provides a more

473 precise chronological constraint than the previously (interpolated sapropel) age given to this
474 tephra (Table 1). The TP 9.70 age also shows very good agreement with the with calibrated
475 ^{14}C age from directly above the OT0702-4/Y-3 layer in the Lake Ohrid record (28,780-29,980
476 cal yrs BP; Vogel et al., 2010). At LGdM the varve age of TM-15/Y-3 (25,900-28,620 yrs BP;
477 Wulf et al., 2012) is slightly younger than the TP 9.70/Y-3 calibrated ^{14}C age. Independent
478 ages for other tephras in this part of the LGdM stratigraphy have suggested that the varve
479 chronology presents a slight underestimate of tephra ages (Brauer et al., 2000; Wulf et al.,
480 2012).

481 Bronk Ramsey et al., (submitted) use a Bayesian age-depth model to generate a marine
482 reservoir corrected, integrated age for the A2/B2 Tyrrhenian Sea tephra layers that were
483 previously correlated to the Ionian Sea Y-3 tephra (Munno and Petrosino, 2004). This yields
484 a modelled age of 28,618-29,541 cal yrs BP, is in perfect agreement with the TP 9.70/Y-3
485 age (28,680-29,420 cal yrs BP). Whilst available geochemical data from these two
486 Tyrrhenian Sea layers (A2 and B2) is not sufficient to precisely validate their affinity to the
487 Ionian Sea Y-3 tephra, this chronological agreement offers significant weight to the existing
488 correlation. The ^{14}C age associated with the extra-caldera SMP1-e ignimbrite deposits
489 (29,390-30,720 cal yrs BP) on the Sorrentine Peninsula (Di Vito et al., 2008) is slightly older
490 than the TP 9.70/Y-3 age. This age and the chemistry does not indicate the existing
491 correlation between the SMP1-e ignimbrite deposits and the distal Ionian Sea Y-3 tephra is
492 wrong (Zanchetta et al., 2008), but indicates further verification is required.

493 In terms of determining the absolute ordering of eruptive events at CF, the ages of the distal
494 Y-3 (TP 9.70; 28,680-29,420 cal yrs BP) and the proximal VRa (29,900-30,700 $^{40}\text{Ar}/^{39}\text{Ar}$ yrs
495 BP; Pappalardo et al., 1999) tephra would suggest that the former represents the younger of
496 two closely spaced CF eruptions. In the central Adriatic marine core, PRAD1-2, the ^{14}C ages
497 below the PRAD1332 cryptotephra layer provide a maximum age (Table 1) that would
498 indicate that this eruption is younger than the eruption of the Y-3 tephra. Whilst geochemical
499 evidence might point to a PRAD1332-VRa correlation, if accepted, their respective ages

500 (Table 1) would suggest that dating either proximally or distally is erroneous. The negligible
501 temporal gaps between the independent ages of the respective tephra deposits mean that
502 validating the absolute ordering of these eruptive events will only be fully established through
503 their identification in the same stratigraphic sequence. Owing to resurgent activity, intra-
504 caldera deposits at CF are complex and consequently establishing the stratigraphic
505 superposition of these eruptive events is more likely to be determined in the distal realm.

506 Where the diagnostic chemistry of the Ionian Sea Y-3 tephra underpins distal correlations,
507 consistency between the TP 9.70/Y-3 (28680-29420 cal yrs BP; this study) and OT0702-4/Y-
508 3 (28780-29980 cal yrs BP; Vogel et al., 2010) ages mean that they currently present the
509 most reliable chronological constraints for the Y-3 tephra marker.

510 **5.3 Climastratigraphic context of the Y-3 tephra**

511 The identification of the Y-3 tephra in the high-resolution TP palaeoenvironmental archive
512 provides detailed information on the environmental conditions at the time of the Y-3 eruption.
513 The Y-3 tephra at TP sits within the latter part of a period marked by a major reduction in
514 tree pollen percentages (Fig. 9) that are related to stadial conditions (Müller et al., 2011).
515 The age-depth model at TP suggests that the Y-3 tephra was deposited ~2,300 years after
516 the onset of stadial conditions. Within this overall period of reduced tree pollen percentages
517 there is a small increase in total tree pollen percentages just below the Y-3 tephra which
518 may reflect a brief climatic amelioration (Fig. 9). Comparisons between the TP and LGdM
519 (Allen et al., 1999) palynological records using the Y-3 tephra marker (See Wutke et al.,
520 submitted) reveals that the tephra occurs within the pollen zone correlated to Greenland
521 Stadial 5 (GS-5) in both archives (Fletcher et al., 2011; Müller et al., 2011). The
522 palaeoenvironmental record from Lake Ohrid sequence JO2004 (Lezine et al., 2010) also
523 suggests the Y-3 sits within a period of reduced total tree pollen. Combined evidence from
524 these sites, in particular the very high resolution pollen stratigraphy from TP, suggest that
525 the Y-3 tephra post-dates the onset of HS3 (*sensu* Sanchez Goni and Harrison, 2010)

526 conditions in the Mediterranean region. The TP environmental record, on an independent
527 time scale, suggests that the Y-3 eruption occurred during or after HE3 (Fig. 9), although
528 given the inherent difficulties comparing marine and terrestrial radiocarbon datasets
529 detection of the Y-3 within a high-resolution marine proxy record might be required to resolve
530 this question. The independent dating evidence put forward for the Y-3 tephra herein
531 strongly suggests that this eruption occurred sometime after the onset of Greenland Stadial
532 5 in the INTIMATE event stratigraphy (Blockley et al., 2012) and before the MIS 3/2
533 transition as defined by Svensson et al. (2006, 2008) (Fig. 9).

534 **6. Conclusions**

535 The Ionian Sea Y-3 tephra (M25/4-12) has a heterogeneous K-trachytic chemistry and glass
536 compositions confirm a source from within the Campi Flegrei caldera. The combined major
537 and trace element glass chemistry shows that the Y-3 tephra is not the distal equivalent of
538 the proximal Tufi Biancastri VRa eruptive unit (30.3 ± 0.2 ka BP). None of the medial tephra
539 investigated from extra-caldera localities fully satisfy the compositional range of the Ionian
540 Sea Y-3 tephra and the precise proximal equivalent of this eruption remains unknown. The
541 diagnostic glass chemistry for the Ionian Sea Y-3 tephra identified and enables us to
542 establish precise distal-distal tephra correlations. A correlative of the Ionian Sea Y-3 tephra
543 is also preserved in the Tenaghi Philippon record, NE Greece, which markedly extends the
544 eastern dispersal of this tephra. A Bayesian-based ^{14}C age depth-model for the Tenaghi
545 Philippon provides a robust age of 28,680-29,420 cal yrs BP for this distal marker tephra,
546 which is in agreement with other distal age estimates. Previous work has shown that at
547 Tenaghi Philippon the Y-3 marker tephra occurs in the later stages of a period linked to
548 stadial conditions (Müller et al., 2011). Dating of the Y-3 would suggest that it was erupted
549 after the onset of Greenland Stadial 5 and post-dates the beginning of Heinrich Stadial 3.
550 Detailed geochemical characterisation and independent dating mean this widespread tephra
551 layer offers both a crucial tephrostratigraphic and chronostratigraphic marker associated with
552 an important climatic event.

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828

829 **Figure Captions**

830

831 **Figure 1:** (a) A map of the central and eastern Mediterranean showing the location of the
832 M25/4-12 core from which the type locality Y-3 tephra was investigated. The localities
833 of other archives where proposed 'Y-3' tephra correlatives have been reported (listed
834 in Table 1). The main volcanic centres active during the Last Glacial (CF, Campi
835 Flegrei; SV, Somma-Vesuvius; IS, Ischia, Ae, Aeolian Islands; Et, Mount Etna; Pa,
836 Pantelleria) are marked. (b) The volcanoes in the Neapolitan volcanic region, Campi
837 Flegrei, Somma-Vesuvius and Ischia (adapted from Tomlinson et al., 2012), and the
838 localities of extra-caldera (SMP1-e and CE1) samples.

839

840 **Figure 2:** SEM and light microscope images of the distal and proximal tephras. Y-3 tephra
841 recovered from M25/4-12 in the Ionian Sea (a-f); (a) shard 40D, a 20µm ablation pit
842 is observed left of a vesicle; (b) shard 29C, with two laser ablation pits (34 and 20 µm
843 diameter); (c) shard 50E showing a 34 µm ablation pit; (d) shard 32D, a thin microlite
844 is observed running through the glass shard, which demonstrates the benefit of
845 preliminary SEM investigations prior to LA-ICP-MS analysis; (e) shard 51E; (f) shard
846 42D showing two ablation pits (34 and 20 µm); (g-h) the lower VRa, highly
847 vesiculated white pumices (i) the upper VRa, less vesiculated, darker pumice; (j-l)
848 SMP1-e (ZS98 262), highly vesicular and stretched pumices; (m) lower CE1 (ZS
849 2506) pumice, vesicular and stretched glasses; (n-m) upper CE1 (ZS 2507) pumices
850 that are less vesiculated than the stratigraphically lower CE1 pumices.

851

852 **Figure 3:** Major element bi-plots showing the glass compositions of Ionian Sea Y-3 tephra
853 (M25/4-12) compared to those of proximal intra-caldera Tufi Biancastri (Tomlinson et
854 al., 2012) and extra-caldera deposits (SMP1-e, CE1; this Study) from Campi Flegrei.
855 Glass data from cryptotephra TP 9.70 layer in Tenaghi Philippon, NE Greece (This

856 study). Published glass compositions of distal layers (Table 1) are also presented to
857 assess distal-distal tephra correlations: References are as follows; (1) Tomlinson et
858 al., (2012); (2) Munno and Petrosino, 2007; (3) Bourne et al., (2010)*; (4) Zanchetta
859 et al., (2008); (5) Munno and Petrosino (2004); (6) Paterne et al. (1988), Zanchetta et
860 al., (2008); (7) Vogel et al., (2010); Caron et al., (2010); (9) Damaschke et al., (2013).
861 Error bars represent 2 standard deviations of repeat analyses of the StHs60-G MPI-
862 DING reference glass. (a) TAS classification after Le Bas et al. (1986). *PRAD 1332
863 data was normalised assuming 0.4 wt.% of Cl (Cl not measured by Bourne et al.,
864 2010) for more reliable water-free comparisons.

865

866 **Figure 4:** Major and trace element glass compositions of the Ionian Sea Y-3 tephra (M25/4-
867 12) compared to proximal Campi Flegrei (Tomlinson et al., 2012), Somma-Vesuvius
868 (Tomlinson et al. unpublished¹) and Ischia (Tomlinson et al., submitted) tephra
869 deposits. (a-c) Y-3 glasses compared to the glass compositional fields of the
870 Neapolitan volcanic centres; and (d) Y-3 glass compositions plotted against those of
871 the pre- and post-Campanian Ignimbrite series (Tomlinson et al., 2012). Error bars
872 represent 2 standard deviations of repeat analyses of the StHs60-G reference glass.
873 ¹ Vesuvius glass data used to evaluate potential geochemical links and to generate
874 this compositional field is available on the RESET database
875 (<https://c14.arch.ox.ac.uk/login/login.php?Location=/resetdb/db.php>)

876

877 **Figure 5:** Major and trace element compositional variation of glasses from the Y-3 tephra
878 recorded in the Ionian Sea compared to potential proximal, medial and distal tephra
879 correlatives. The Ionian Sea Y-3 glasses are dominated by two compositional end-
880 members, whilst the Tenaghi Philippon tephra spans the full compositional range.
881 Refer to Figure 3 figure caption for published data references.

882

883 **Figure 6:** Trace element glass compositions of the Ionian Sea Y-3 tephra (M25/4-12)
884 compared to those of proximal intra-caldera Tufi Biancastri (Tomlinson et al., 2012)
885 and extra-caldera deposits (SMP1-e, CE1; this study) from Campi Flegrei. The glass
886 compositions of distal tephra layers TP 9.70, PRAD 1332 (this study), and LGdM TM-
887 15 (Tomlinson et al., 2012). Error bars represents 2 standard deviations of repeat
888 analyses of both the StHs60-8G and ATHO-1 reference glasses.

889

890 **Figure 7:** The revised known dispersal and distribution of the Y-3 tephra.

891

892 **Figure 8:** Posterior probability density function generated for TP 9.70, the Tenaghi Philippon
893 cryptotephra layer. The brackets at the base of the distribution represent the 95.4%
894 and 99.7% probability ranges.

895

896 **Figure 9:** Climatostratigraphic context of the Y-3 tephra at Tenaghi Philippon, NE Greece.
897 Shown from left to right: (a) Marine Isotope Stage (Svensson et al., 2006; 2008); (b)
898 The INTIMATE event stratigraphy (Lowe et al., 2008; Blockley et al., 2012). (c) The
899 NGRIP oxygen isotope record based upon the GGIC05 timescale from Andersen et
900 al. (2006) and Svensson et al. (2006). (d) Calibrated age ranges for Heinrich Event 3
901 from western European records (Bard et al., 2000; de Abreu et al., 2003) (thin line)
902 and calculated ages (Thouveny et al., 2000) (black box). (e) Tenaghi Philippon (TP-
903 2005) total tree pollen percentage curve and the stratigraphic position of the Y-3
904 cryptotephra. (f) The age–depth model (at 95.4 % probability range) and radiocarbon
905 ages on which it is based are also shown (green) alongside the original Müller et al.
906 (2011) chronology (grey green line). Note all data have been adjusted to a yrs BP
907 (1950) timescale.

908

909 **Table Captions**

910

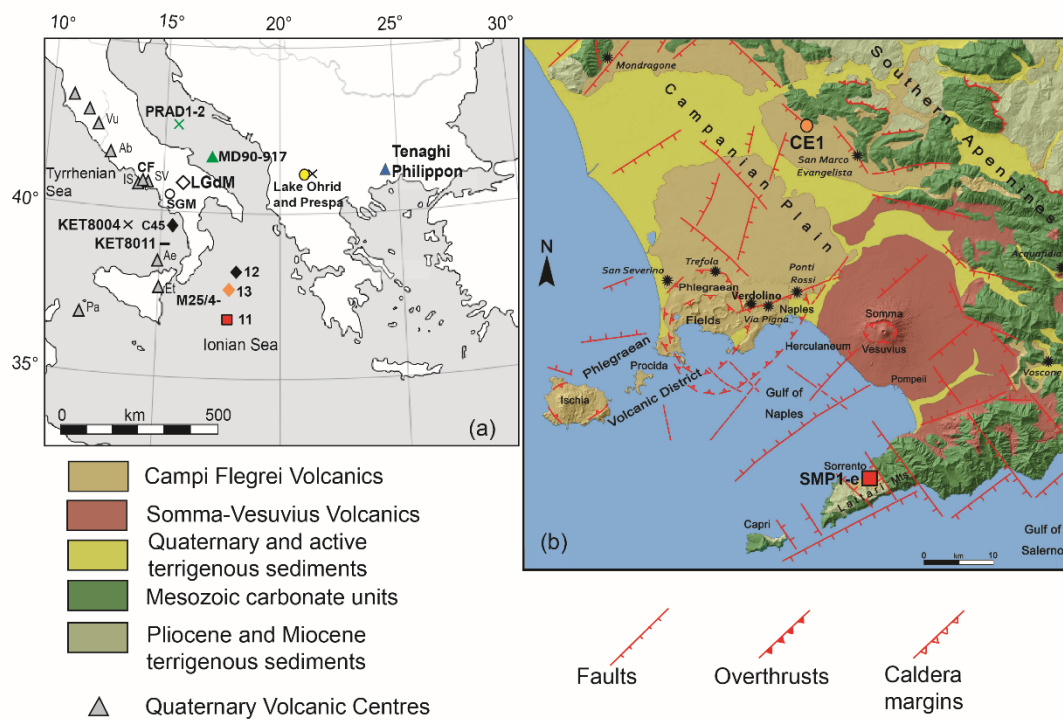
911 **Table 1:** Reported occurrence of the Y-3 tephra and proposed correlatives from across the
912 Central Mediterranean region. Dispersal is given relative to Campi Flegrei caldera.
913 Presented are ages associated with these tephra deposits. Terrestrial and marine
914 radiocarbon ages are have been calibrated using atmospheric and marine data sets
915 respectively incorporated within IntCal13. LGdM, Lago Grande di Monticchio and
916 SGM, San Gregorio Magno basin.

917 **Table 2:** Representative shard-specific major, minor (EMP) and trace element (LA-ICP-MS)
918 glass data from the Ionian Sea Y-3 tephra (M25/4-12), extra-caldera CF deposits
919 (SMP1-e and CE1), and other distal tephra deposits considered as potential
920 correlatives of the Y-3 tephra including TP 9.70 (Tenaghi Philippon, NE Greece) and
921 PRAD 1332 (PRAD 1-2, Central Adriatic). (LA, LA-ICP-MS analyses). A full grain-
922 specific glass data set is presented in the supplementary information. * water-free
923 major element data for PRAD1332 (Bourne et al., 2010) was calculating assuming a
924 0.4 wt.% Cl prior to normalisation.

925 **Table 3:** Diagnostic concentrations and ratios for geochemical fingerprinting the Y-3 tephra,
926 errors are 2 standard deviations. TAS classification is based on (Na₂O + K₂O) versus
927 SiO₂ using Le Bas et al. (1986). TP 9.70 data are representative analysis from each
928 compositional end-member.

929

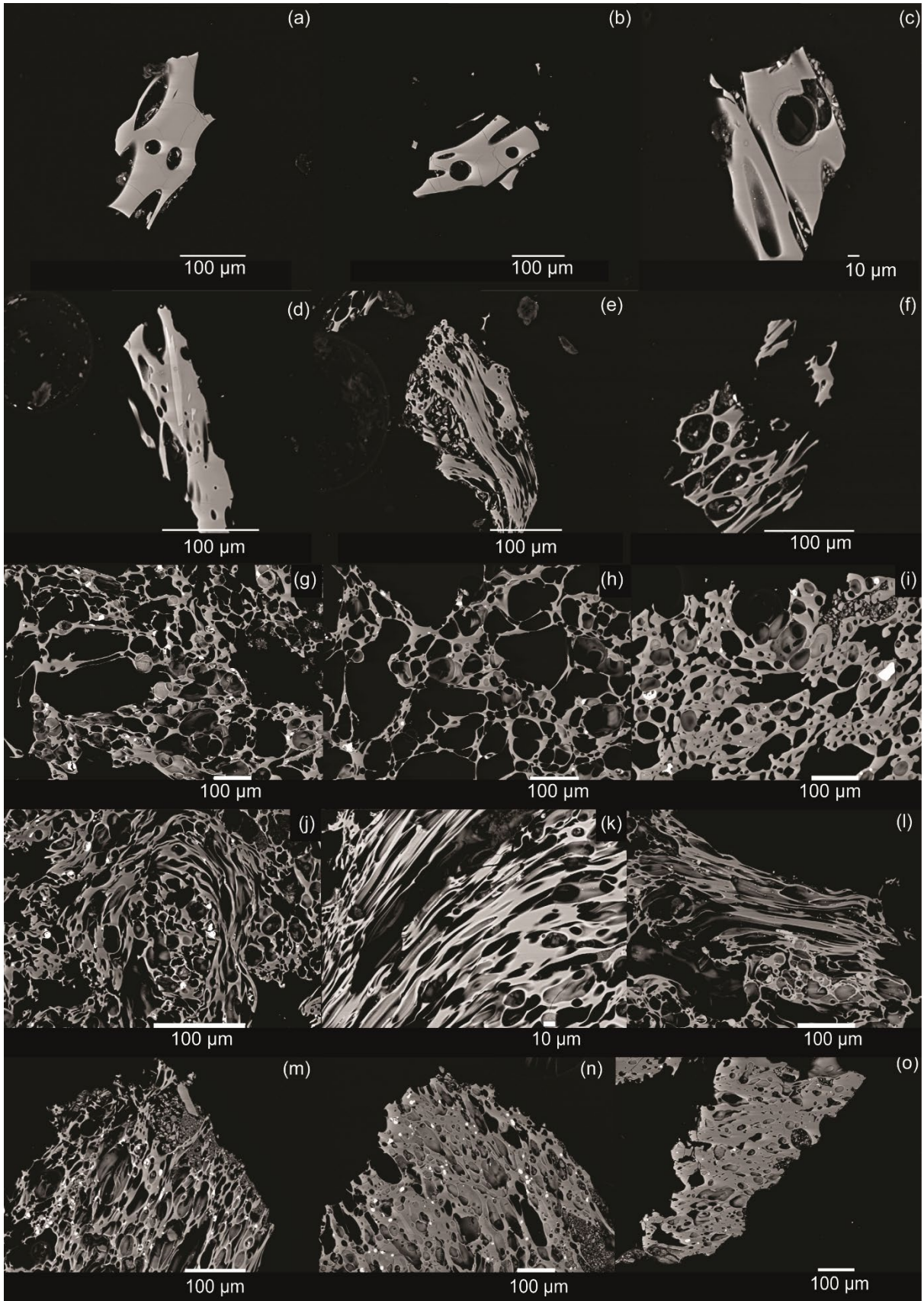
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931

932 **Figure 1.**

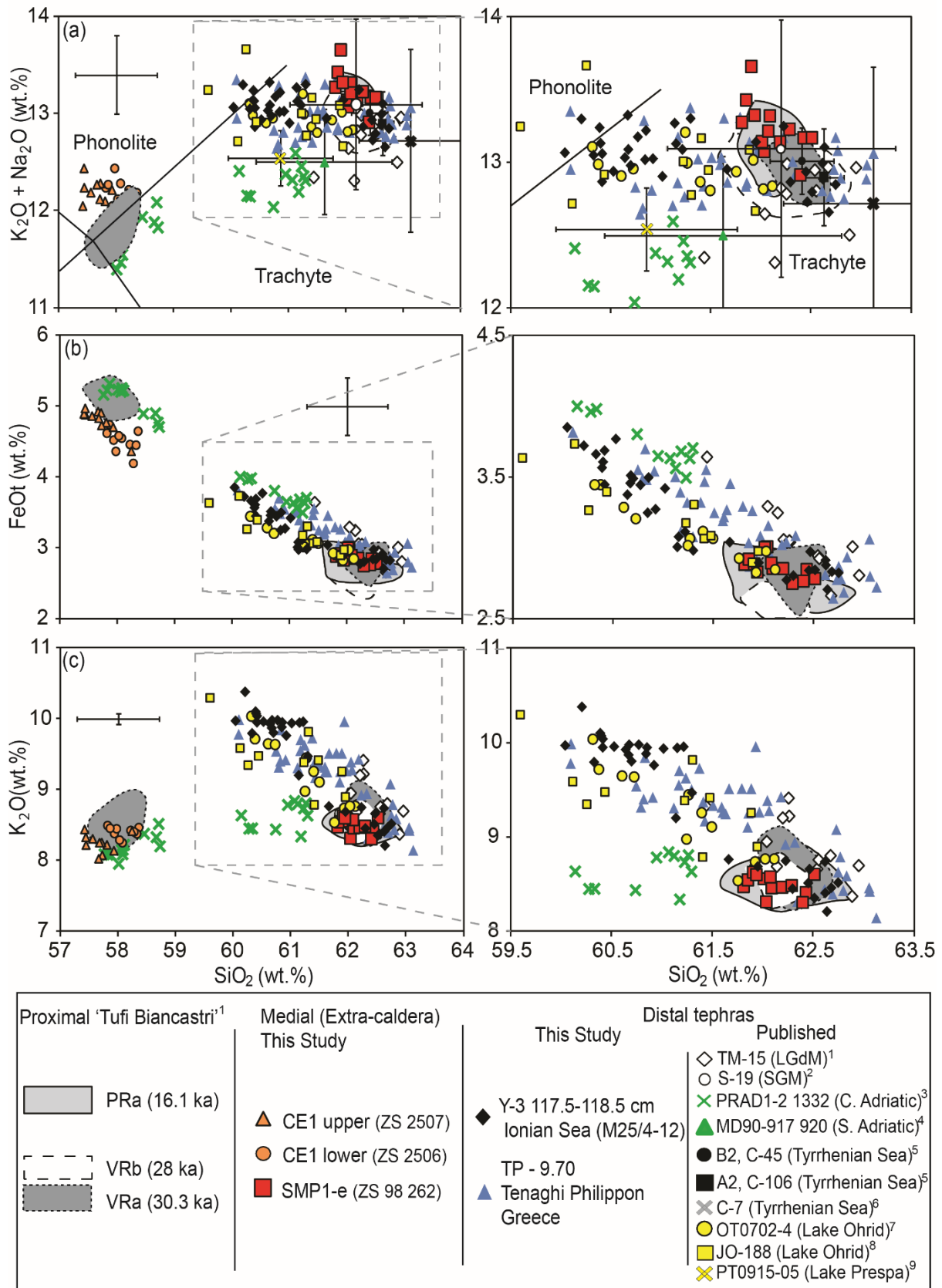
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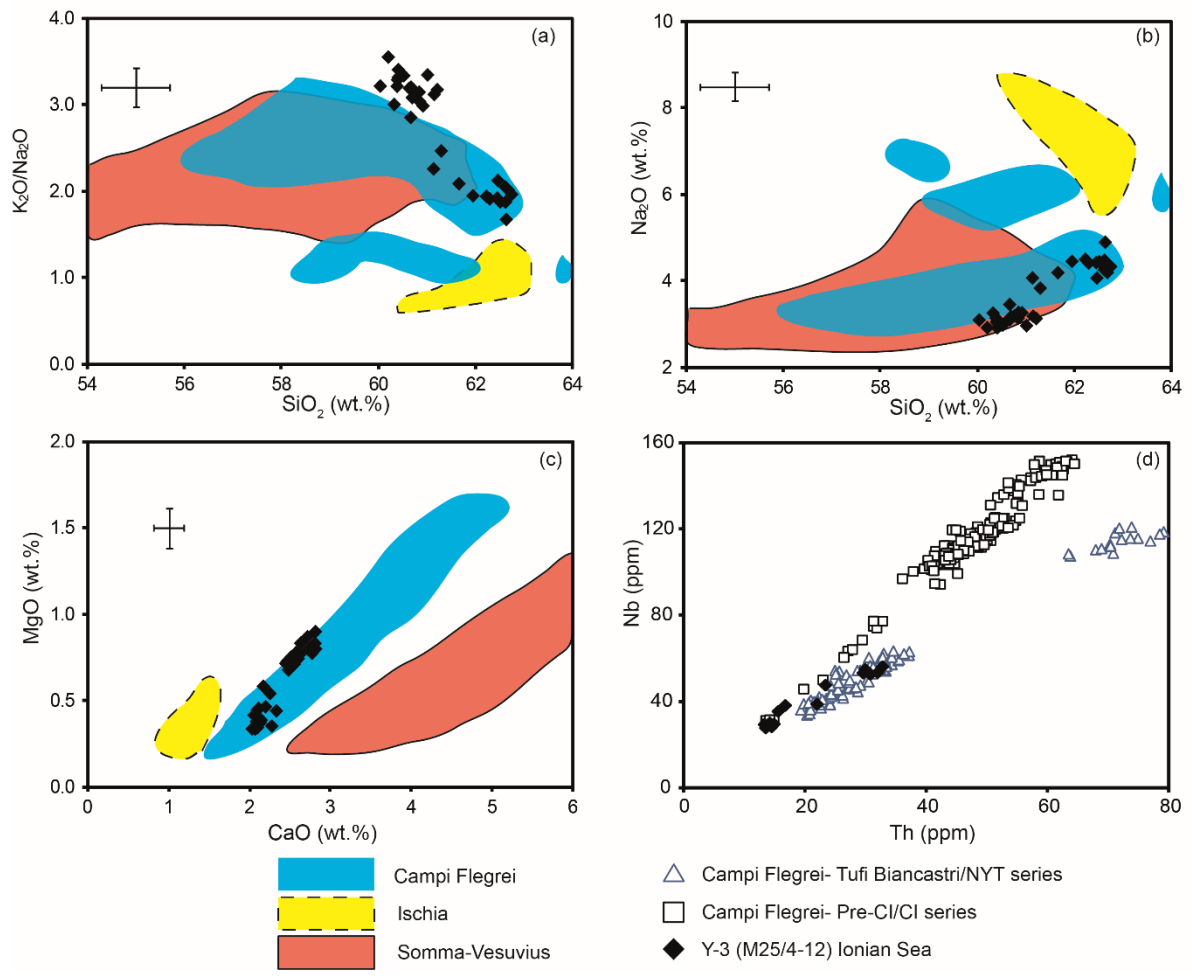
936 **Figure 2**



937

938 **Figure 3**

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942 **Figure 4**

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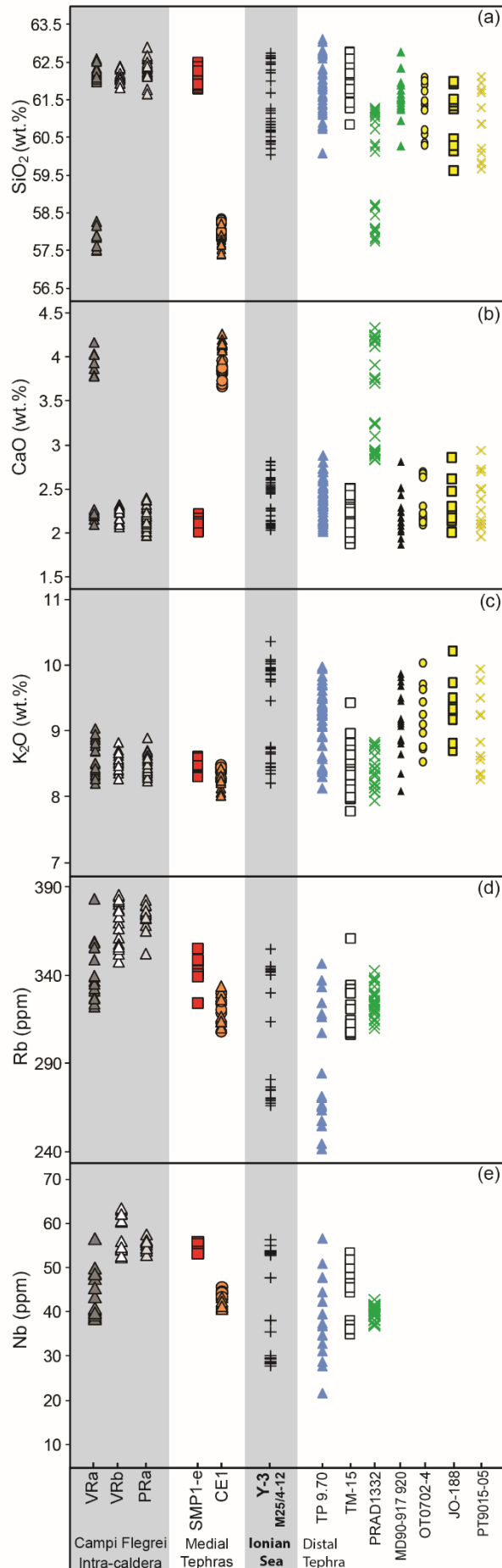
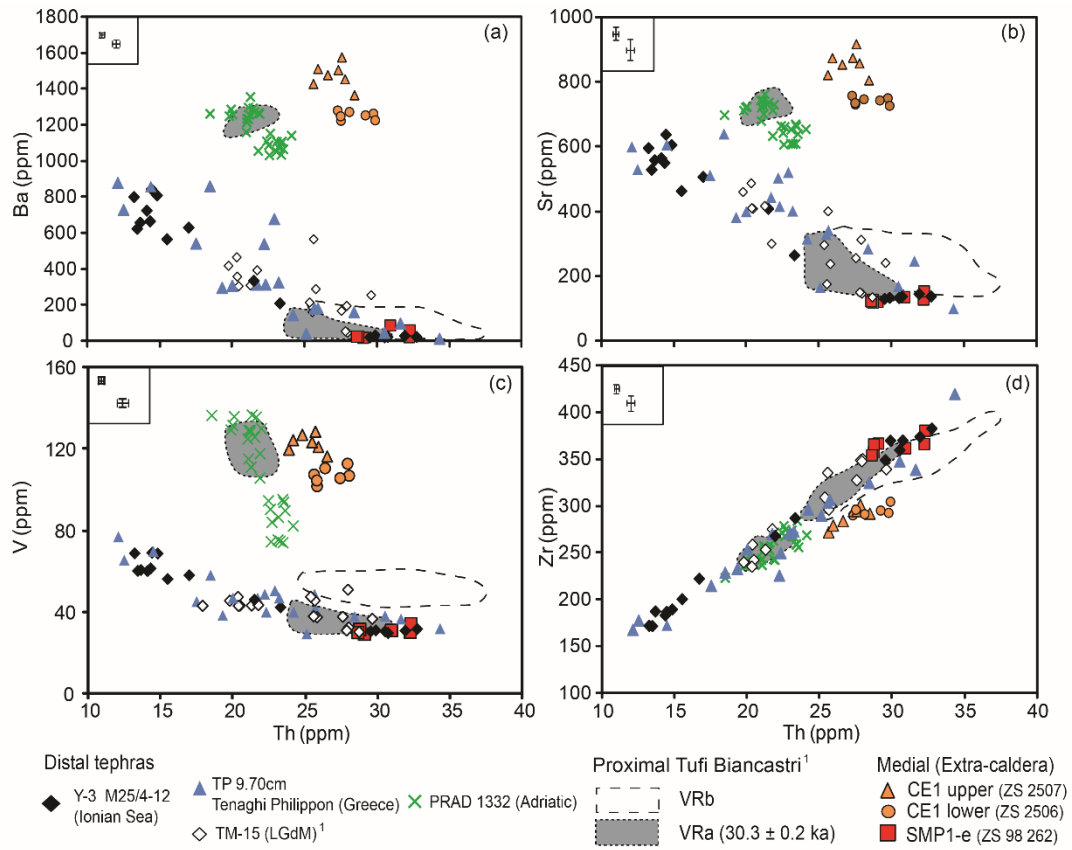
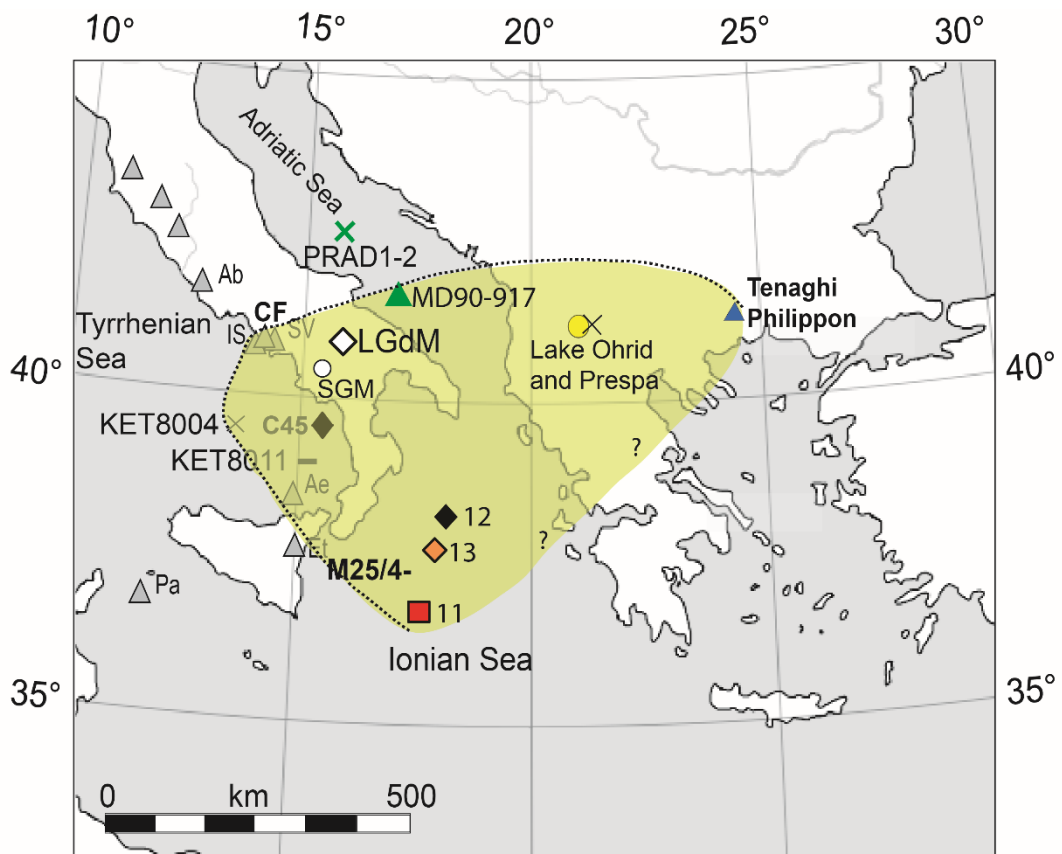


Figure 5



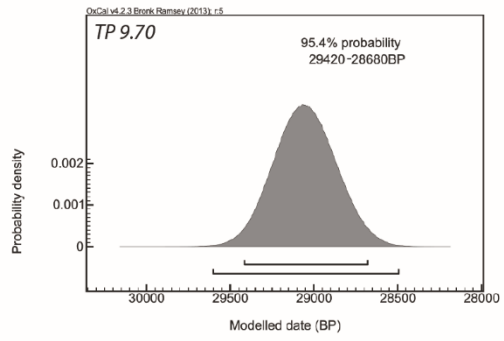
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946 **Figure 6**



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948 **Figure 7**

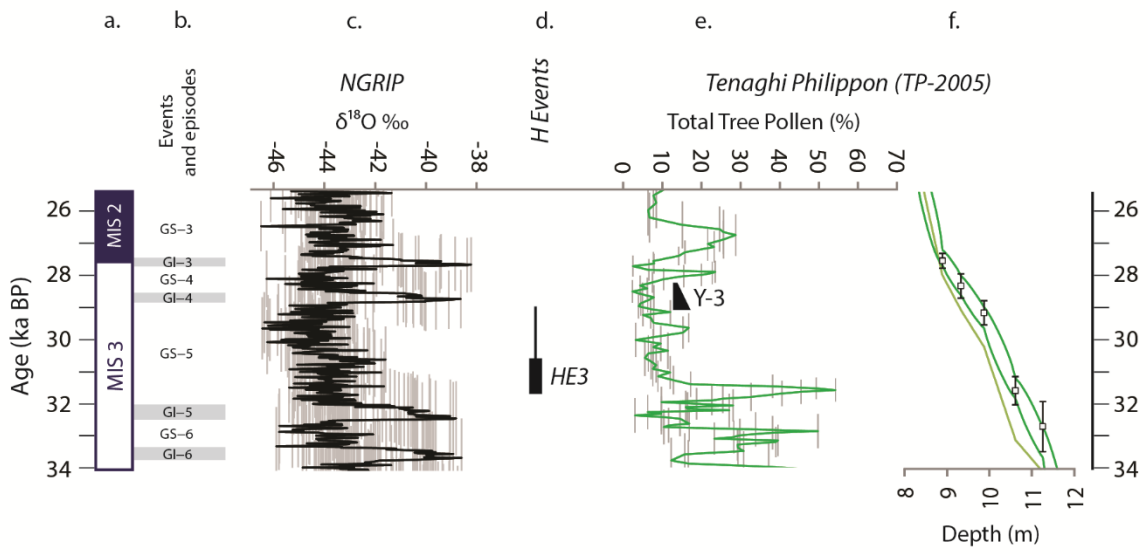


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951 Figure 8

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954 Figure 9.

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Tephra/Archive	Depth (cm)	Thickness (cm)	Description	Colour	Phenocrysts	Dispersal	Distance from CZ (km)	Dating (method)	¹⁴ C (uncal)	Age (cal) vs BP	Reference
V3 type locality/ Ionian Sea											
V3, Ionian Sea (RC 191)	245-244	1	-	yellowish-grey layer	Kf, ap, bt	SE		-	-	-	Keller et al. (1978)
V3 (M25/4-12, M25/4-13 *)	117-118	1	(1) HV tubular shards; (2) MW blocky shards	yellowish-grey layer; clear to brown shards	Kf, ap, bt	SE	450	¹⁴ C interpolation	-	*22300-28300	Kami (1997), This Study
Proximal V3 correlative											
Vna (Verdello Valley, CF)	-	450	HV punice	light beige (base) to brown (top)	plg, K, cpx, bt, mag, minor ap	-	-	⁴⁰ Ar/ ³⁹ Ar	-	30300 ± 400 (2σ) (29900-30700)	Orsi et al. (1996), Pappalardo et al. (1999*), Tomlinson et al. (2012)
Medial V3 correlative											
SMP1-e tephra											
SMP1-e, Sorrentine Peninsula (ZS98262)	-	40-50	HV punice	light grey	aphytic, minor kf	SE	30	AMS ¹⁴ C on charcoal from palaeosol directly beneath the tephra	25820 ± 270	29390-30720	Sulpice et al. (2003), Di Vito et al. (2008)
CE1, Cervino (ZS2506 lower; ZS2507 upper)		100	HV punice (lower) to MW punice (upper)	light grey	aphytic, minor kf	NE	32	-	-	-	
Discal V3 correlative											
Marine											
Thyrrhenian Sea											
A2, Salerno Gulf core C-106	565-579	14	-	-	Kf, bt, cpx	S		AMS ¹⁴ C on foraminifera 4 cm below the tephra	26030 ± 150	29350-30310	Buccheri et al. (2002), Munno and Petrosino (2004)
B2, Salerno Gulf core C-45	380-383	3	-	-	Kf, bt, cpx	S		AMS ¹⁴ C on foraminifera 3 cm below the tephra	25570 ± 110	28890-29530	Buccheri et al. (2002b), Munno and Petrosino (2004)
C7, KER004	264-274	10	-	-	-	S	100	A2/B2 integrated age (Bayesian age-depth model)	28618 ± 295 ± 41	-	Bronk/Ramsey et al. (submitted)
C7, KER011	205	1	-	-	-	S	180	-	-	-	Palermi et al. (1988), Zanchetta et al. (2008)
Adriatic Sea											
920, M090-917	920	-	-	-	-	E	260	-	-	-	Zanchetta et al. (2008)
PR0D 1332, PR0D1-2	1332	<1	-	vesicular shards	Clear	NE	190	AMS ¹⁴ C on foraminifera 8 cm below the tephra	24130 ± 150	27550-28120	Bourne et al. (2010), Piva et al. (2008)
								AMS ¹⁴ C on foraminifera 8 cm below the tephra	23390 ± 150	27020-27600	Bourne et al. (2010), Piva et al. (2008)
Terrestrial (Italy/Balkans)											
TM-15, Lago Grande di Monticchio											
TM-15, Lago Grande di Monticchio	1471, 9	28.6	HV mp	Beige	Kf, plg, bt, cpx	E	120	Varve Chronology	-	27260 ± 1360	Wulf et al. (2004, 2012*), Tomlinson et al. (2012)
S19, SGI basin	665-700	35	HV mp	-	Kf, bt	SE	90	-	-	-	Munno and Petrosino (2007)
996 Oh, Lake Ohrid	896-987	1	HV tubular shards/mp	yellowish-grey	Kf	E	550	-	-	-	Wagner et al. (2008)
10-186, Lake Ohrid	185.5-188.5	3	HV tubular shards/mp	transparent to brown	-	E	550	-	-	-	Carone et al. (2010), Ledine et al. (2010)
OM702.4, Lake Ohrid	617-620	3	(1) HV tubular; (2) Cusate shards/mp	light to brown	Kf, plg, cpx	E	550	AMS bulk ¹⁴ C on sediment directly above the tephra layer	25260 ± 210	28780-29890	Vogel et al. (2010)
PT091.5-05, Lake Prespa	616.8-617.8	1	(1) HV tubular; (2) Cusate shards/mp	transparent to brown	-	E	570	-	-	-	Damache et al. (2013)

Table 1

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Sample Core/Locality	Y-3 Ionian Sea (M25/4-12)						Extra-caldera, Tuft Binaestri						TP-9, TP-2005 Tenaghi Philippou, Greece						PRAD 1332 Central Adriatic					
	31C	29C	188	28C	34µm LA	34µm LA	34µm LA	34µm LA	34µm LA	34µm LA	34µm LA	34µm LA	34µm LA	34µm LA	5µm SIMS	5µm SIMS	34µm LA	5µm SIMS	5µm SIMS	25µm LA	25µm LA	25µm LA	25µm LA	
Major/minor	60.53	60.92	62.23	62.75	62.51	62.03	57.93	58.35	58.23	57.44	60.54	61.59	62.73	62.83	58.07	57.77	61.19	61.28	0068_107	0068_8	0068_6	0068_113		
SiO ₂	0.40	0.36	0.34	0.34	0.36	0.41	0.53	0.51	0.54	0.59	0.35	0.31	0.36	0.38	0.64	0.61	0.45	0.43	Bourne et al., (2010)*					
TiO ₂	18.35	18.43	17.82	17.81	17.79	17.90	19.16	18.60	18.78	18.66	18.71	18.29	18.16	18.36	18.26	18.52	18.30	18.13						
Al ₂ O ₃	3.77	3.27	2.77	2.82	2.79	3.01	4.53	4.65	4.37	4.88	3.63	3.19	2.78	2.94	5.20	5.16	3.68	3.63						
FeO _t	0.11	0.12	0.10	0.08	0.12	0.14	0.15	0.11	0.15	0.21	0.10	0.09	0.09	0.12	0.15	0.12	0.10	0.11						
MnO	0.83	0.74	0.35	0.45	0.43	0.40	1.16	1.06	1.26	1.36	0.81	0.55	0.49	0.42	1.51	1.53	0.75	0.75						
CaO	2.63	2.60	2.28	2.10	2.03	2.22	3.70	3.67	3.98	4.27	2.71	2.30	2.33	2.14	4.12	4.25	2.91	2.90						
Na ₂ O	2.98	3.27	4.50	4.33	4.57	4.76	3.67	3.72	3.43	3.58	2.66	3.67	3.87	4.16	3.47	3.54	3.86	3.55						
K ₂ O	9.96	9.76	8.75	8.51	8.61	8.32	8.38	8.47	8.42	8.22	9.96	9.31	9.08	8.59	8.15	8.09	8.34	8.81						
P ₂ O ₅	0.14	0.14	0.05	0.09	0.10	0.09	0.28	0.34	0.30	0.30	0.18	0.13	0.10	0.06	0.41	0.41	0.42	0.42						
Cl	0.30	0.38	0.81	0.71	0.71	0.72	0.52	0.52	0.55	0.50	0.34	0.55	-	-	-	-	-	-						
Analytical Total	95.26	97.79	98.59	94.91	95.87	94.06	93.93	95.11	99.02	98.32	98.47	97.58	97.15	94.99	97.69	98.29	95.99	96.007						
Trace																								
Spot size	34µm	34µm	34µm	20µm	34µm	34µm	34µm	34µm	34µm	34µm	34µm	34µm	34µm	34µm	5µm	5µm	34µm	5µm	5µm	25µm	25µm	25µm	25µm	
Method	LA	LA	LA	LA	LA	LA	LA	LA	LA	LA	SIMS	SIMS	LA	SIMS	SIMS	LA	LA	LA	LA	LA	LA	LA		
V	68.9	56.3	29.8	31.1	31.7	30.4	106.0	110.9	121.3	127.3	77.0	40.1	41.9	38.2	128.4	136.1	90.1	75.7						
Rb	268	276	344	355	356	324	326	309	317	314	216	272	347	334	330	310	321	337						
Sr	597	508	138	134	128	123	697	701	799	795	599	415	341	168	742	723	639	609						
Y	19.2	22.0	31.5	29.8	29.7	29.9	25.8	26.4	26.4	25.0	17.5	23.6	27.9	27.1	24.6	24.1	23.8	24.1						
Zr	172	223	370	370	358	354	277	273	281	264	168	250	307	348	257	237	256	267						
Nb	30	38	53	55	56	53	45	44	43	41	22	38	48	51	40	37	40	41						
Ba	801	636	24	32	26	26	1176	1194	1354	1376	880	315	182	47	1296	1257	1064	1051						
La	43	48	71	67	65	66	69	66	65	63	41	45	63	56	50	48	50	51						
Ce	79	96	134	131	129	128	130	129	132	126	78	89	116	122	97	94	99	97						
Pr	8.4	10.0	13.9	13.3	13.2	12.7	13.6	13.4	13.6	12.9	-	-	12.8	-	11.0	9.9	10.2	10.4						
Nd	31.1	37.2	49.3	48.8	47.7	48.6	51.9	48.4	50.2	47.8	25.9	31.5	43.9	48.8	40.3	39.5	39.4	38.5						
Sm	6.1	6.3	8.3	10.1	8.3	9.2	9.2	9.3	9.5	9.1	4.3	4.7	8.3	8.4	7.8	7.9	6.8	6.4						
Eu	1.9	1.8	1.6	1.5	1.6	1.6	2.2	2.1	2.2	2.1	1.3	1.4	1.9	1.7	2.0	1.9	1.9	1.7						
Gd	4.5	5.0	6.0	8.1	6.3	6.5	6.7	6.5	6.6	6.7	3.7	3.8	7.2	6.5	5.5	5.9	5.7	5.5						
Dy	3.6	4.9	5.6	5.7	5.0	5.3	5.3	5.2	5.1	4.9	3.5	3.8	5.3	5.3	4.9	4.6	4.7	4.3						
Er	1.8	2.3	3.1	3.1	2.9	3.1	2.7	2.7	2.6	2.4	2.0	2.2	2.7	2.9	2.4	2.5	2.4	2.3						
Yb	1.7	2.3	3.3	2.9	2.9	3.0	2.5	2.7	2.3	2.3	1.8	1.7	2.8	2.7	2.3	2.5	2.4	2.2						
Lu	0.27	0.28	0.45	0.49	0.39	0.39	0.43	0.40	0.38	0.36	-	-	0.32	-	0.34	0.38	0.34	0.32						
Ta	1.5	1.9	2.7	2.8	2.7	2.4	2.4	2.2	2.1	2.1	-	-	2.2	-	1.8	1.9	2.0	2.0						
Th	13.2	16.7	30.7	29.9	28.7	28.6	27.3	26.3	25.9	24.7	12.1	22.3	25.7	30.5	21.2	20.1	23.6	23.1						
U	4.4	5.8	9.9	10.1	9.6	8.8	9.1	8.6	7.9	7.9	3.5	6.3	8.9	8.9	7.3	7.2	8.0	8.0						

Table 2

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Sample Locality	M25/4-12, 118.5- V3 Ionian Sea (This Study)	VRB Verdoline Valley	VRa Verdoline Valley	C.1. (Fall) Voscone & Aquafidia	C.1 (Main flows) Mondragone	C.1. Upper Flow Mondragone
TAS classificatio	Low SiO ₂ Trachyte	High SiO ₂ Trachyte	Low SiO ₂ Trachyte	High SiO ₂ Trachyte	Phonolite to Trachyte	Phonolite to Trachyte
	(This Study)		Tomlinson et al. (2012)		Tomlinson et al. (2012)	
FeO/CaO	1.3 ± 0.1	1.3 ± 0.1	1.2 ± 1	1.3 ± 0.1	1.3 ± 1	1.6 ± 0.2
Cl wt. %	0.36 ± 0.06	0.72 ± 0.11	0.45 ± 0.06	0.42 ± 0.05	0.49 ± 0.11	0.63 ± 0.06
V (ppm)	63 ± 10	31 ± 1	51 ± 8	121 ± 14	38 ± 5	16 ± 4
Zr/Sr	0.34 ± 0.11	2.69 ± 0.15	0.92-2.42	0.35 ± 0.02	0.94-2.58	5-31
Nb/Th	2.1 ± 0.2	1.8 ± 0.1	1.9 ± 0.3	1.9 ± 0.1	1.9 ± 0.1	2.4 ± 0.1
Zr/Th	13.1 ± 0.6	11.9 ± 0.5	11.2 ± 1.1	12.0 ± 0.4	12.2 ± 0.3	13.4 ± 0.6
Ta/Th	0.10 ± 0.02	0.09 ± 0.01	0.09 ± 0.01	0.09 ± 0.003	0.09 ± 0.01	0.11 ± 0.004
Nb/Zr	0.16 ± 0.02	0.15 ± 0.01	0.17 ± 0.01	0.16 ± 0.004	0.15 ± 0.01	0.18 ± 0.01
Y/Th	1.41 ± 0.12	1.00 ± 0.03	1.01 ± 0.1	1.15 ± 0.04	1.06 ± 0.1	1.04 ± 0.1
La/Yb	24.5 ± 3.6	22.7 ± 2.7	25.5 ± 1.4	22 ± 1.7	23.1 ± 1.5	22.8 ± 1.7
Sample Locality	SMP1-e	SMP1-e	CE1	TM-15 LGDM	PRAD1-2 PRAD1332	Tenaghi Phillipon TP-970
TAS classificatio	Trachyte	Phonolite	Trachyte	Low SiO ₂ Trachyte	High SiO ₂ Trachyte	Low SiO ₂ Trachyte
	(ZS 98262; This Study)	(ZS 2506; This Study)	(ZS 2507; This Study)	Tomlinson et al. (2012)	(This Study)	(This Study)
FeO/CaO	1.4 ± 0.1	1.2 ± 0.1	1.2 ± 0.1	1.3 ± 0.1	1.2 ± 0.1	1.2 ± 0.1
Cl wt. %	0.71 ± 0.08	0.54 ± 0.06	0.50 ± 0.05	0.57 ± 0.21	-	-
V (ppm)	31 ± 3	107 ± 7	122 ± 10	40 ± 10	128 ± 16	87 ± 20
Zr/Sr	2.82 ± 0.41	0.40 ± 0.03	0.34 ± 0.03	0.52-2.65	0.34 ± 0.03	0.42 ± 0.04
Nb/Th	1.8 ± 0.2	1.7 ± 0.1	1.7 ± 0.2	1.8 ± 0.1	1.9 ± 0.1	1.8 ± 0.1
Zr/Th	12 ± 1.0	10.4 ± 0.7	10.9 ± 1.3	12.2 ± 0.5	11.8 ± 0.7	11.6 ± 1.0
Ta/Th	0.09 ± 0.01	0.08 ± 0.01	0.08 ± 0.02	0.09 ± 0.01	0.09 ± 0.01	0.09 ± 0.01
Nb/Zr	0.15 ± 0.01	0.16 ± 0.01	0.16 ± 0.01	0.15 ± 0.01	0.16 ± 0.01	0.15 ± 0.01
Y/Th	1.00 ± 0.09	0.99 ± 0.08	1.03 ± 0.08	1.09 ± 0.1	1.17 ± 0.07	1.04 ± 0.09
La/Yb	22.5 ± 2.4	27.1 ± 3.4	25.5 ± 2.7	23.2 ± 2.9	21.0 ± 2.9	21.3 ± 2.9

Table 3

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